

# Introduction to CCB Observing

BSM

23feb09

1. Quick overview of the instrument
2. Observing: Configuring
3. Observing: Pointing and Focus
4. Looking at the data
5. Sensitivity
6. Current work

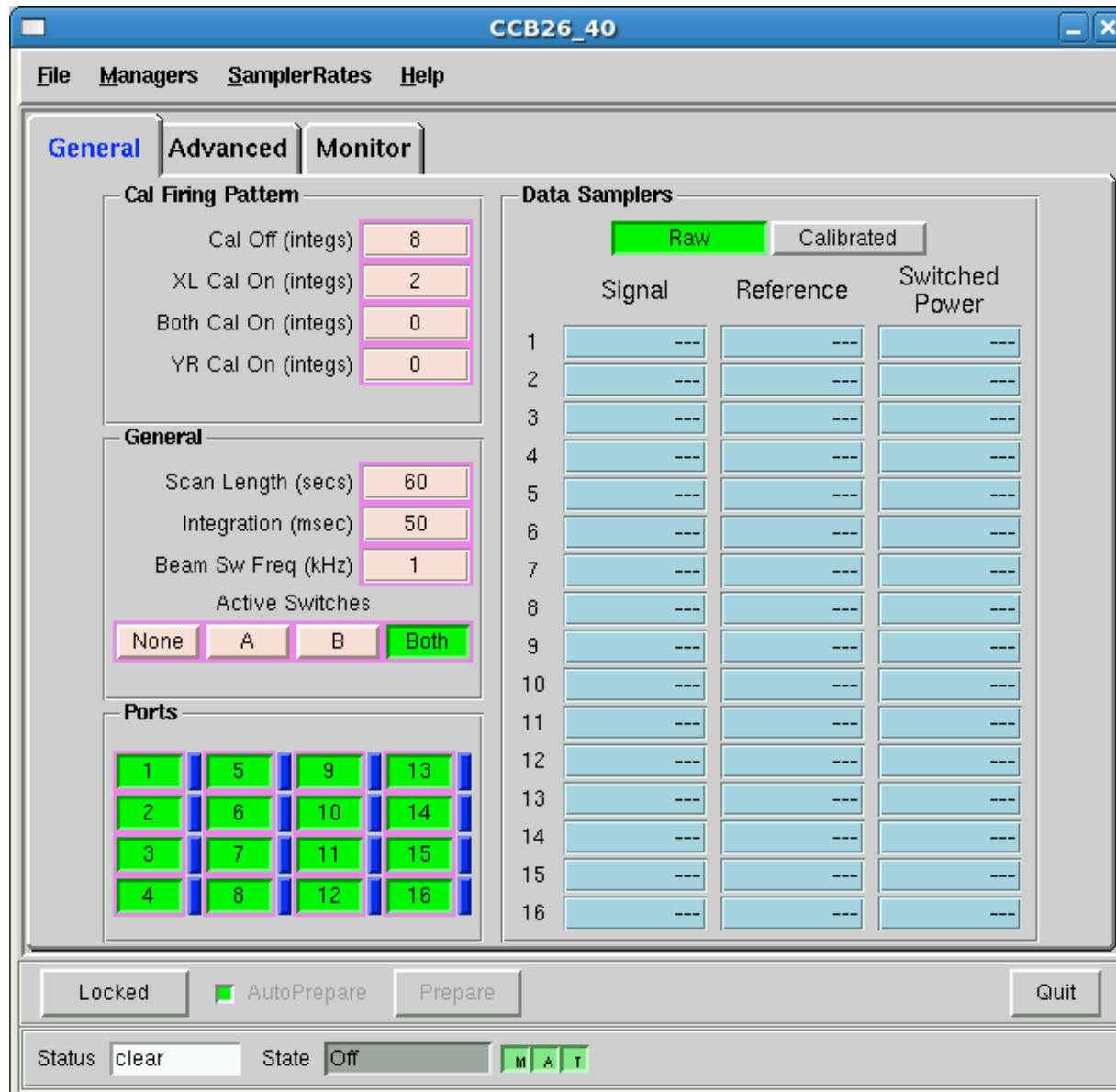
# Caltech Continuum Backend

Goal: provide sensitive broad-band continuum capability

- Direct RF detection -- reduce system complexity, improve stability
  - no IF's , mixers, etc.
  - Ka receiver: 16 (8) channels of 3.5 GHz bandwidth each
    - 2 beams x 2 (1) pol'n x 4 frequency channels
- Rapid beamswitching (4 kHz is standard)
  - cal and beam switch (aka phase switch) control is by CCB/Ka direct connection -- doesn't use switching signal master
  - arrangement of cal switching slightly different: done on a whole integration basis, not within an integration
  - (10% blanking warning is standard)
- Cals are controlled independently rather than together to allow full calibration (needed because of leakage terms inherent in the rx architecture)

# Configuration

- receiver = 'Rcvr26\_40'
- beam = 'B12'
- obstype = 'Continuum'
- backend = 'CCB'
- nwin = 4
- restfreq = 27000.0,32000.0,35000.0,38000.0
- deltafreq = 0.0,0.0,0.0,0.0
- bandwidth = 600.0,600.0,600.0,600.0
- swmode = "sp"
- swtype = "bsw"
- ccb.cal\_off\_integs = 20
- ccb.XL\_on\_integs=2
- ccb.YR\_on\_integs=2
- ccb.both\_on\_integs=2
- tint = 0.005
- Ccb.bswfreq = 4
- pol = "Circular"
- vdef='Radio'
- vframe='topo'



# Pointing & Focus with GFM

- With about the reliability of other receivers, “Just works”
- Similar problems occur (cabling/beam IDs can be confused after new installation, requiring the usual workarounds)

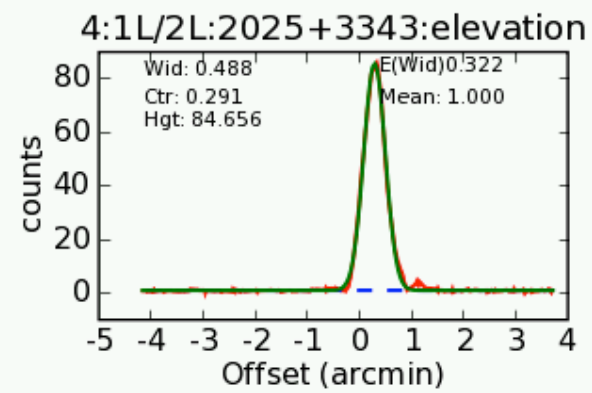
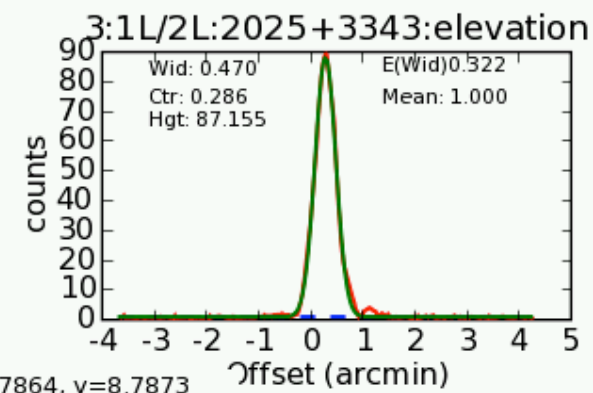
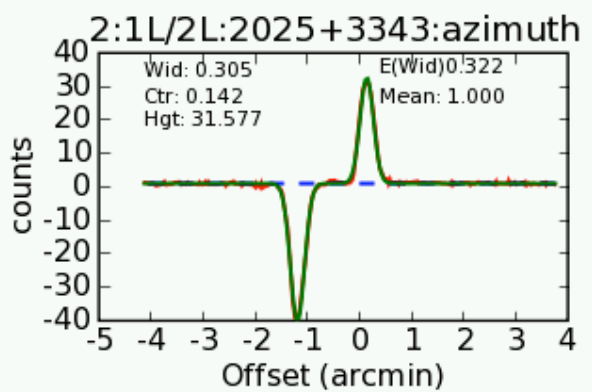
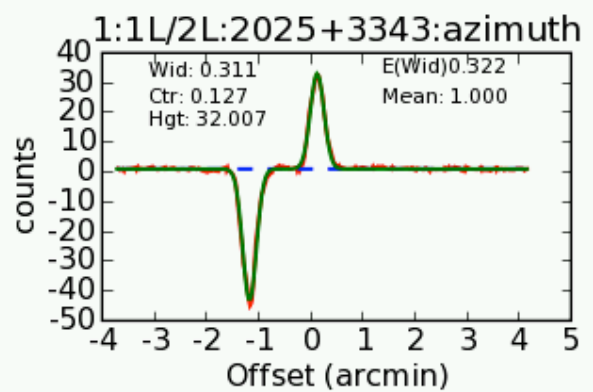
But:

- Data are “raw” (in counts), beamswitched
  - we should fix this
- Because Ka only has one polarization GFM often issues lots of harmless complaints.

ools Help



Pointing Focus OOF Continuum Spectral Line Beta



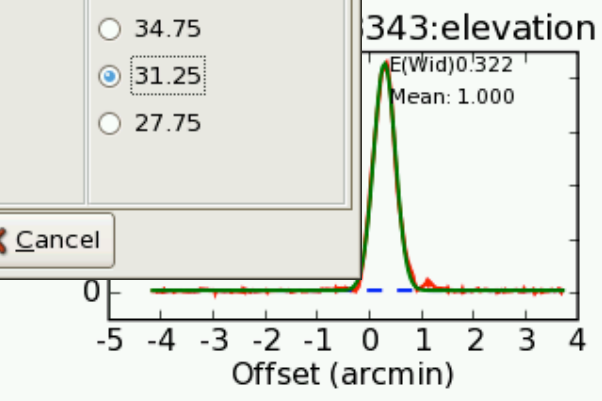
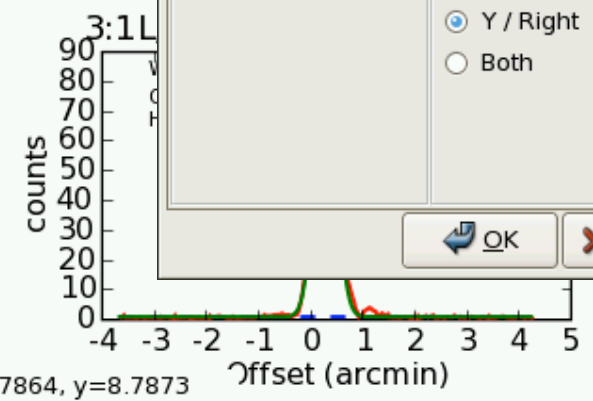
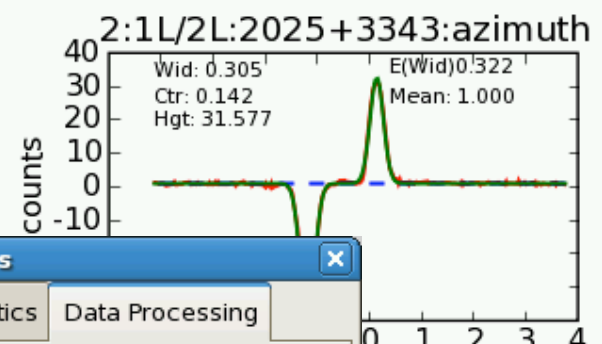
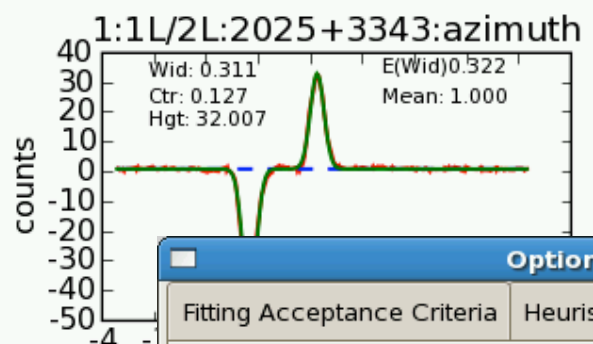
x=-4.7864, y=8.7873

Rcvr26\_40 Feeds = [1, 2] Raw Polarizations = ['XL'] Center Sky Frequency = 38.25

```
del2: 0.291 tel2: 2.304 *** heuristics failed ***
-----
ldAz2: 0.000 oldEl: 0.000
dAz2: 0.134 dEl: 0.288
ewAz2: 0.134 NewEl: 0.288 *** El heuristics failed
```



Pointing Focus OOF Continuum Spectral Line Beta



**Options**

Fitting Acceptance Criteria Heuristics Data Processing

Processing:  Raw

Polarization:  X / Left  Y / Right  Both

Frequencies (GHz):  38.25  34.75  31.25  27.75

OK Cancel

Rcvr26\_40 Feeds = [1, 2] Raw Polarizations = ['XL'] Center Sky Frequency = 38.25

```
del2: 0.291 tel2: 2.304 *** heuristics failed ***
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ldAz2: 0.000 oldEl: 0.000
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# Template SBs

In ~bmason/ccbPub (\*.turtle)

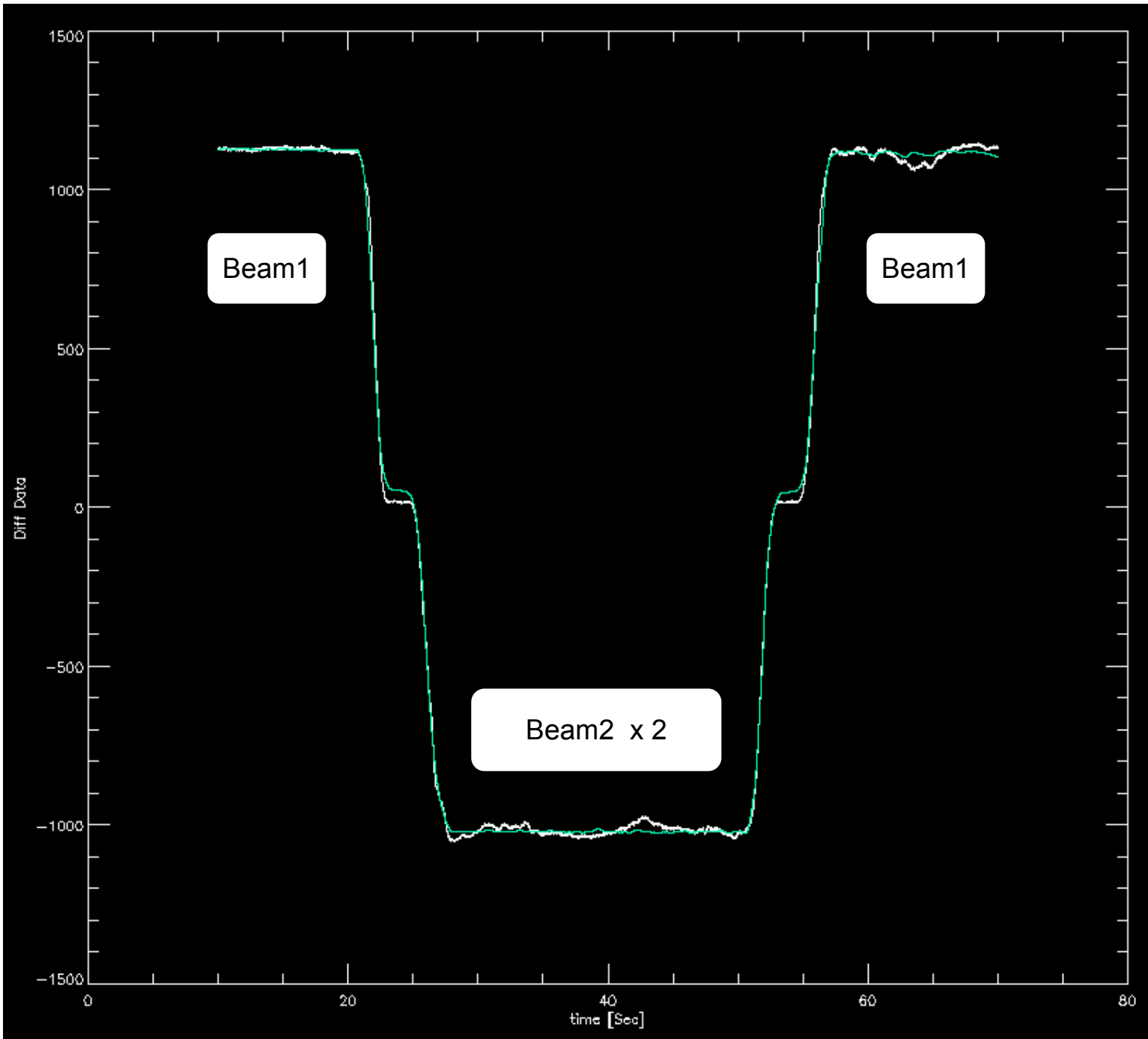
- Pointing/Focus
- Point source photometry **75%+ of observations**
- Mapping
  - data reduction is experimental (my scripts;OBIT)
  - key is correcting for the beamswitching.
- skydips
  
- ALWAYS recommend to observe primary flux calibrator  
If feasible (3c48, 3c147, 3c286) -- calibration transferrable.  
peak/focus  
Nod (included in peak/focus template SB)
- However, Tcals and RF gains are stable so calibration is transferrable over days to weeks.

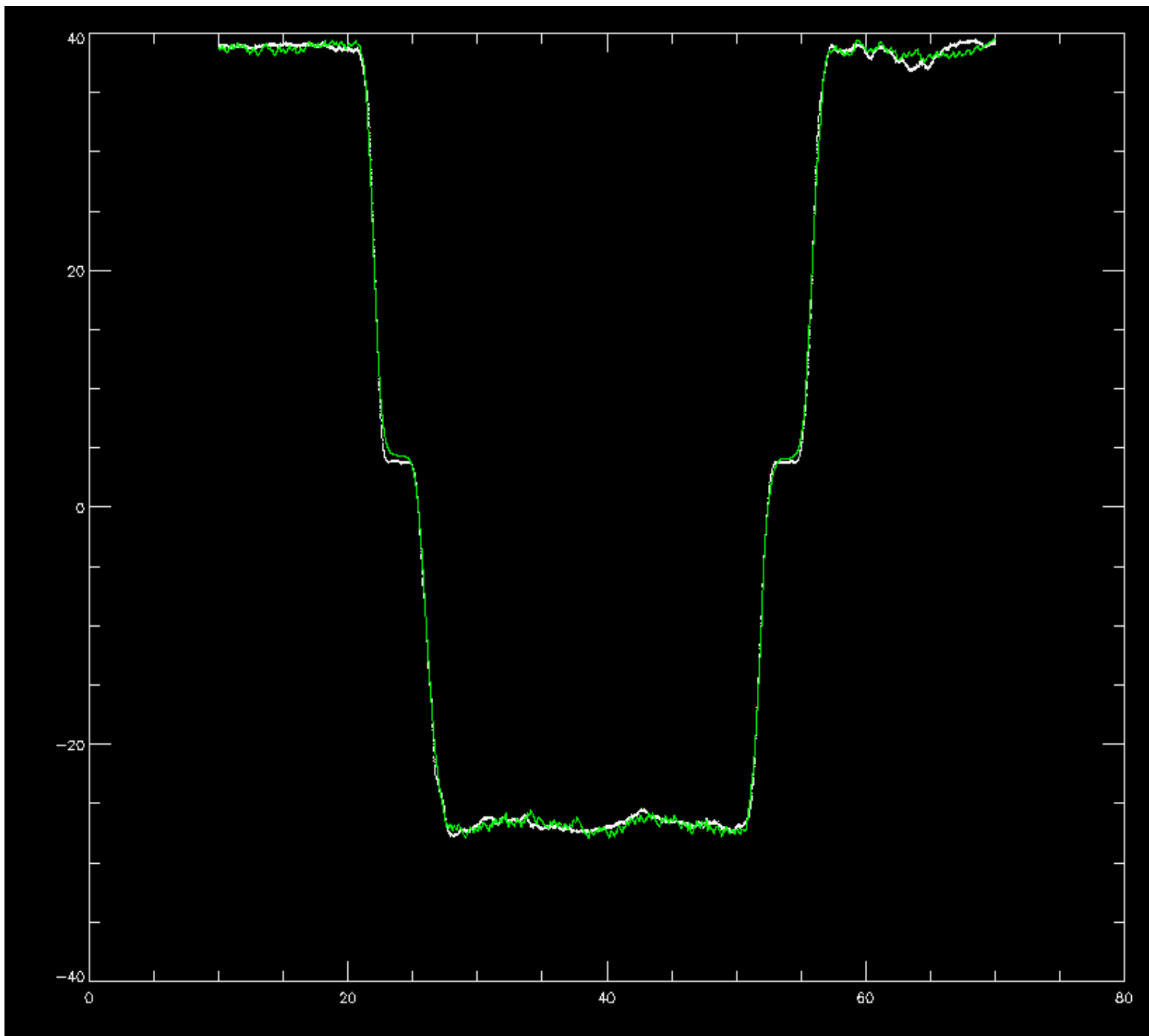


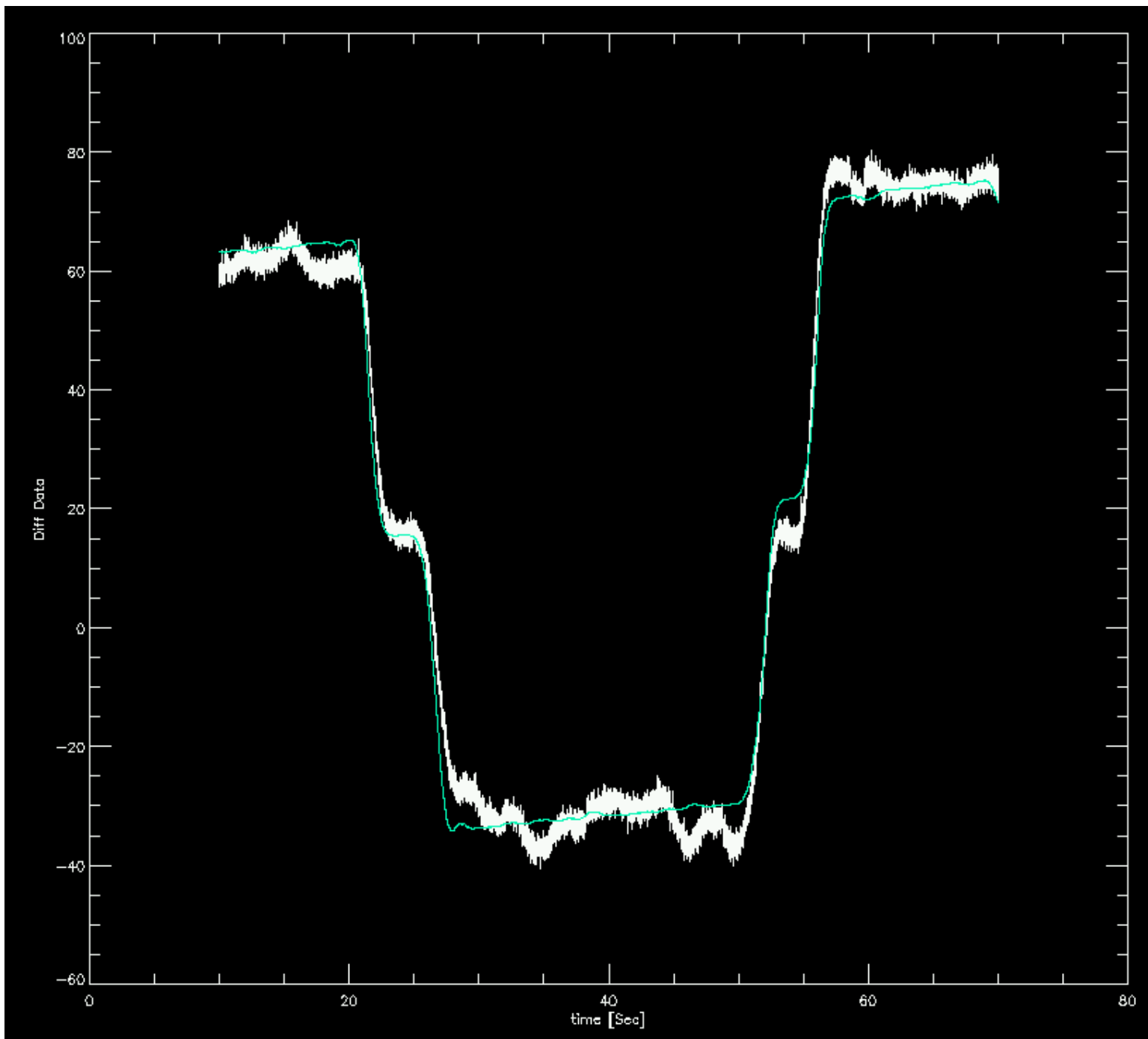
# Quick Look Data Reduction

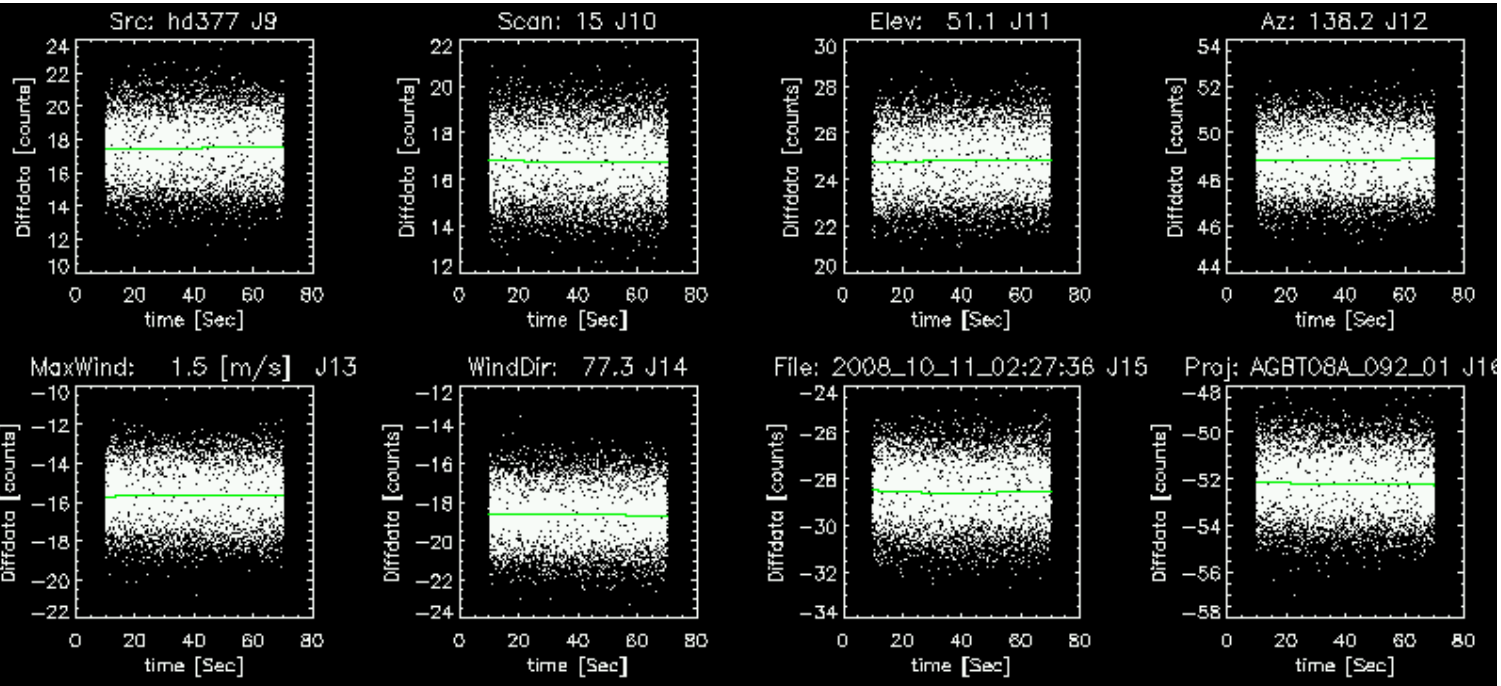
- Uses my IDL scripts
- `~bmason/ccbPub/ccbidl`
  - Shell script to launch regular IDL
- Basic commands in `~bmason/ccbPub/README.txt`
  
- Basic observation: on-the-fly nod
  - 70 seconds: 10sec dwell, On/Off/Off/On (symmetric nod), 10 sec slews.
  - Compute template source response (given source Ra/Dec and telescope position as a function of time) and fit
  - Can be generalized to fit for pointing offset in each observation if  $\text{SNR} > 1$

# Beamswitched power: Beam 1 - Beam 2

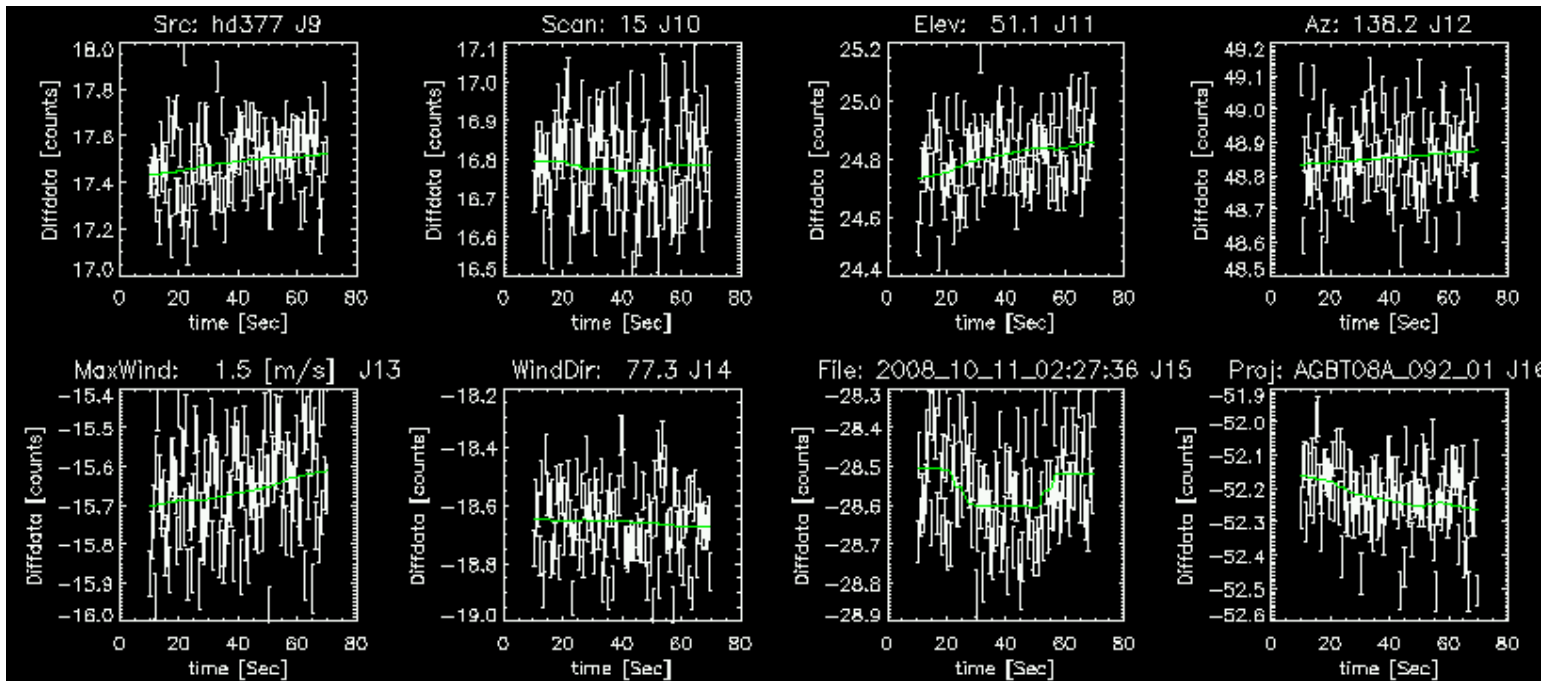




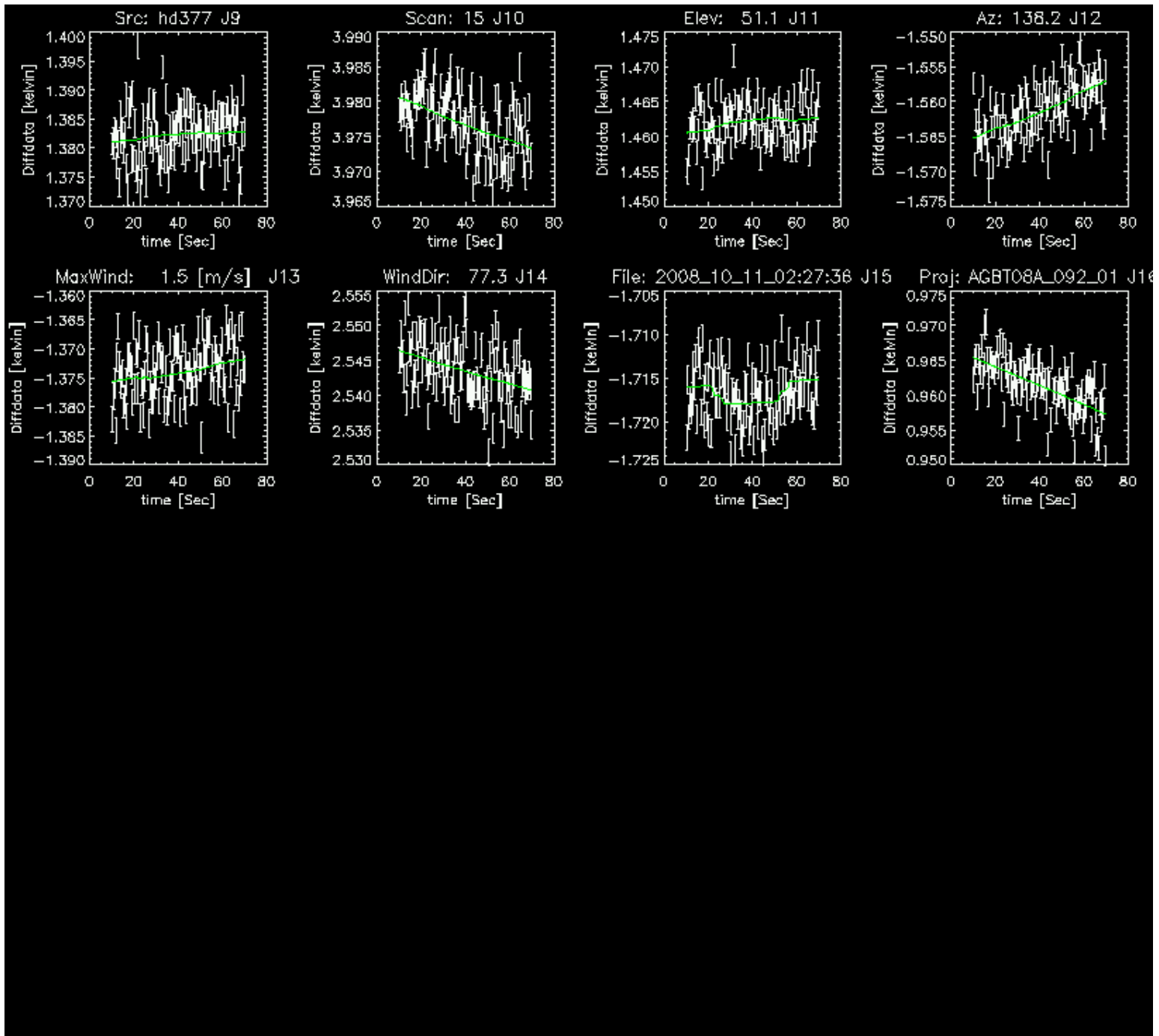


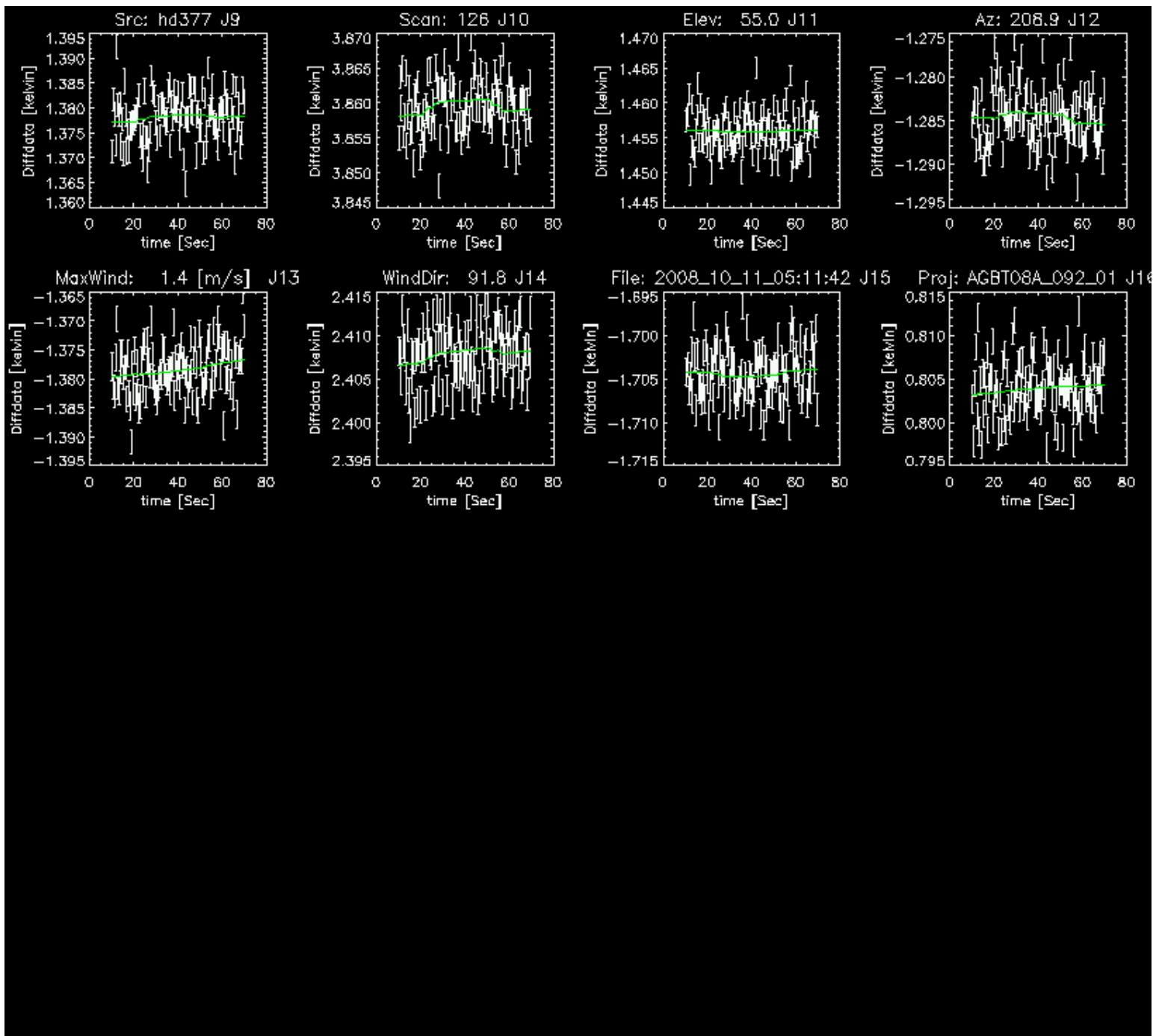


Indexscans,si            -- analagous to filling  
 Readccbotfnod,si[12],q   -- read scan 12 into variable q  
 Fitccbotfnod,q            -- fit the nod



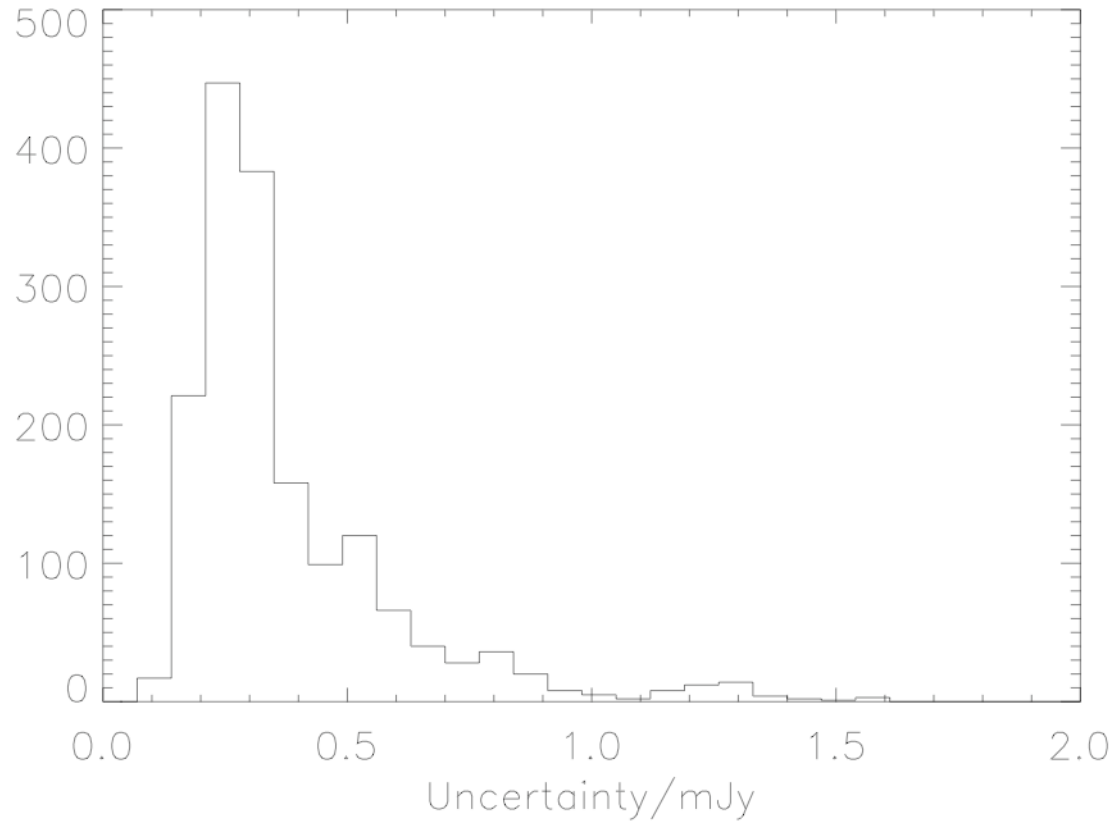
Indexscans,si  
 Readccbotfnod,si[12],q  
 Fitccbotfnod,q,binw=0.5 bin into 0.5 sec ints before fit





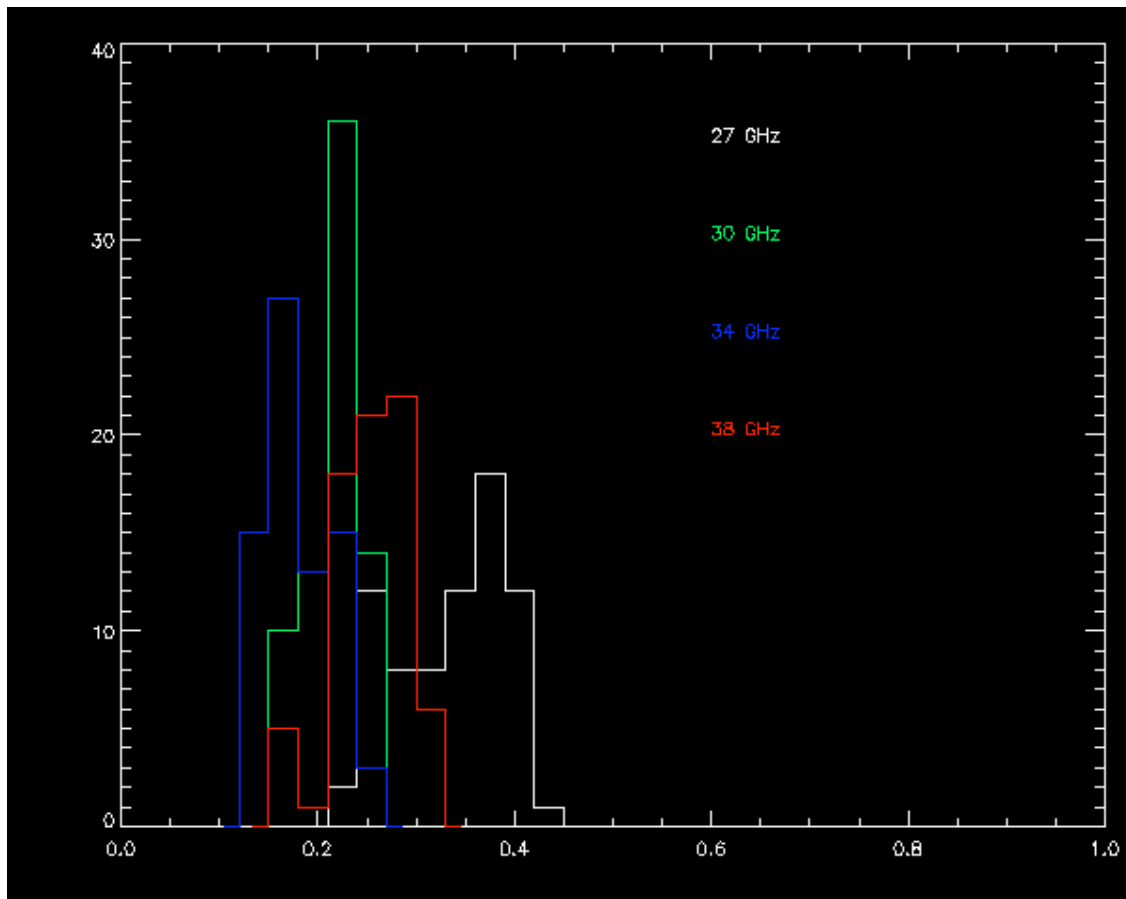


# Sensitivity



RMS uncertainty in mean (31 GHz channel -- old best): 70 second Nod measurement

# New Sensitivity



150  $\mu$ Jy RMS (34 GHz)

100  $\mu$ Jy RMS (all combined)

# Current Work

- Improve integration with GBT system
  - Tcals into Measurements database
  - GFM to calibrate data (instead of raw beamswitched)
- Streamline postprocessing data reduction
  - Quick look stable and easy to use
- Develop more precise photometry techniques
  - Impose small nutation in az and el on Nod and fit for pointing offsets, beam offsets
  - Preliminary: 3% RMS photometry at 38 GHz.