Observing with the Green Bank Telescope

Some import terms...

- Astrid: <u>Astronomers Integrated Desktop</u>; Interface for controlling the GBT and monitoring observations
- Scheduling Block: A list of observing commands submitted through Astrid
- Configuration/Config: A special set of keywords and values that are used to setup the GBT for various types of observations
- **CLEO**: <u>Control Library for Operators and Engineers</u>; Useful tools + expert control of all GBT systems
- Talk & Draw: A CLEO chat tool

Tabs for different Astrid windows



Tabs for different Astrid windows



Session Info (Editable)

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Observer: Project ID: Status: LST:	unknown AGBT14B_342_29 Fault 06:23:33	Last Update: UTC Date: UTC Time: MJD:	2015-05-20 15:50:19 2015/05/20 19:50:19 57162.8266152	GBT State: Ready GBT Status:
Az commanded (deg): Az actual (deg): Az error (arcsec): Coordinate Mode: Major Coord: Major Cmd Coord: LPCs Az/XEI/EI ("): DC Az/XEI/EI ("): AS FEM Model: AS Zernike Model: AS Zernike Thrm MdI:	162.652 162.6524 -0.631 J2000 07:31:06 (RA) 07:31:06 (RA) (0.000 0.000 0.000) (0.000 -6.727 11.441) On On On	El commanded (deg): El actual (deg): El error (arcsec): Antenna State: Minor Coord: Minor Cmd Coord: LFCs (XYZ mm): LFCs (XYZ deg): DC Focus Y (mm): AS offsets:	25.808 25.8079 -0.758 Tracking -23:41:47 (Dec) -23:41:47 (Dec) (0.00 0.00 0.00) (0.00 0.00 0.00) -2.322169312 On	Fault Queue Control: Halt Queue Observation Control: Pause Stop
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- Key parts of scheduling block are
 - Catalog(<catalog>): Loads a list of sources
 - Slew(source): Moves the telescope to a source
 - config: Defines backend and IF/LO parameters
 - Configure(config): Executes config instructions
 - Balance: Adjusts power levels to optimal range
 - Observing Directive
- There are *lots* of Astrid commands, and you can use python to make fairly complicated scripts
- Let's look at the config portion more carefully...

receiver = 'Rcvr1_2'	vegas.obsmode = 'search'
restfreq = 1500.0	vegas.numchan = 2048
obstype ='Pulsar'	vegas.polnmode = 'full_stokes'
backend = 'VEGAS'	vegas.scale = 10000
pol ='Linear'	vegas.outbits = 8
swmode = 'tp_nocal'	
swtype = 'none'	vegas.fold_dumptime = 10
noisecal = 'off'	vegas.fold_bins = 256
swper = 0.04	vegas.fold_parfile = "/users/rlynch/myparfile.par"
swfreq = 0.0	vegas.dm = 50
nwin = 1	
deltafreq = 0	
ifbw = 0	
bandwidth = 800	
tint = 81.92e-6	
vlow = 0	
vhigh = 0	
vframe = 'topo'	
vdef = 'Radio'	

- swmode: Switching mode ('tp' or 'tp_nocal')
- swtype: Only used for frequency switching
- noisecal: Controls noise diode ('lo' or 'hi' or 'off')
- swper: Switching period (always 0.04 [25 Hz] for pulsar observing)
- swfreq : Only used for frequency switching
- nwin: Number of spectral windows (usually 1 for pulsar observing)
- deltafreq: Offset of spectral windows (always 0 for pulsar observing)
- ifbw: BW of filter in IF rack (80 for 100 MHz total bandwidth, otherwise 0)

- vegas.obsmode
 - fold, cal, search, coherent_{fold,cal,search}
- vegas.numchan
 - 2ⁿ, n = 6...11 (64...4096 [8192])
 - Max depends on bandwidth
- vegas.polnmode
 - full_stokes or total_intensity
- vegas.scale
 - Varies with setup, but fairly stable
- vegas.outbits
 - Always 8

- vegas.fold_dumptime
 - Typically 10
- vegas.fold_bins
 - Typically 256, higher for coherent modes
- vegas.fold_parfile
 - Always make sure this is correct!!
- vegas.dm = 50
 - Specific to coherent_search

A word on tint

tint = acc_len * vegas.numchan/bandwidth

- acc_len is a hardware parameter not directly set by the user
 - acc_len > 4 for 100 and 200 MHz modes
 - acc_len > 16 for 800 MHz mode
 - acc_len < 100 recommended</p>

- Write a simple observing script to observe B1937+21 in search mode for 30 minutes
 - 2048 channels, 40.96 us time resolution, total intensity, 800 MHz bandwidth
- Ask yourself
 - How will Astrid know the location of the source?
 - How do things need to be configured?
 - Is the system balanced?
 - How do I tell Astrid to start taking data?

- Write a simple observing script to observe J1713+0747 using coherent fold mode for 20 minutes
 - 512 channels, 800 MHz bandwidth, (what is tint?)
 - Include all the necessary steps to fully calibrate the data

- How do we calibrate pulsar data?
 - Raw data are recorded in units of "counts", which change depending on how the system is balanced
 - We want physical units (Jansky)
- 1) Observe a standard source of known flux density
- 2) Use this to measure the strength of the GBT noise diode
- 3) Balance the system at the position of the pulsar and observe the noise diode
- 4) Observe the pulsar **without rebalancing**, and use the noise diode to calibrate the pulsar data

• Scheduling blocks use Python!

```
sources = ["B1937+21", "J1713+0747"]
for src in sources:
   Slew(src)
   Balance()
   Track(src, None, 20*60)
```

```
for src in sources:
  config = """
   . . .
  vegas.fold_parfile = "/home/gpu/tzpar/{src}.par"
   . . .
  """.format(src=src)
srcdict = {
  "B1937+21":{"DM":50,"parfile":"mypar"},
   "J1713+0747":{"DM":100,"parfile":"mypar2"}
for src in srcdict:
  config =
           || || ||
  vegas.fold_parfile = {parfile}
  """.format(parfile=src["parfile"])
```

- Astrid has a number of useful convenience functions and options
 - Track() can take startTime and stopTime instead of scan length
 - Horizon() objects can be used to define times based on a source's elevation (.e.g., start observing when a source rises or stop when it sets)
 - OnOff() is useful for flux calibration scans
 - Observe an On source and Off source position automatically
 - AutoPeakFocus() necessary at high frequencies (roughly Cband and above)

- Track(src, None, 3600, startTime="15:15:00")
- Track(src, None, startTime="15:15:00", stopTime="16:15:00")

- myhorizon = Horizon(6.0)
 - Track(src, None, stopTime=myhorizon)
 - Track(src, None, startTime="15:15:00", stopTime=myhorizon)
- OnOff("3C48", Offset(0,1,cosv=False), 90)

- Write an observing script that will loop over B1937+21, J1713+0737, and J1124+78
 - Perform a cal scan and online folding scan on each using coherent modes
 - You may specify the parfile in any way you choose
 - Do a 90 second cal scan and a 20 minute pulsar scan

- VEGAS is controlled through the GBT but there are standalone tools used to monitor observations
- Some are always used, some are specific to coherent and incoherent modes

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• Launched by typing vpmStatus

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vegas-hpc7:	1250.78	exiting	exiting	7		
vegas-hpc8:	1150.78	exiting	exiting	7		
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vegas-hp	C6:	This bank is	not curre	ntly	/ in use				
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Launched by typing vpmHPCStatus

vpmHPCStatus



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www.gb.nrao.edu/vpm

- Final task: Write an observing script to observe some test pulsars (we will specify them later) using coherent fold mode
 - Don't worry about the Catalog and specific sources for now
 - Include a cal scan
 - You can be as pythonic as you want, or you can hard code things by hand
- When you are done, ask me to check your work

- Astrid has a number of useful convenience functions and options
 - Can supply startTime and stopTime instead of scan length
 - Horizon() objects can be used to define times based on a source's elevation (.e.g., start observing when a source rises or stop when it sets)
 - AutoPeakFocus() necessary at high frequencies (roughly C-band and above)
 - Never hurts at lower frequencies, but it has little benefit and takes time

- Track(src, None, 3600, startTime="15:15:00")
- Track(src, None, startTime="15:15:00", stopTime="16:15:00")

- myhorizon = Horizon(6.0)
 - Track(src, None, stopTime=myhorizon)
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