



Combining Single Dish and Interferometer Data

Natalie Butterfield



Talk Outline

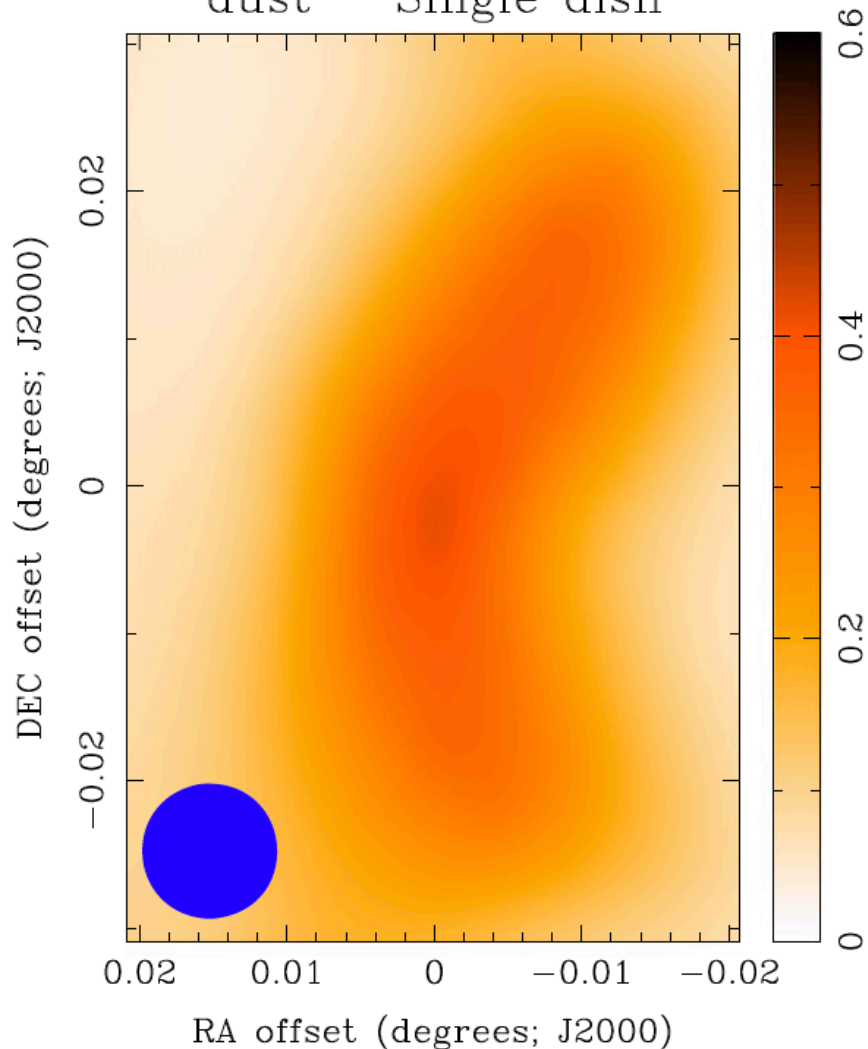


- Imaging SD vs Interferometer Data
 - discuss: resolution, size scales, etc
- Imaging artifacts in interferometer data
 - Negative bowls, Data cube example
- Science behind negative bowls
- Measuring visibilities
 - Example of what SD+VLA data means in UV plane
- Real-life example in science
 - Rathborne 2015 example; measurements
- How to do feathering
 - Casa task: feathering example
- Baselines for VLA and ALMA: which arrays to use with GBT data
- Summary slide

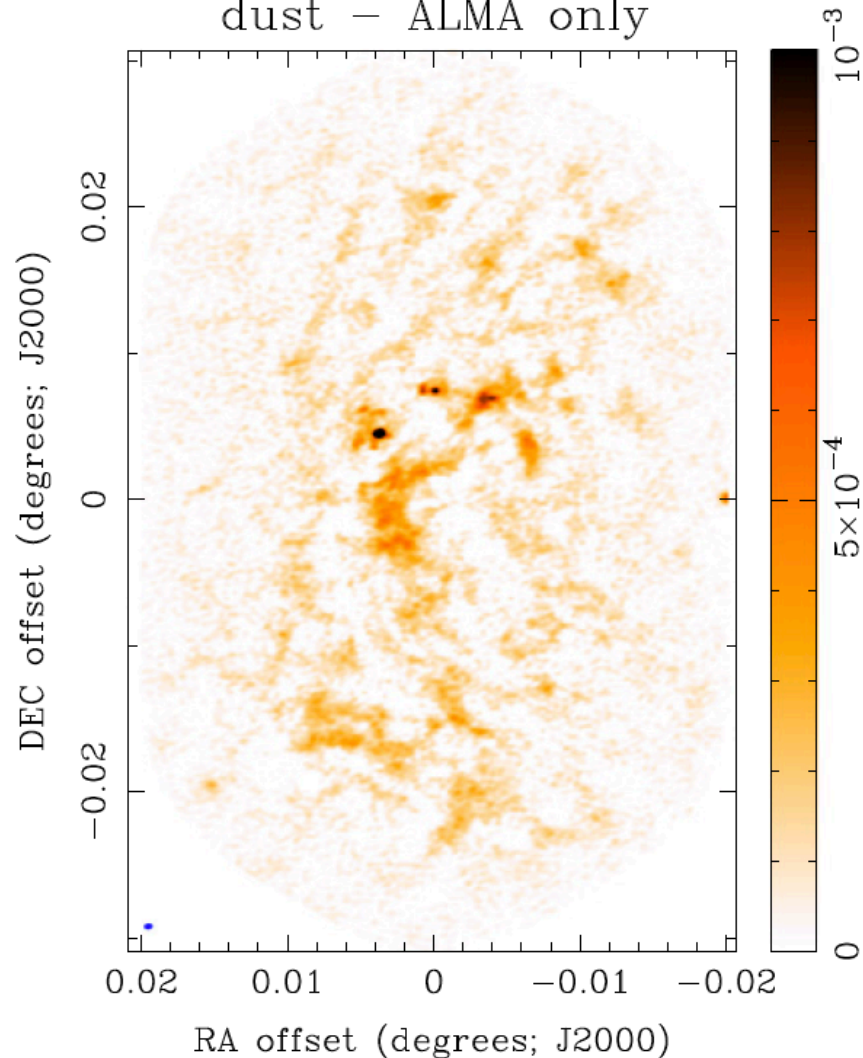


Imaging SD vs Interferometer Data

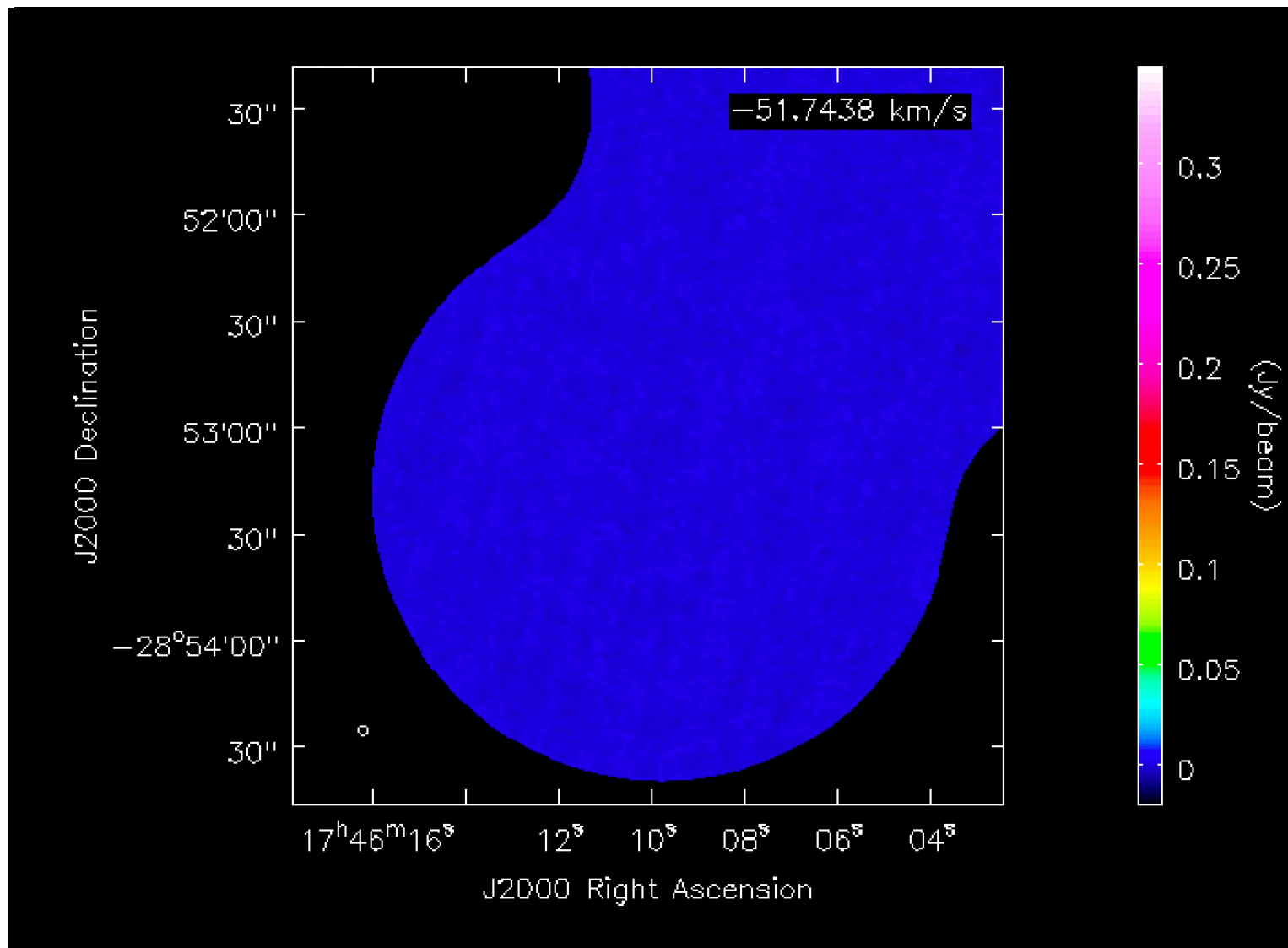
dust - Single dish



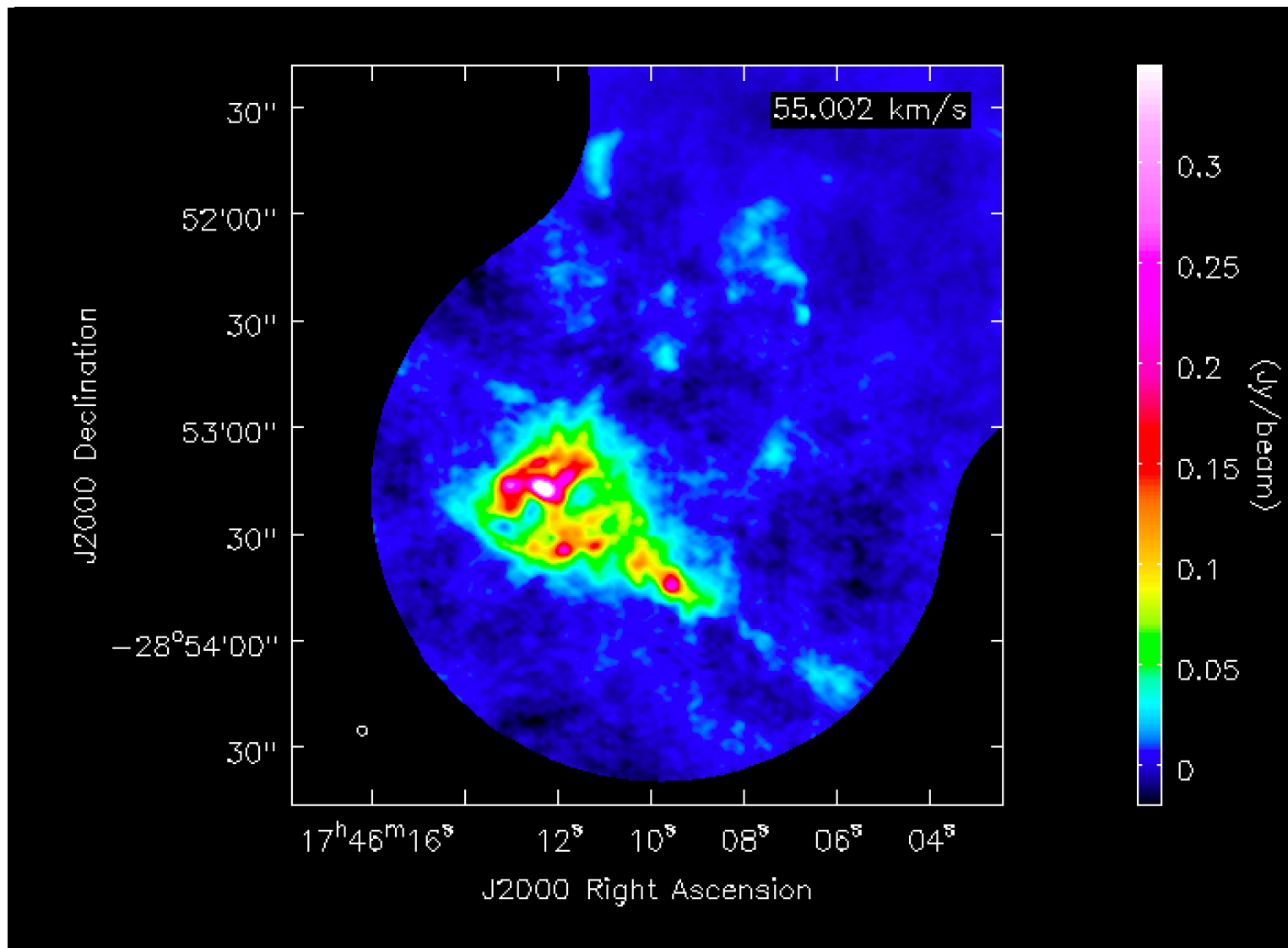
dust - ALMA only



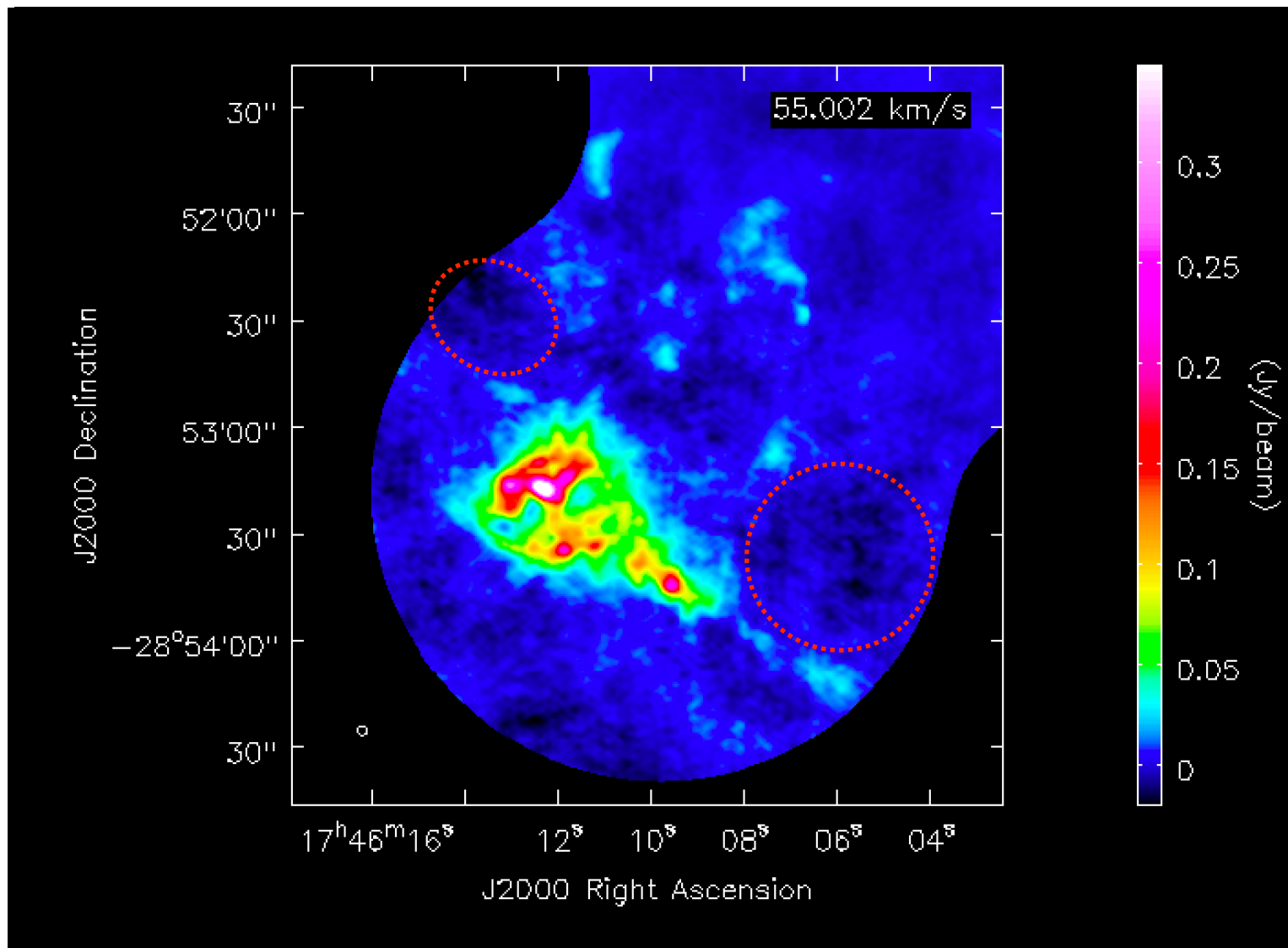
Imaging Artifacts in Interferometer Data



Imaging Artifacts in Interferometer Data



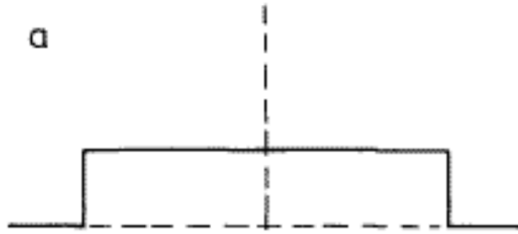
Imaging Artifacts in Interferometer Data



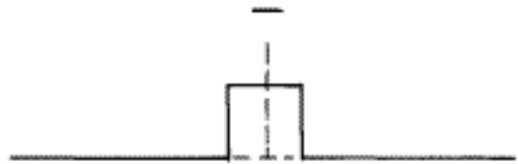
Science behind ‘Negative Bowls’

- UV plane

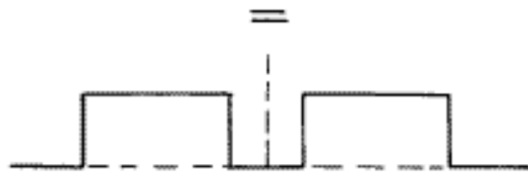
ideal



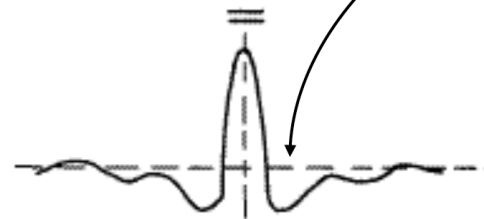
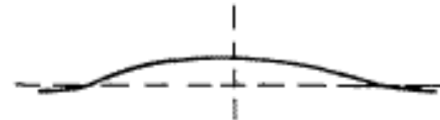
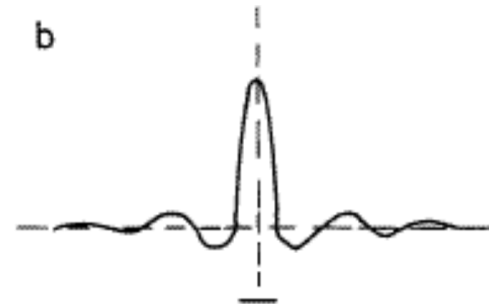
Central hole



Typical interferometer



- PSF

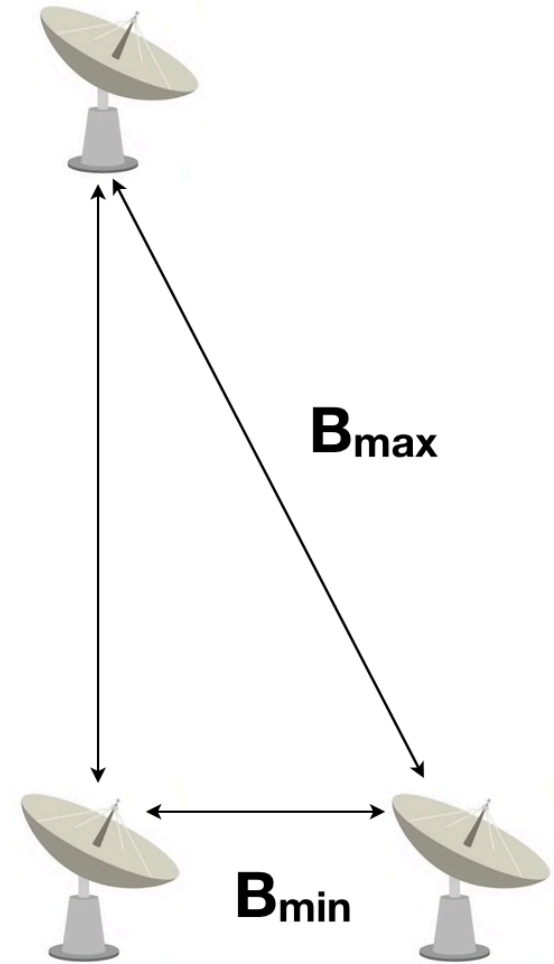
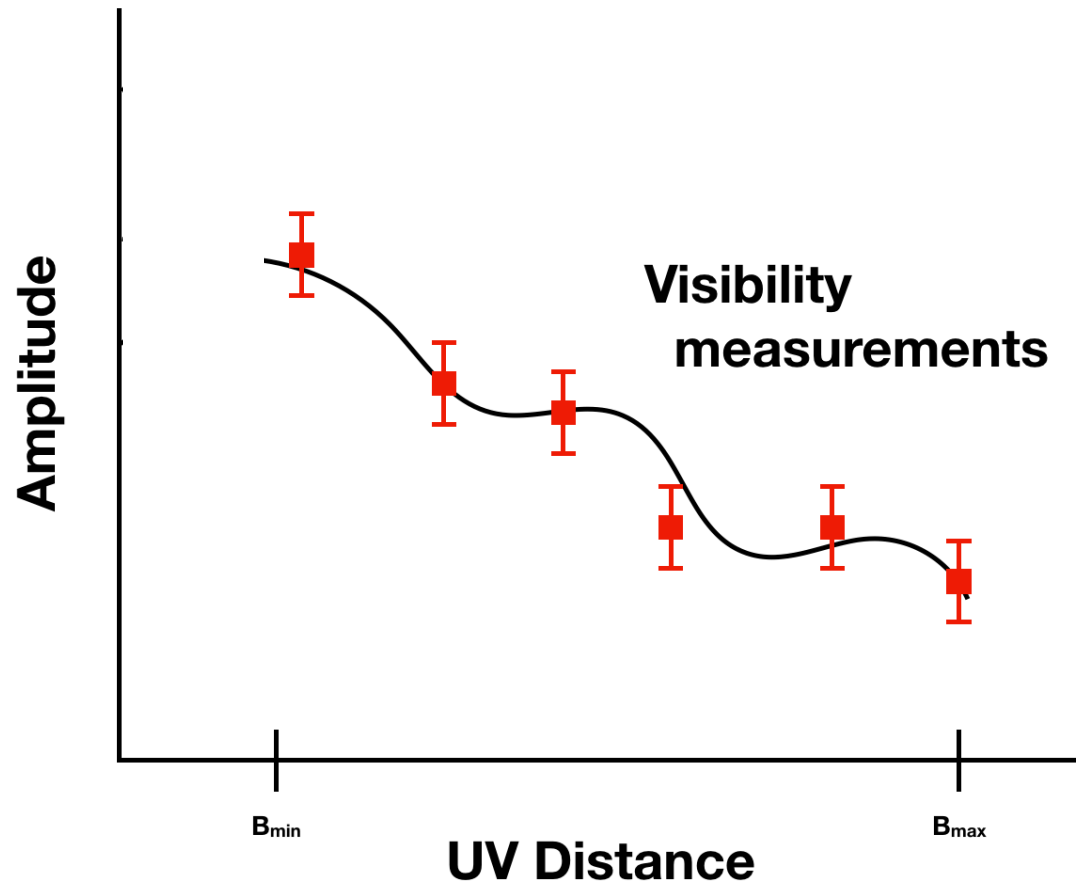


- Braun & Walterbos (1985)

- Negative ‘bowl’

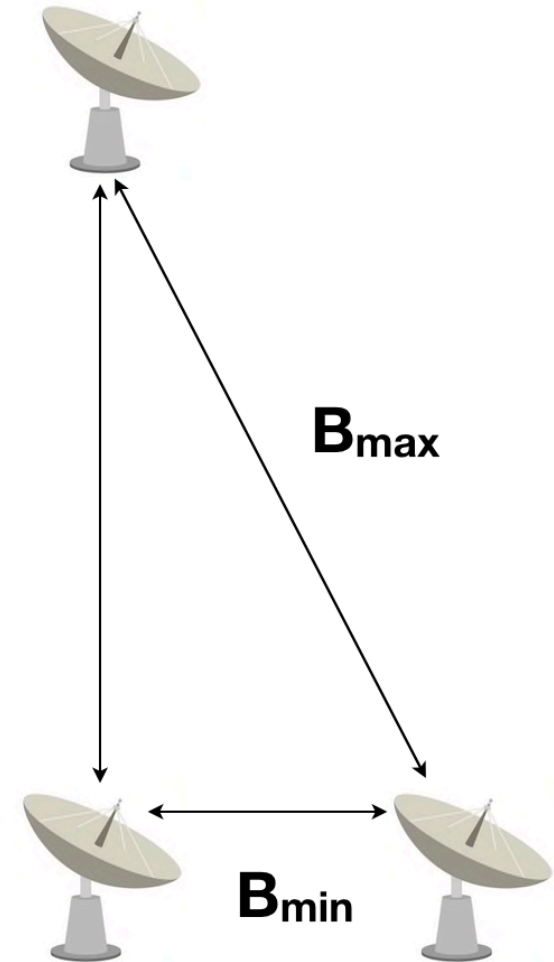
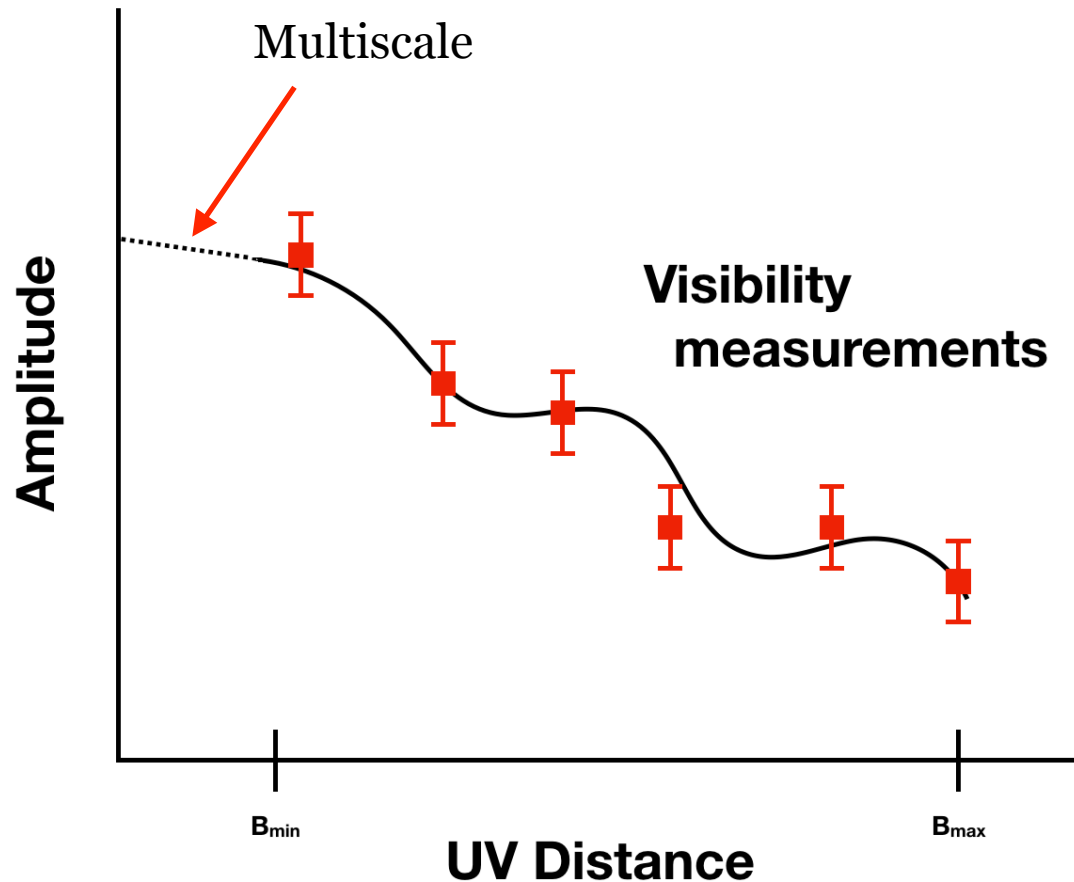
Slide Credit: Brian mason

Measuring Visibilities

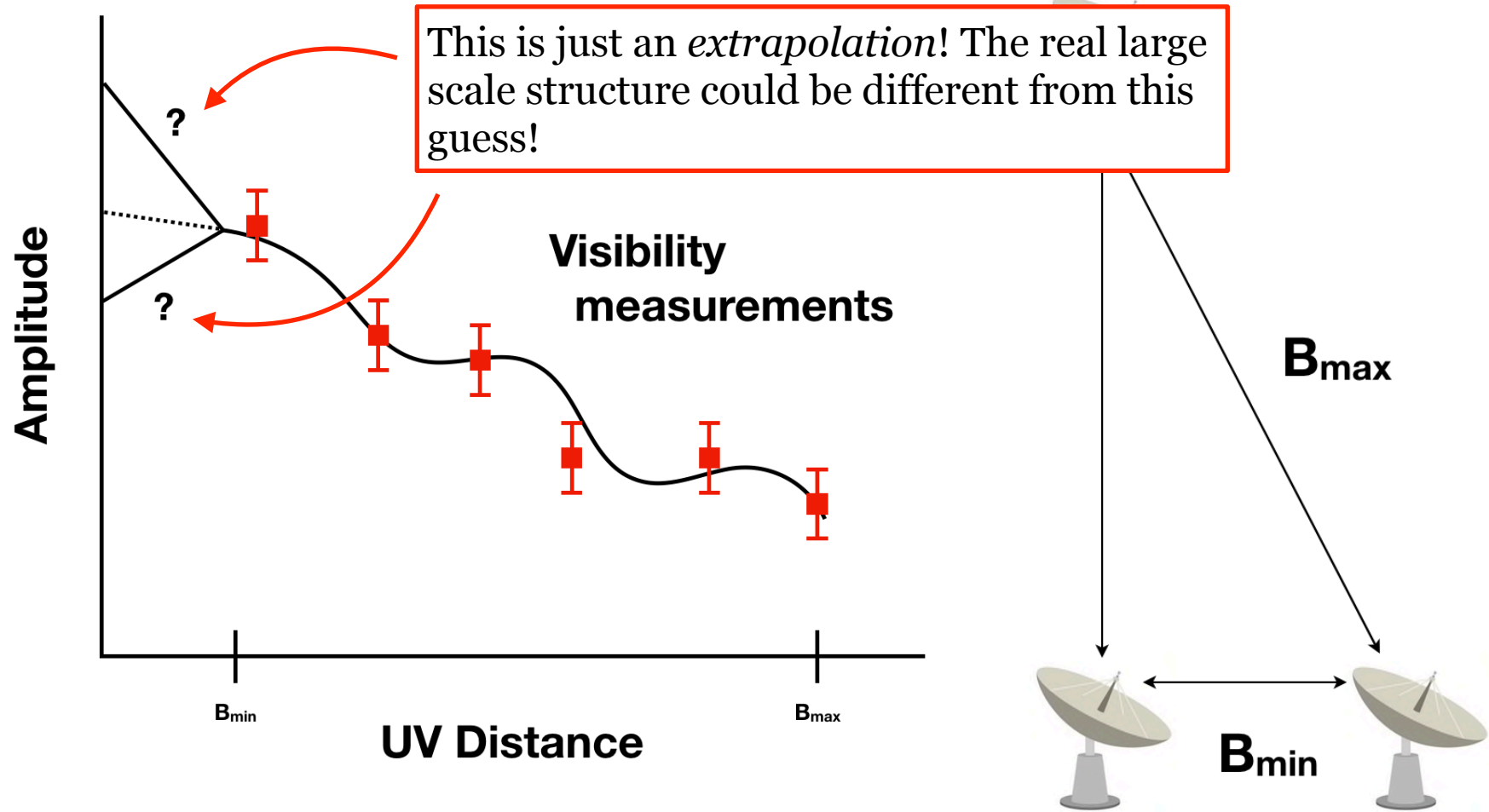


Measuring Visibilities

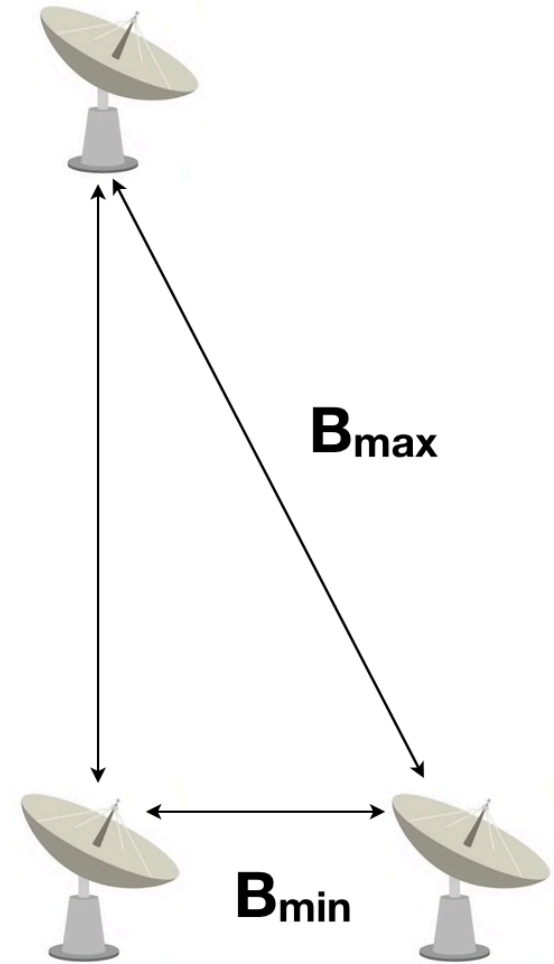
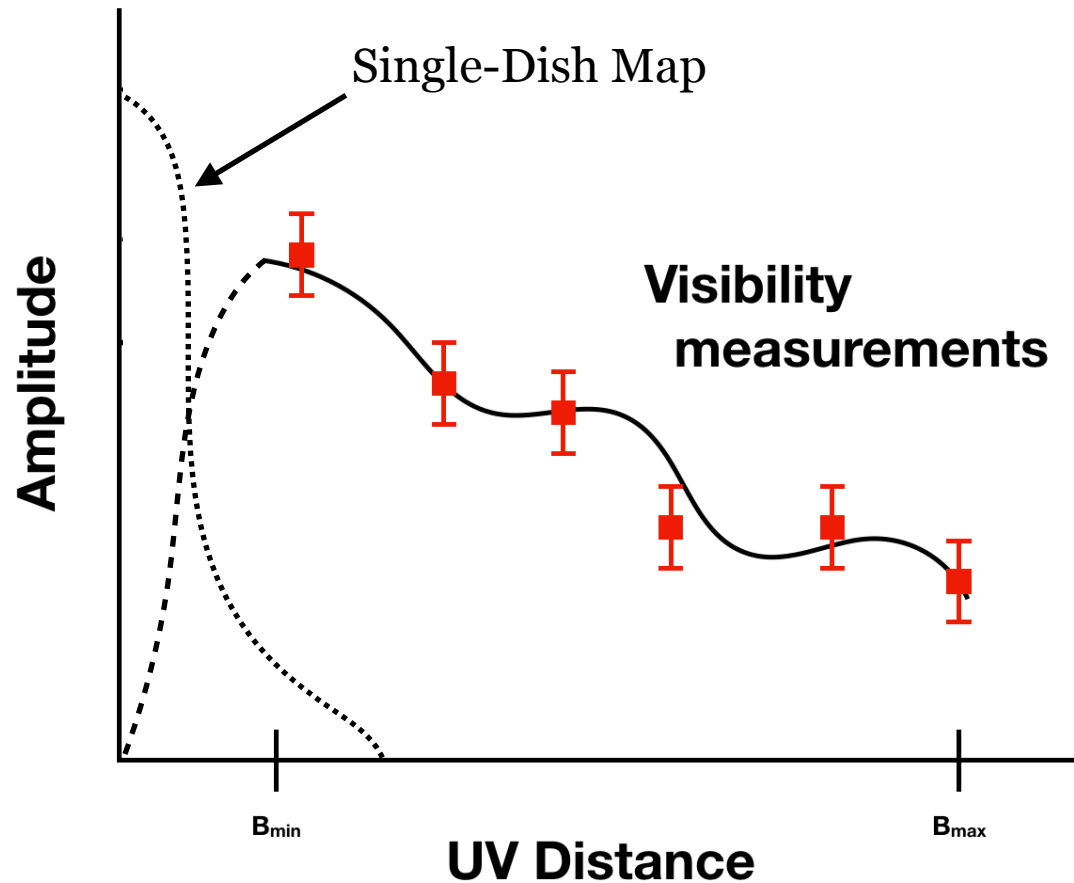
Multiscale: CASA clean parameter;
recover some large scale structure.



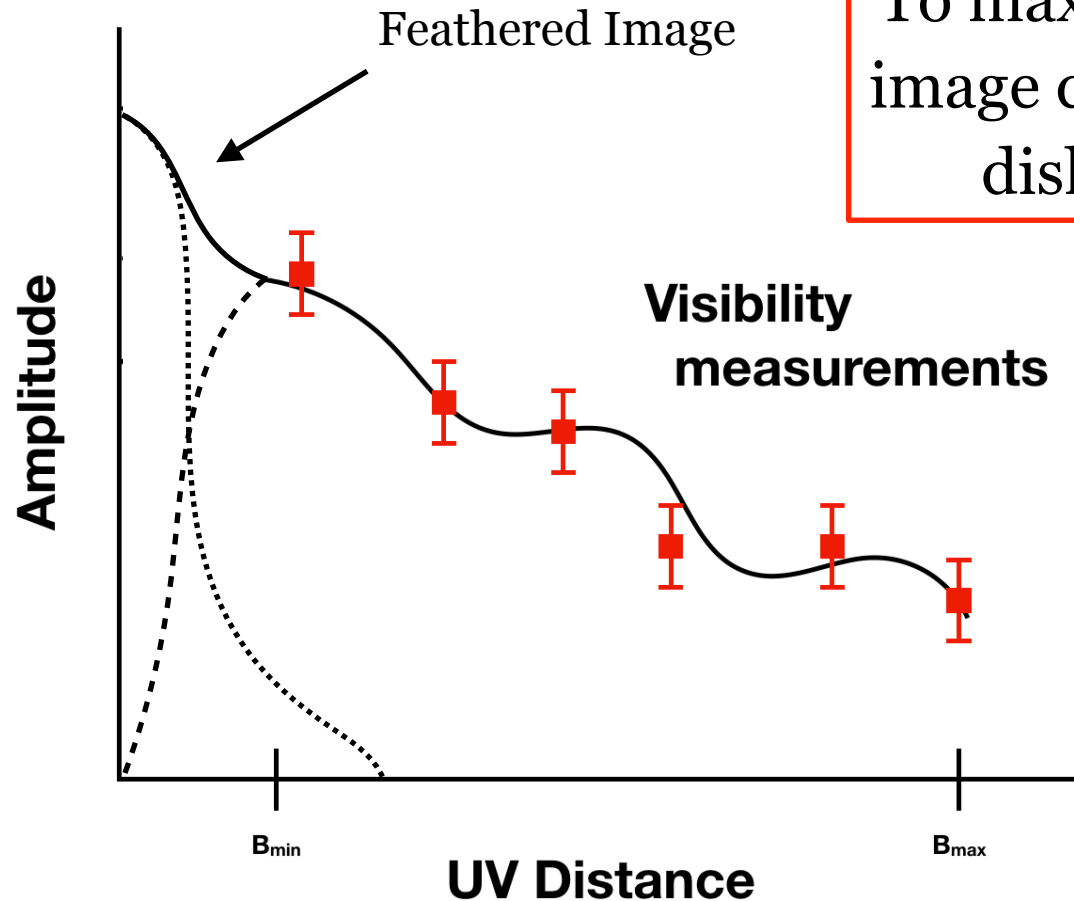
Measuring Visibilities



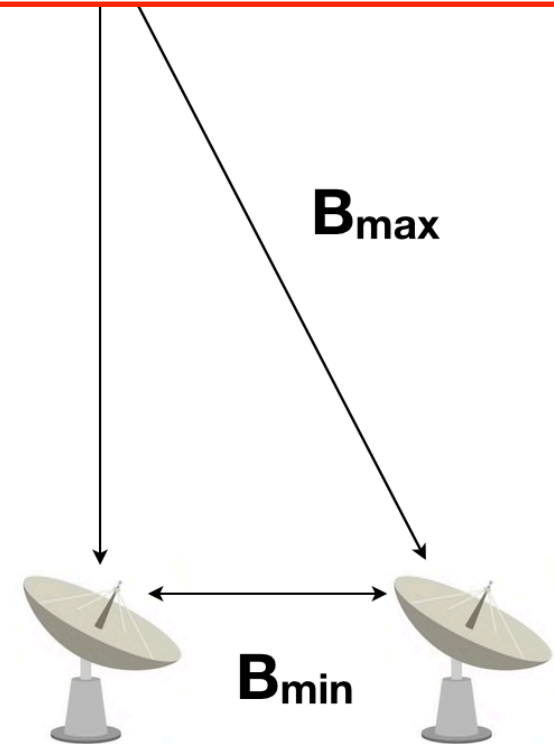
Measuring Visibilities



Measuring Visibilities

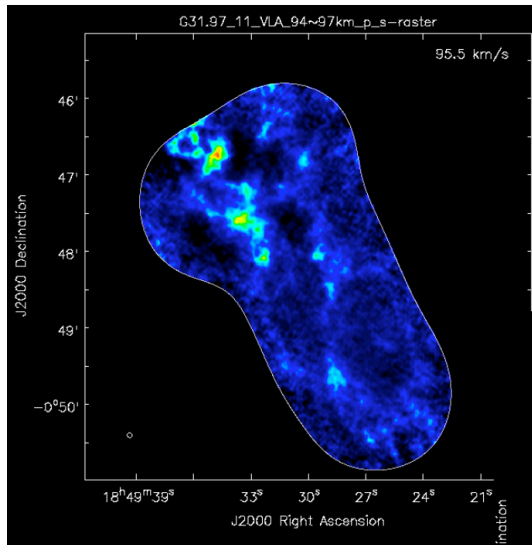


To maximize flux recovery and image quality you want a single dish size of $D > 1.5 \times B_{min}$



Example of Feathering

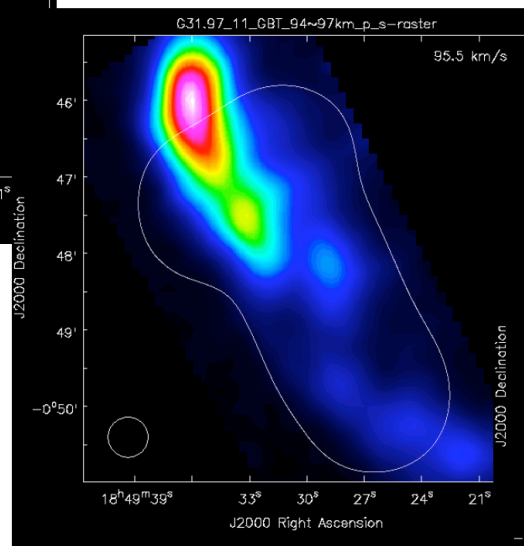
EVLA NH₃ (multi-scale CLEANed)



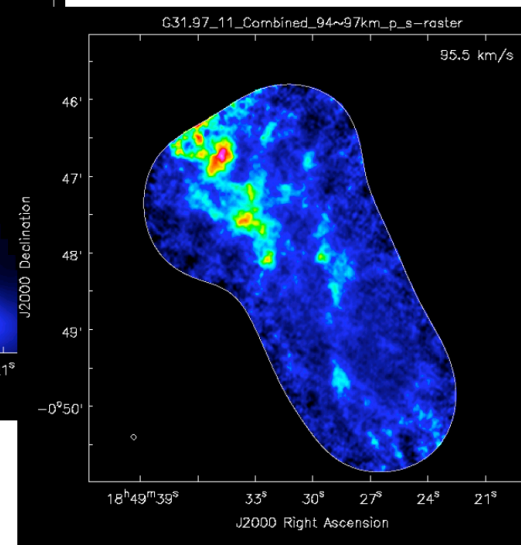
DiRienzo et al. (2015)

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GBT NH₃

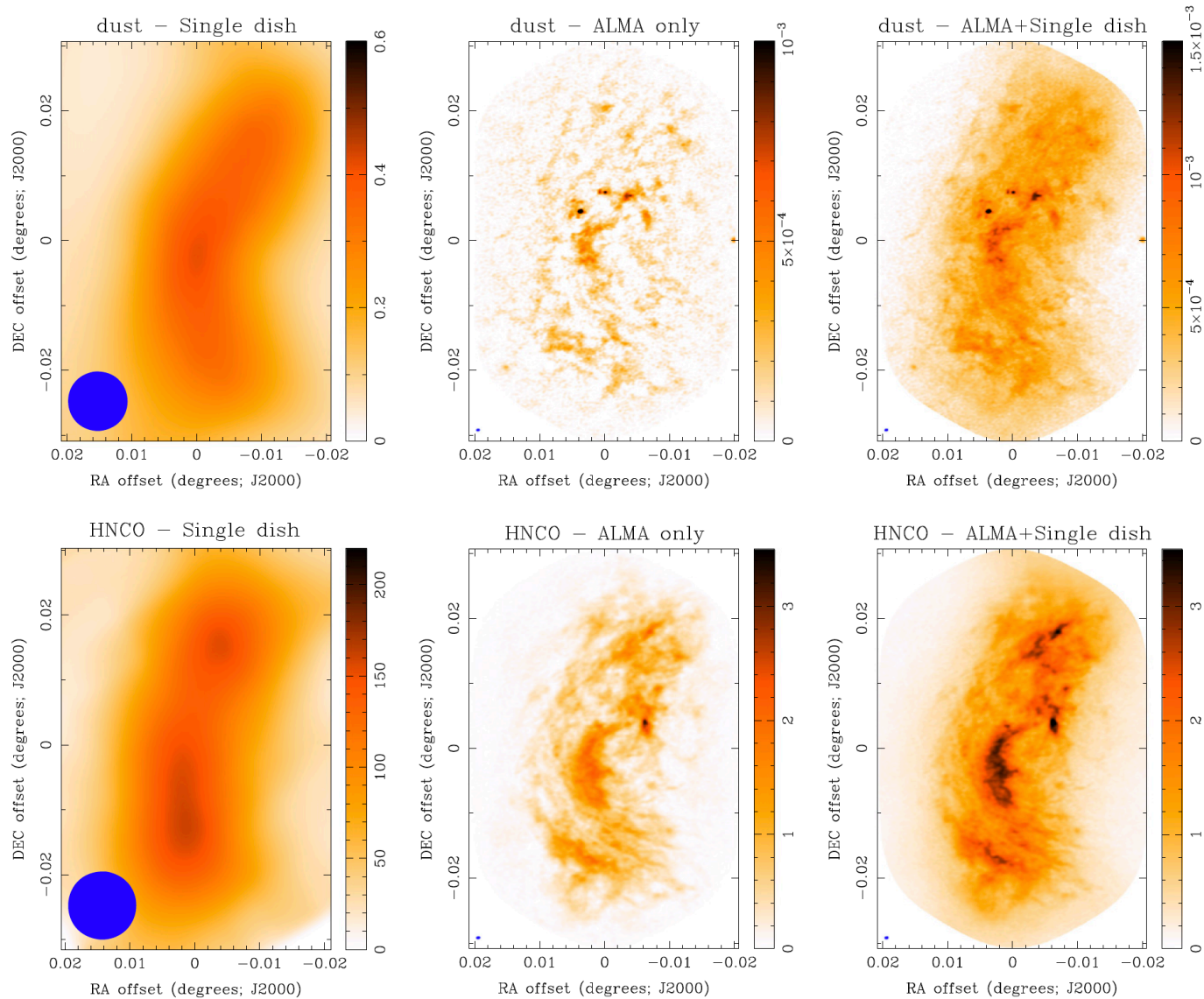


Feathered

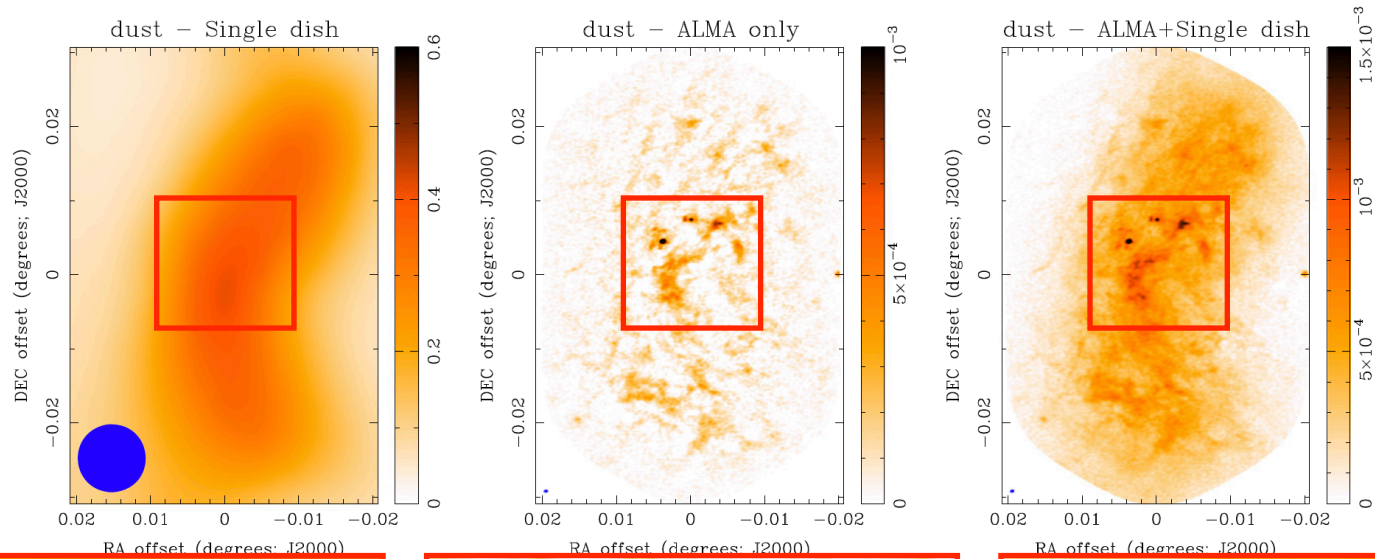


Slide Credit: Brian mason

Rathborne et al. (2014) Example



Rathborne et al. (2014) Example



Single-Dish

Interferometer

Feathered

Valid Flux Measurements

Missing flux from large scale structures can effect measurements!

THE ASTROPHYSICAL JOURNAL, 805:72 (25pp), 2015 May 20

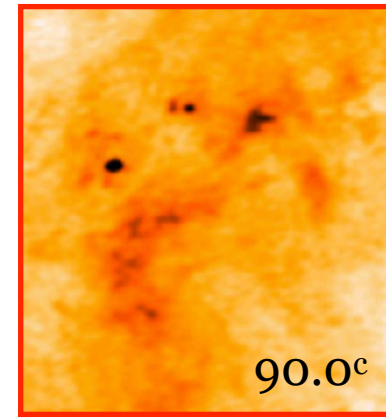
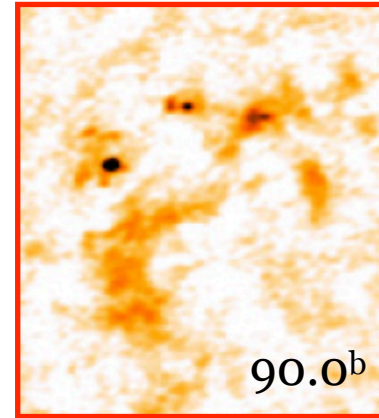


Table 3
Continuum Regions

	Measured Flux (mJy) ^a						Spectral Index				
	Area	Cont.	24.1	25.4	27.5	36.4	90.0 ^b	90.0 ^c	(24–90 GHz)		log N_{Lyc}
	(sq'')	Level	(GHz)	(GHz)	(GHz)	(GHz)	(GHz)	(GHz)	Uncorrected ^b	Corrected ^c	(phot s ^{−1})
C1	35.7	10σ	4.6 ± 0.1	4.6 ± 0.2	4.2 ± 0.2	2.6 ± 0.1	3.1 ± 0.1	6.6 ± 0.3	−0.29 ± 0.01	0.27 ± 0.03	46.5
C2	279.1	6σ	18.8 ± 0.1	18.7 ± 0.1	15.6 ± 0.1	8.2 ± 0.2	10.5 ± 0.1	27.3 ± 0.3	−0.43 ± 0.01	0.28 ± 0.03	47.2
C3	27.6	10σ	2.3 ± 0.1	2.4 ± 0.1	2.1 ± 0.1	2.6 ± 0.2	5.9 ± 0.1	11.3 ± 0.6	0.68 ± 0.01	1.17 ± 0.05	45.9
C4	14.8	6σ	1.0 ± 0.1	0.9 ± 0.1	0.9 ± 0.1	0.5 ± 0.1	1.9 ± 0.1	4.2 ± 0.5	0.52 ± 0.09	1.1 ± 0.1	45.9
C5	16.1	10σ	6.3 ± 0.1	5.6 ± 0.2	4.0 ± 0.1	NA	1.0 ± 0.2	1.9 ± 0.4	−1.31 ± 0.03	−0.86 ± 0.05	...
C6	81.6	6σ	4.5 ± 0.1	4.2 ± 0.1	5.5 ± 0.1	6.1 ± 0.1	11.8 ± 0.1	26.6 ± 0.5	0.74 ± 0.07	1.34 ± 0.09	46.5
C7	161.9	6σ	10.7 ± 0.1	8.4 ± 0.2	10.1 ± 0.2	5.9 ± 0.2	8.3 ± 0.1	28.4 ± 0.4	−0.1 ± 0.15	0.8 ± 0.18	46.9
C8	164.2	6σ	8.4 ± 0.1	5.7 ± 0.1	6.3 ± 0.1	2.8 ± 0.1	6.4 ± 0.1	32.3 ± 0.5	−0.1 ± 0.17	1.1 ± 0.14	46.8
C9	521.8	6σ	43.4 ± 0.1	34.1 ± 0.2	37.7 ± 0.2	35.8 ± 0.2	24.4 ± 0.3	80.9 ± 0.3	−0.3 ± 0.25	0.6 ± 0.3	47.5
C10	7.7	10σ	1.6 ± 0.1	1.6 ± 0.2	NA	NA	0.6 ± 0.1	1.0 ± 0.1	−0.73 ± 0.03	−0.35 ± 0.01	...

^a “NA” indicates this region was outside or near the edge of the field of view.

^b Values from 3 mm ALMA-only image of Rathborne et al. (2014b).

^c Values from single-dish-corrected ALMA image of Rathborne et al. (2014b).

Mills et al. (2015)



How to Feather Images

CASA task: feather

- Parameters:
 - ‘lowres’ - Low-resolution image
 - ‘highres’ - High-resolution image
 - ‘imagename’ - output feathered image
- Assumptions:
 - Overlapping spatial frequencies
 - Well-defined primary beams
- Requirements:
 - Low resolution, single dish (SD) image
 - High-resolution, interferometric (IF) image
 - Primary beam of the high-resolution image (PB)

How to Feather Images

CASA task: feather

See Hoffmann & Kepley (2018)
for more details (GBT memo #300)

- Parameters:
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 - Overlapping spatial frequencies
 - Well-defined primary beams
- Requirements:
 - Low resolution, single dish (SD) image
 - High-resolution, interferometric (IF) image
 - Primary beam of the high-resolution image (PB)

Baselines for ALMA and VLA

Table A-1: Angular Resolutions (AR) and Maximum Recoverable Scales (MRS) for the Cycle 7 Array configurations

Config	Lmax		Band 3	Band 4	Band 5	Band 6	Band 7	Band 8	Band 9	Band 10
	Lmin		100 GHz	150 GHz	183 GHz	230 GHz	345 GHz	460 GHz	650 GHz	870 GHz
7-m Array	45 m	AR	12.5"	8.4"	6.8"	5.4"	3.6"	2.7"	1.9"	1.4"
	9 m	MRS	66.7"	44.5"	36.1"	29.0"	19.3"	14.5"	10.3"	7.7"
C43-1	161 m	AR	3.4"	2.3"	1.8"	1.5"	1.0"	0.74"	0.52"	0.39"
	15 m	MRS	28.5"	19.0"	15.4"	12.4"	8.3"	6.2"	4.4"	3.3"
C43-2	314 m	AR	2.3"	1.5"	1.2"	1.0"	0.67"	0.50"	0.35"	0.26"
	15 m	MRS	22.6"	15.0"	12.2"	9.8"	6.5"	4.9"	3.5"	2.6"
C43-3	500 m	AR	1.4"	0.94"	0.77"	0.62"	0.41"	0.31"	0.22"	0.16"
	15 m	MRS	16.2"	10.8"	8.7"	7.0"	4.7"	3.5"	2.5"	1.9"
C43-4	784 m	AR	0.92"	0.61"	0.50"	0.40"	0.27"	0.20"	0.14"	0.11"
	15 m	MRS	11.2"	7.5"	6.1"	4.9"	3.3"	2.4"	1.7"	1.3"
C43-5	1.4 km	AR	0.54"	0.36"	0.30"	0.24"	0.16"	0.12"	0.084"	0.063"
	15 m	MRS	6.7"	4.5"	3.6"	2.9"	1.9"	1.5"	1.0"	0.77"
C43-6	2.5 km	AR	0.31"	0.20"	0.16"	0.13"	0.089"	0.067"	0.047"	0.035"
	15 m	MRS	4.1"	2.7"	2.2"	1.8"	1.2"	0.89"	0.63"	0.47"
C43-7	3.6 km	AR	0.21"	0.14"	0.11"	0.092"	0.061"	0.046"	0.033"	0.024"
	64 m	MRS	2.6"	1.7"	1.4"	1.1"	0.75"	0.56"	0.40"	0.30"
C43-8	8.5 km	AR	0.096"	0.064"	0.052"	0.042"	0.028"	N/A	N/A	N/A
	110 m	MRS	1.4"	0.95"	0.77"	0.62"	0.41"			
C43-9	13.9 km	AR	0.057"	0.038"	0.031"	0.025"	0.017"	N/A	N/A	N/A
	368 m	MRS	0.81"	0.54"	0.44"	0.35"	0.24"			
C43-10	16.2 km	AR	0.042"	0.028"	0.023"	0.018"	0.012"	N/A	N/A	N/A
	244 m	MRS	0.50"	0.33"	0.27"	0.22"	0.14"			

ALMA

11 arrays

4 arrays

VLA

GBO = 100 m
Arecibo = 300 m

Table 3.1.1: Configuration Properties

Configuration	A	B	C	D
B _{max} (km ¹)	36.4	11.1	3.4	1.03
B _{min} (km ¹)	0.68	0.21	0.035 ⁵	0.035
Band	Synthesized Beamwidth θ_{HPBW} (arcsec) ^{1,2,3}			
74 MHz (4)	24	80	260	850
350 MHz (P)	5.6	18.5	60	200
1.5 GHz (L)	1.3	4.3	14	46
3.0 GHz (S)	0.65	2.1	7.0	23
6.0 GHz (C)	0.33	1.0	3.5	12
10 GHz (X)	0.20	0.60	2.1	7.2
15 GHz (Ku)	0.13	0.42	1.4	4.6
22 GHz (K)	0.089	0.28	0.95	3.1
33 GHz (Ka)	0.059	0.19	0.63	2.1
45 GHz (Q)	0.043	0.14	0.47	1.5
Band	Largest Angular Scale θ_{LAS} (arcsec) ^{1,4}			
74 MHz (4)	800	2200	20000	20000
350 MHz (P)	155	515	4150	4150
1.5 GHz (L)	36	120	970	970
3.0 GHz (S)	18	58	490	490
6.0 GHz (C)	8.9	29	240	240
10 GHz (X)	5.3	17	145	145
15 GHz (Ku)	3.6	12	97	97
22 GHz (K)	2.4	7.9	66	66
33 GHz (Ka)	1.6	5.3	44	44
45 GHz (Q)	1.2	3.9	32	32



Baselines for ALMA and VLA

Table A-1: Angular Resolutions (AR) and Max

Config	Lmax		Band 3	Band 4
	Lmin		100 GHz	150 GHz
7-m Array	45 m	AR	12.5"	8.4"
	9 m	MRS	66.7"	44.5"
C43-1	161 m	AR	3.4"	2.3"
	15 m	MRS	28.5"	19.0"
C43-2	314 m	AR	2.3"	1.5"
	15 m	MRS	22.6"	15.0"
C43-3	500 m	AR	1.4"	0.94"
	15 m	MRS	16.2"	10.8"
C43-4	784 m	AR	0.92"	0.61"
	15 m	MRS	11.2"	7.5"
C43-5	1.4 km	AR	0.54"	0.36"
	15 m	MRS	6.7"	4.5"
C43-6	2.5 km	AR	0.31"	0.20"
	15 m	MRS	4.1"	2.7"
C43-7	3.6 km	AR	0.21"	0.14"
	64 m	MRS	2.6"	1.7"
C43-8	8.5 km	AR	0.096"	0.064"
	110 m	MRS	1.4"	0.95"
C43-9	13.9 km	AR	0.057"	0.038"
	368 m	MRS	0.81"	0.54"
C43-10	16.2 km	AR	0.042"	0.028"
	244 m	MRS	0.50"	0.33"

by configurations



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10 GHz (X)	5.3	17	145	145
15 GHz (Ku)	3.6	12	97	97
22 GHz (K)	2.4	7.9	66	66
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Configuration	A	B	C	D
1	36.4	11.1	3.4	1.03
			0.035 ⁵	0.035
Bandwidth $\theta_{\text{HPBW}}(\text{arcsec})^{1,2,3}$				
	260	850		
	60	200		
	14	46		
	7.0	23		
	3.5	12		
10 GHz (X)	0.20	0.60	2.1	7.2
15 GHz (Ku)	0.13	0.42	1.4	4.6
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10 GHz (X)	5.3	17	145	145
15 GHz (Ku)	3.6	12	97	97
22 GHz (K)	2.4	7.9	66	66
33 GHz (Ka)	1.6	5.3	44	44
45 GHz (Q)	1.2	3.9	32	32

4 arrays

VLA

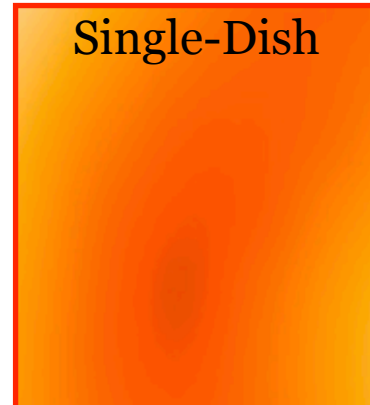
GBO = 100 m
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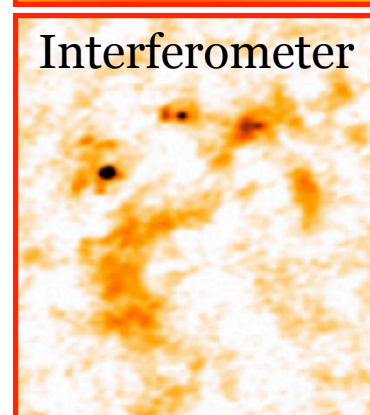
Summary

- Combination of SD+IF data is ‘feathering’
- To maximize flux recovery and image quality you want a single dish size of $D > 1.5 \times B_{\min}$
 - For the GBT: VLA arrays D & C; ALMA arrays C43-1 —C43-7
 - For Arecibo: VLA array D, C, & B
- CASA task: ‘Feather’
- Valid Flux Measurements
 - Need single-dish data to get valid flux measurements

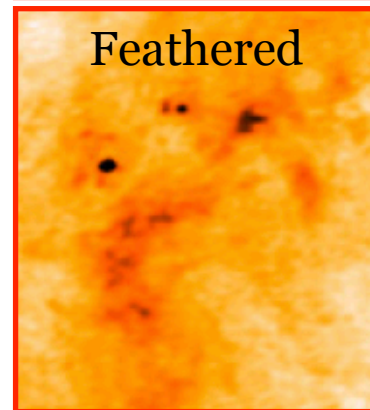
Single-Dish



Interferometer



Feathered



Additional References

- Rathborne et al. (2014)
 - Example of combining ALMA and Mopra Data
- Hoffmann & Kepley (2018)
 - GBT memo #300: Correcting ALMA 12 m Array Data for missing short spacings using the GBT
 - CASA scripts
- CASA Guide: 'M100 Band3 Combine'
 - Steps on combining TP+7m+12m ALMA data
- DiRienzo et al. (2015)
 - Feathering VLA+GBT data
- Brian Mason slides on Mosaicking
- Braun & Walters (1985)
 - Science behind negative 'bowls'





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