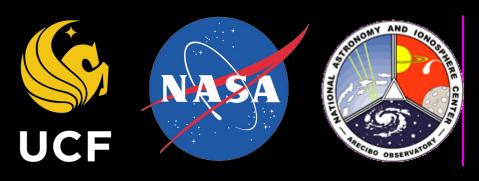
Shape Modeling of Potentially Hazardous Asteroid (85989) 199 JD6 from Radar and Lightcurve Data



Sean Marshall

And 25 others



JD6 Coauthors





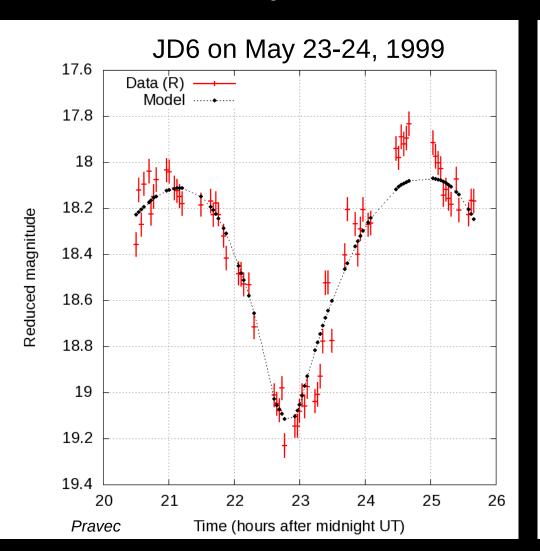


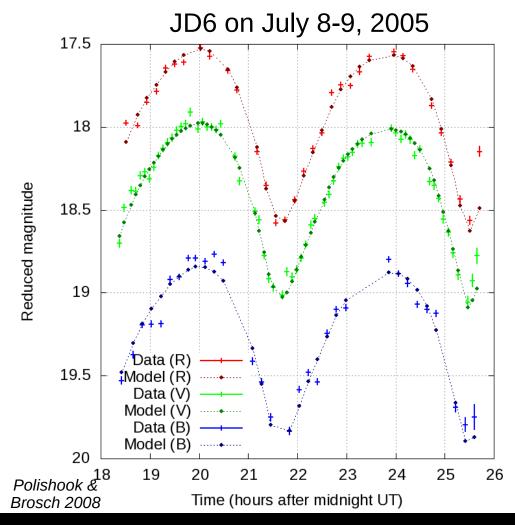


- Marina Brozović, Lance Benner, Shantanu Naidu, Jon Giorgini, Joseph Jao, Clement Lee, Michael Hicks (JPL)
- Donald Campbell (Cornell University)
- Christopher Magri (University of Maine at Farmington); Patrick Taylor, Linda Rodriguez-Ford (LPI); James Richardson (PSI); Ronald Vervack, Jr. (Johns Hopkins U. / APL); Yanga Fernandez (U. of Central Florida); Frank Ghigo, Adam Kobelski (Green Bank Obs.); Michael Busch (SETI Institute); Petr Pravec (Academy of Sciences of the Czech Republic); Benjamin Sharkey (U. of Minnesota); Ellen Howell, Michael Nolan, Jenna Crowell (U. of Arizona); Brandon Bozek (U. of Texas, Austin); Brice-Olivier Demory (U. of Bern & Geneva Obs.); Raoul Behrend (Geneva Obs.)

1999 JD6: Lightcurves

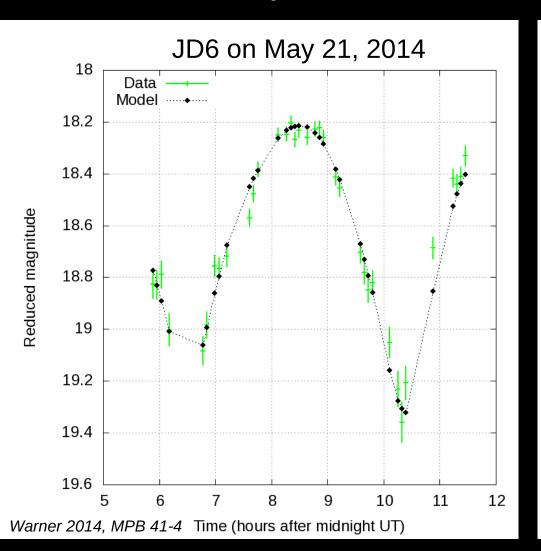
- Have lightcurves from six different years
- Rotation period ~7.66 hours

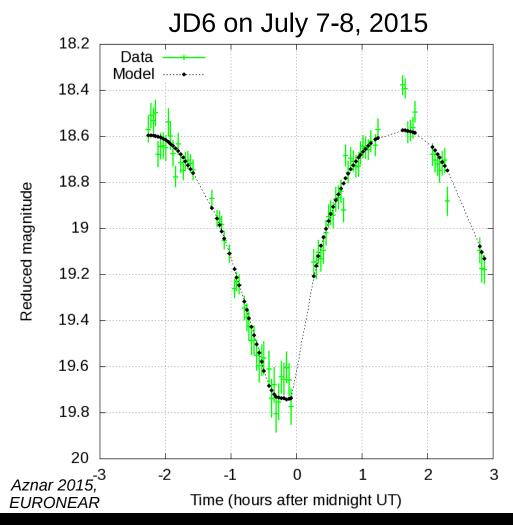




1999 JD6: Lightcurves

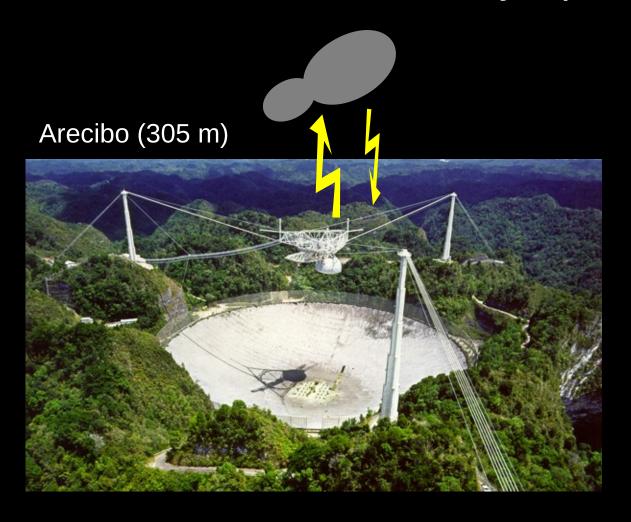
- Have lightcurves from six different years
- Rotation period ~7.66 hours

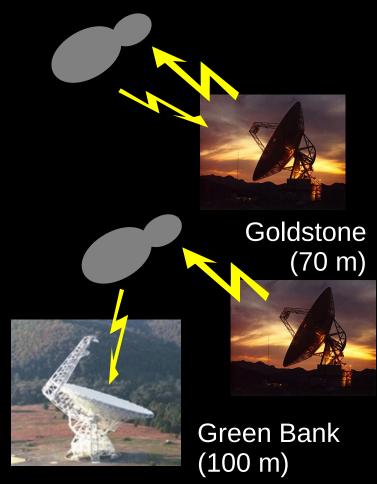




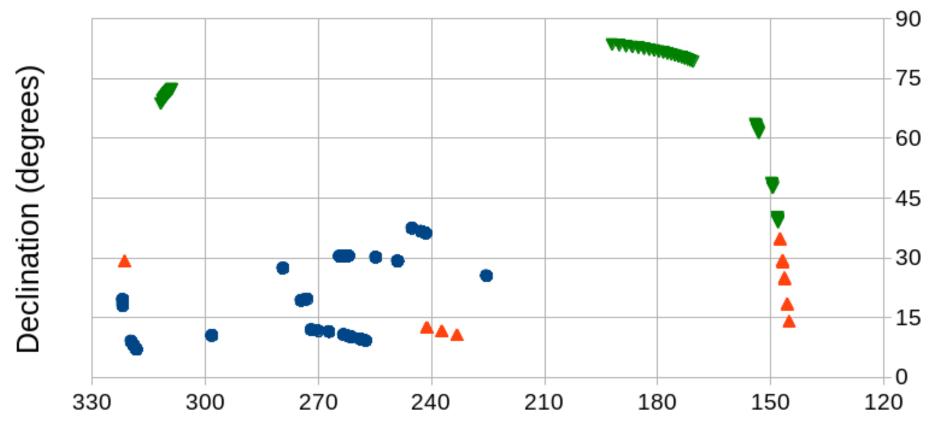
1999 JD6: Radar Observations

- From Arecibo: 3 days in 2010, 6 days in 2015
- From Goldstone: 5 days (2 bistatic) in 2015





Sky positions of 1999 JD6 during observations



Right ascension (degrees)



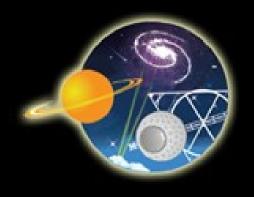




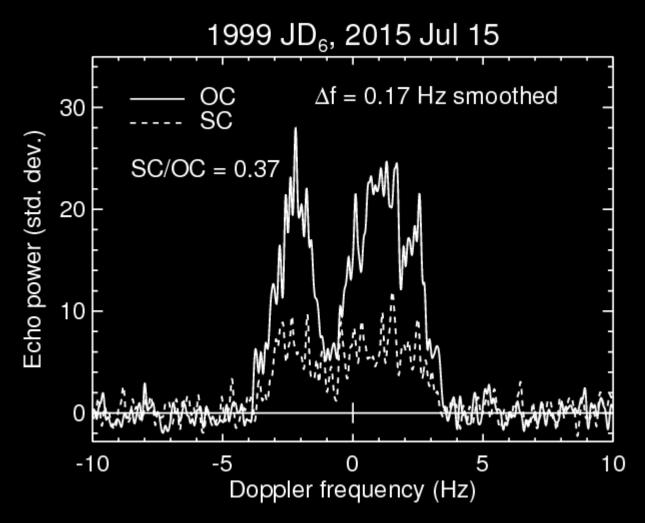


JD6: Radar Observations

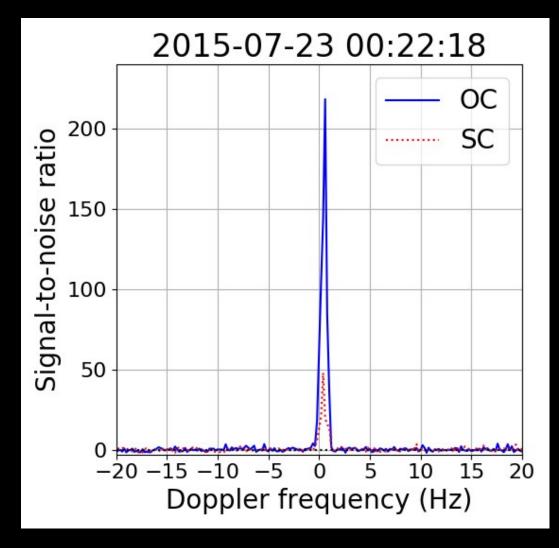
- From Arecibo (monostatic), July 15
 - Asteroid was 0.13 au away



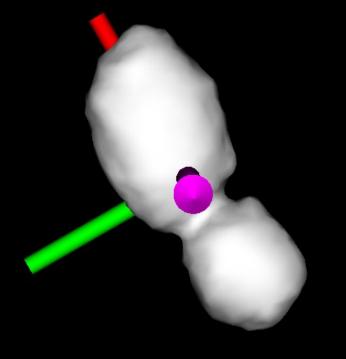




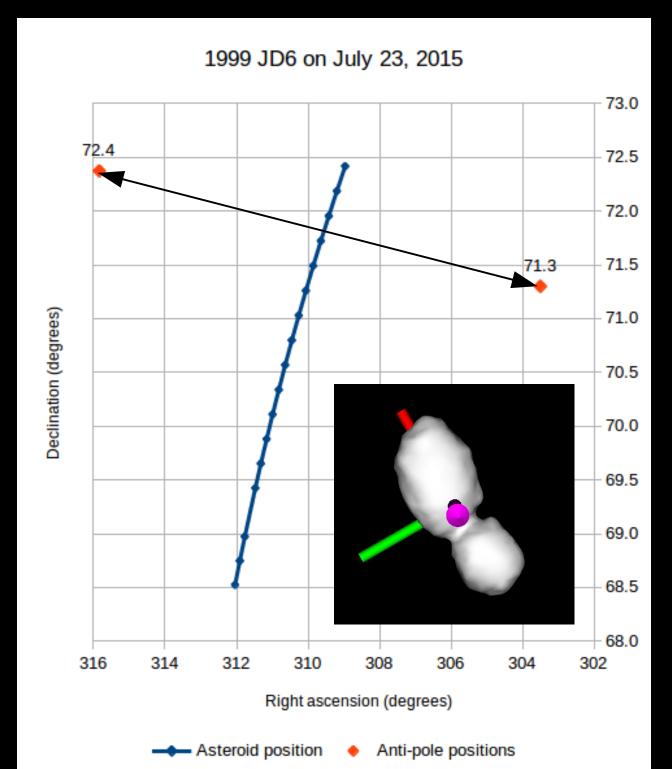
• July 23, 2015 (DSS-14 monostatic): Very narrow bandwidth! Line of sight 2° from pole



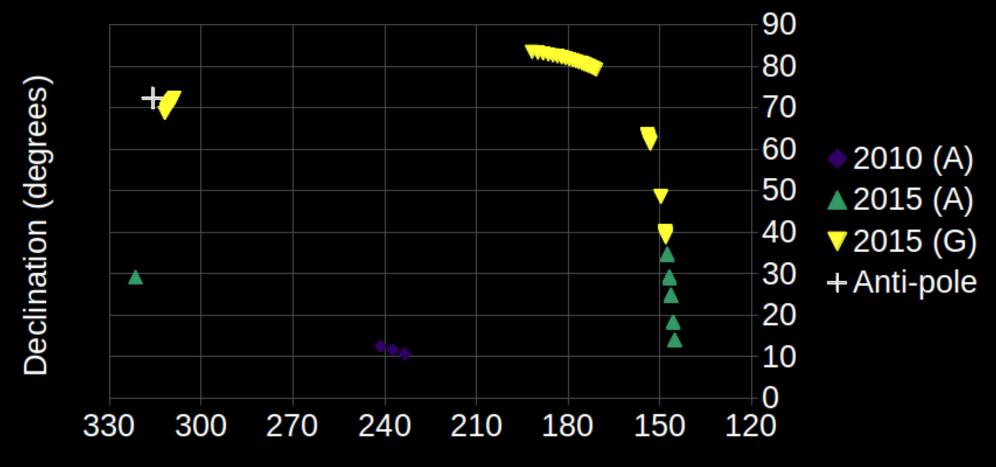
$$B = \frac{4\pi D \cos \phi}{\lambda_0 P_{spin}}$$



- July 23, 2015: Line of sight passed about 2° from the pole position
- The two marked positions are equal distances from the arc; only one of them matches the observed rotation phases



Sky position of 1999 JD6 during radar observations



Right ascension (degrees)

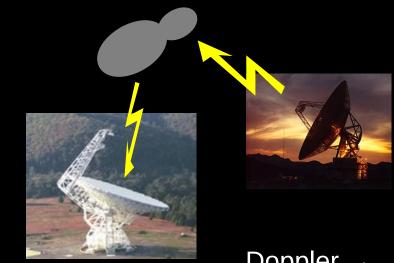




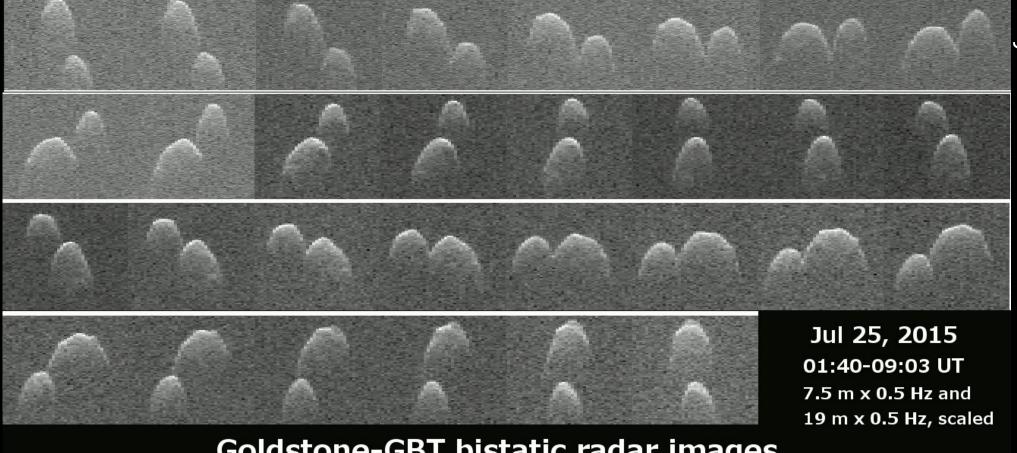


Goldstone to GBT

 Night of closest approach 0.048 au = 19 lunar distances



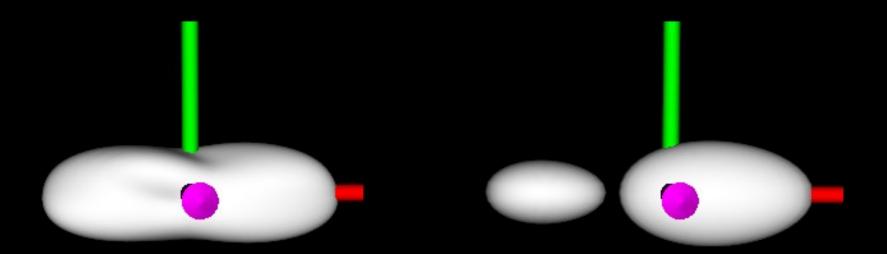
Doppler →



Goldstone-GBT bistatic radar images

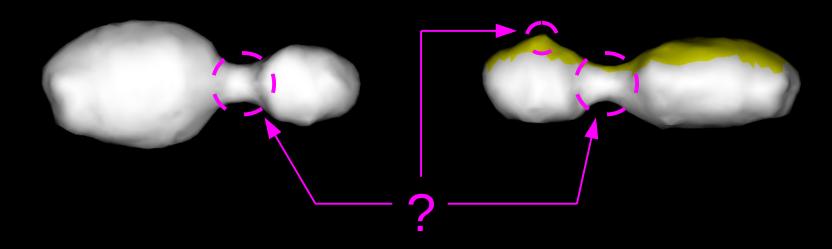
1999 JD6: First Shape Models

- Left: Spherical harmonic
- Right: Two ovoids (not overlapping)



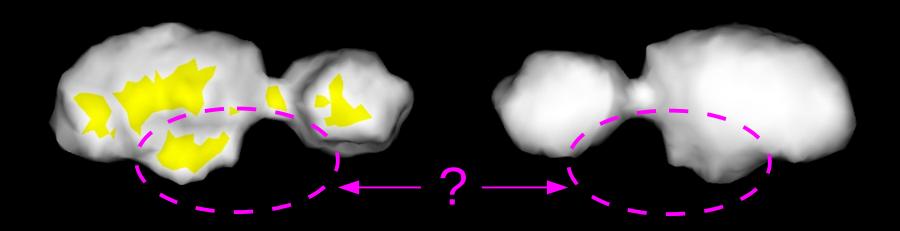
1999 JD6: Early Shape Model

- Dimensions: 3.0 km × 1.2 km × 0.9 km
 - Volumetric mean diameter 1.4 km
- Sidereal rotation period ~7.66 hours
 - Already known from lightcurves
- Pole position most likely $(\lambda, \beta) \approx (36, +77)$
 - Or possibly (220, -73)



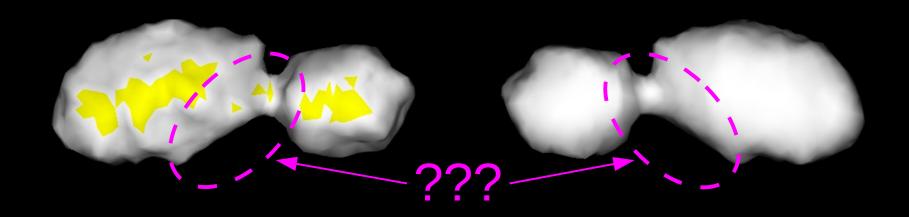
1999 JD6: Later Shape Model

- Dimensions: 2.75 km × 1.21 km × 0.98 km
 - Volumetric mean diameter: 1.33 km
- Sidereal rotation period: 7.6644 hours
 - Uncertainty: 0.0003 hours ≈ 1 second
- Pole position: $(\lambda, \beta) = (220.5^{\circ}, -73.4^{\circ}) \pm 0.5^{\circ}$



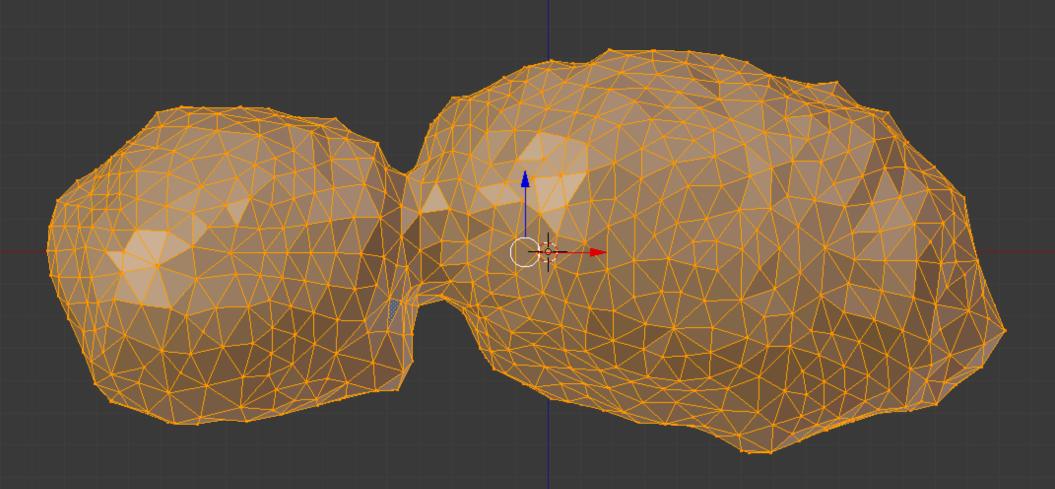
1999 JD6: Later Shape Model

- Dimensions: 2.74 km × 1.11 km × 0.99 km
 - Volumetric mean diameter: 1.34 km
- Sidereal rotation period: 7.6644 hours
 - Uncertainty: 0.0003 hours ≈ 1 second
- Pole position: $(\lambda, \beta) = (220.5^{\circ}, -73.4^{\circ}) \pm 0.5^{\circ}$



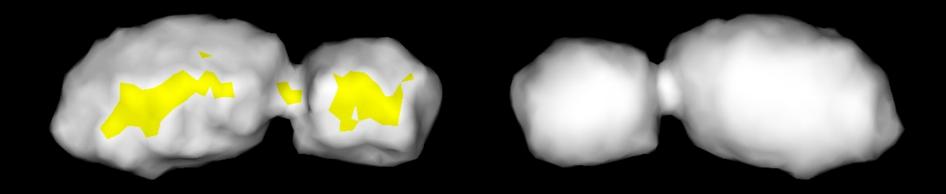
1999 JD6: Penultimate Shape Model

- Individual triangular facets, as seen with Blender
 - https://www.blender.org/



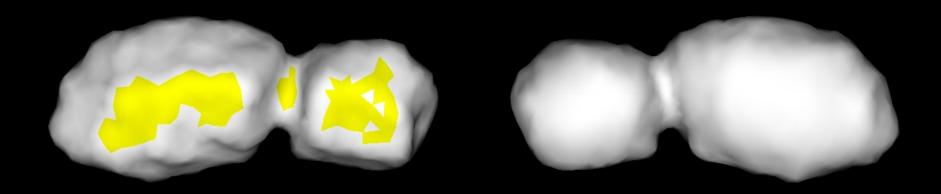
1999 JD6: Penultimate Shape Model

- Dimensions: 2.97 km × 1.25 km × 1.03 km
 - Volumetric mean diameter: 1.45 km
- Sidereal rotation period: 7.6643464 hours
 - Uncertainty: 0.0000056 hours ≈ 20 milliseconds
- Pole position: $(\lambda, \beta) = (220.3^{\circ}, -73.43^{\circ}) \pm 0.2^{\circ}$



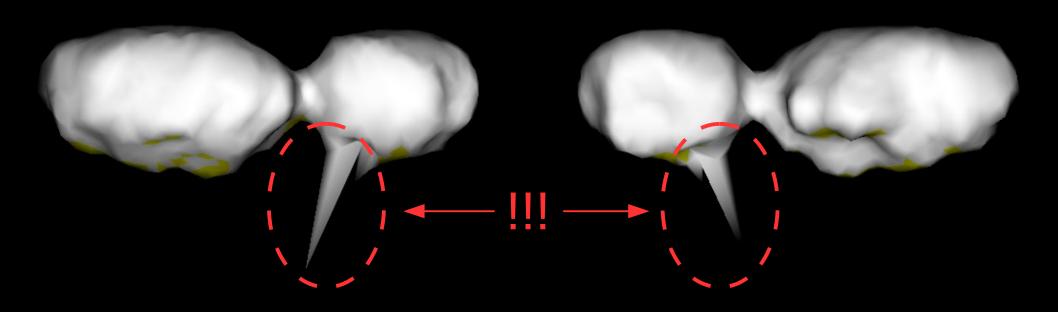
1999 JD6: Final Shape Model

- Dimensions: 2.93 km × 1.26 km × 1.05 km
 - Volumetric mean diameter: 1.48 km
- Sidereal rotation period: 7.6643463 hours
 - Uncertainty: 0.0000056 hours ≈ 20 milliseconds
- Pole position: $(\lambda, \beta) = (220.4^{\circ}, -73.41^{\circ}) \pm 0.2^{\circ}$

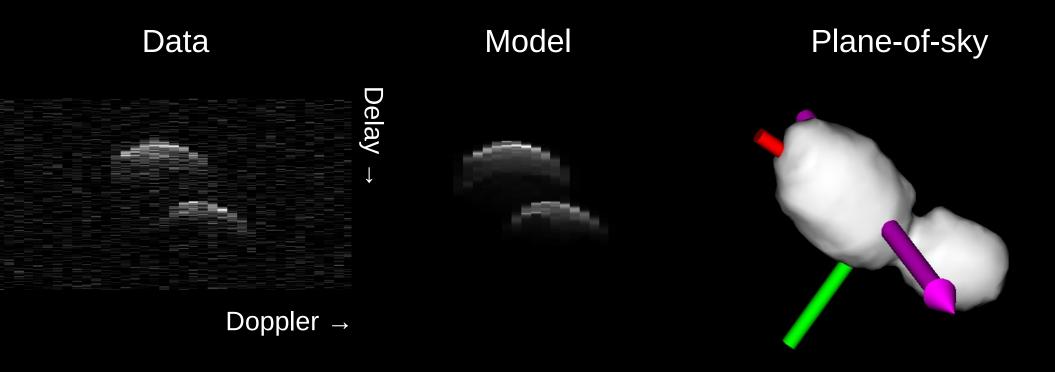


1999 JD6: Bad Shape Model

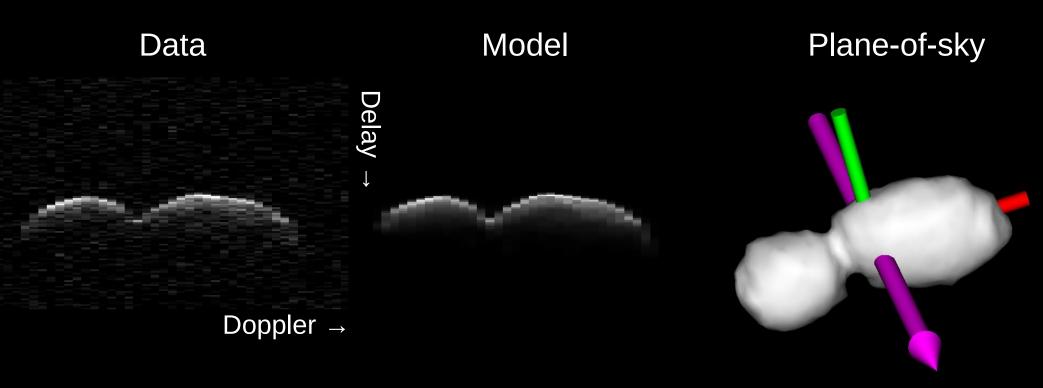
• That spike probably isn't real...



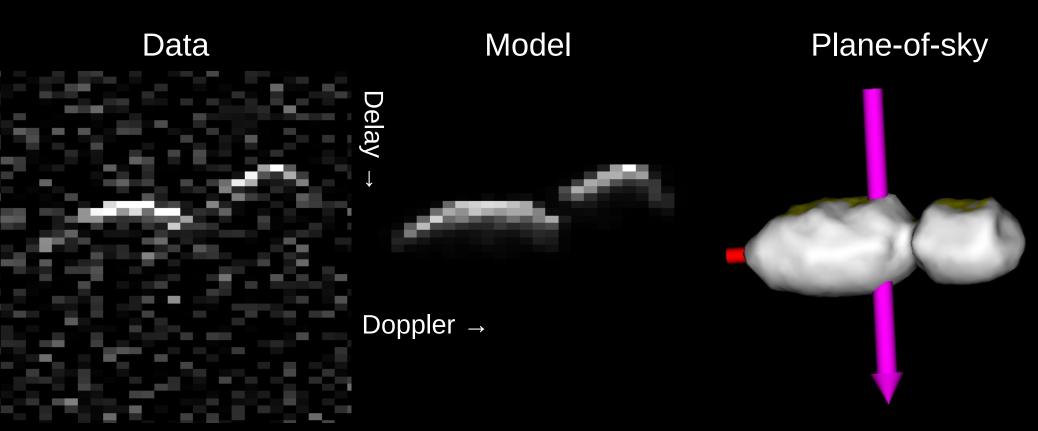
- July 25, 2015: Goldstone DSS-14 bistatic to Green Bank Telescope
 - $0.050 \mu s \times 0.50 Hz$
 - Radar image shows two lobes, but not the "neck"



- July 25, 2015: DSS-14 bistatic to GBT
 - $0.125 \,\mu s \times 0.50 \,Hz$
 - Is there really a "neck"? Or should the lobes overlap?

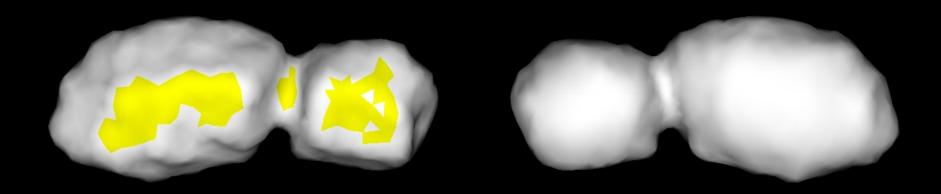


- August 2, 2015: Arecibo monostatic
 - $-0.500 \mu s \times 0.48 Hz$
 - Best views of this side were when JD6 was farther away (could only get images from Arecibo)



1999 JD6: Final Shape Model

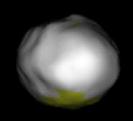
- Dimensions: 2.93 km × 1.26 km × 1.05 km
 - Volumetric mean diameter: 1.48 km
- Sidereal rotation period: 7.6643463 hours
 - Uncertainty: 0.0000056 hours ≈ 20 milliseconds
- Pole position: $(\lambda, \beta) = (220.4^{\circ}, -73.41^{\circ}) \pm 0.2^{\circ}$

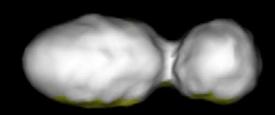


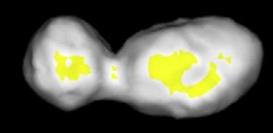
Shape Model for (8567) 1996 HW1

• From Magri et al. 2011







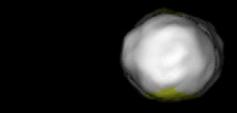


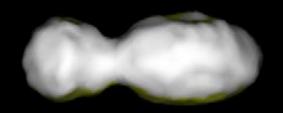
L y view from +x

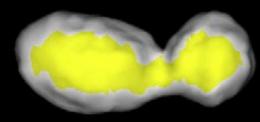


view from +y









view from -x





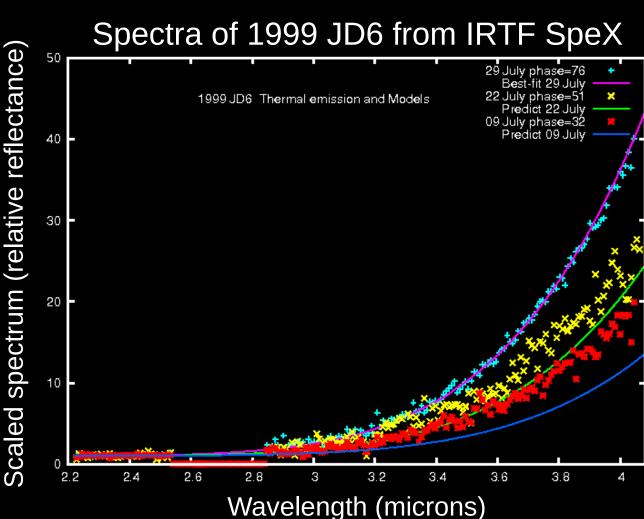


Infrared Observations

NASA

- Using the NASA InfraRed Telescope Facility (IRTF)
- SpeX
 - Full range 0.8 to 5.0 µm





(switch to third file)