

High Performance, Full-Digital Control on the LMT and KVN Telescopes

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The LMT/GTM

50m Diameter (MT-M)

1-4mm (0.85mm)

Fully steerable (Az/EI)

Mexico at 4,600m

Primary surface in five
rings (12, 24, 48, 48,
48 = 180 segments)

Three rings in use

Full-Digital Control



LMT Measurements

Main Axis

Encoders

Tilt meters

Pointing observations

M1 Segments

Laser metrology

Photogrammetry

LMT Measurements

M1 Surface:

Laser metrology

Photogrammetry

Holography (12 GHz, three elevations)

Out of focus beam maps (incl. M2)

Temperature sensors (>60)

LMT Surface Metrology

Site: Total station to install to $<500\mu\text{m}$

INAOE: 6 Metrology technicians + analyst

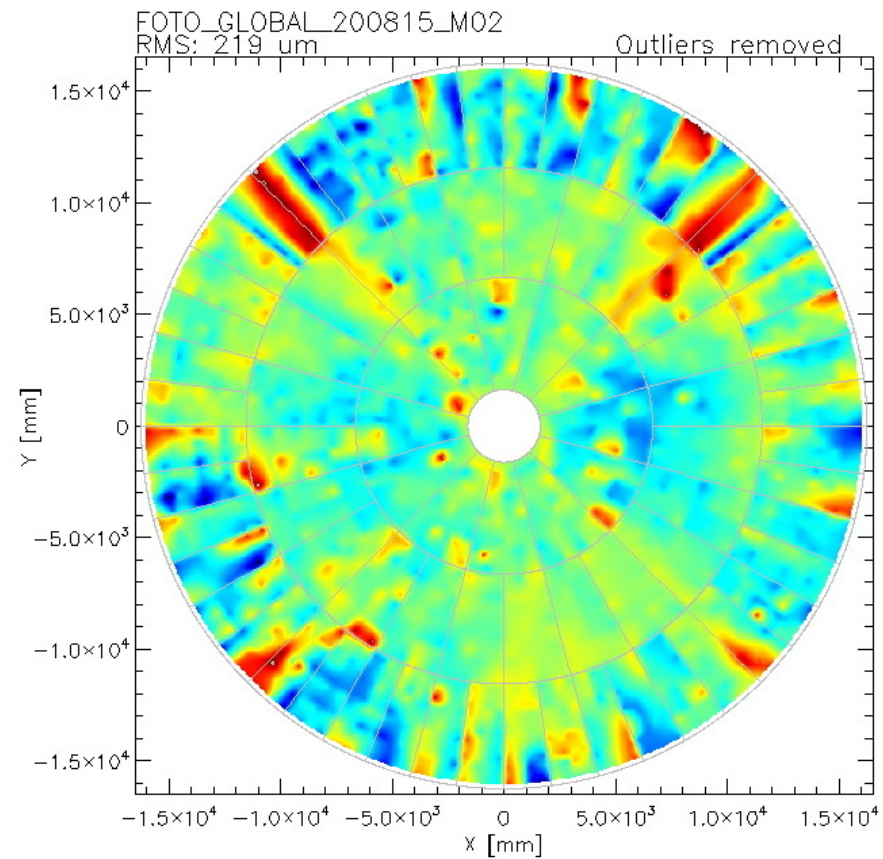
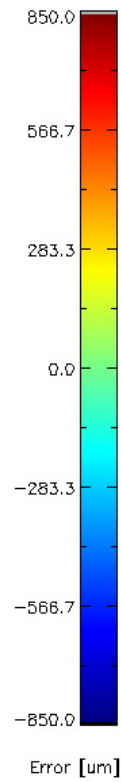
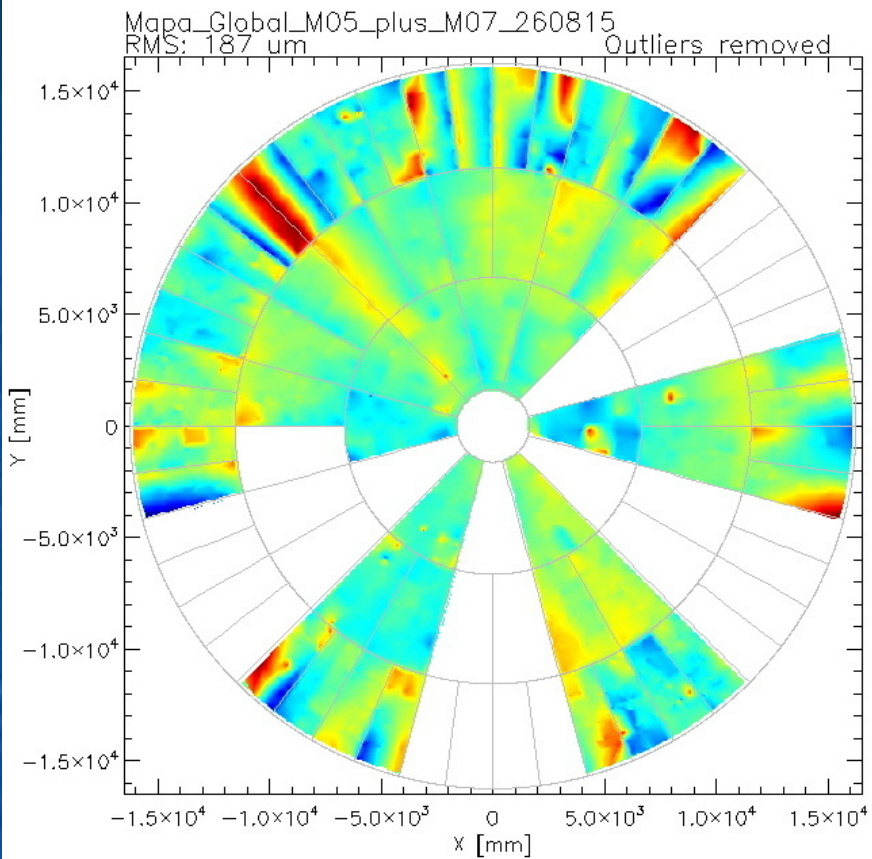
- Three laser trackers

- Photogrammetry camera

- Calibration scale bars, etc.

Segment setting to $20\mu\text{m}$

Surface mapping to $100\mu\text{m}$

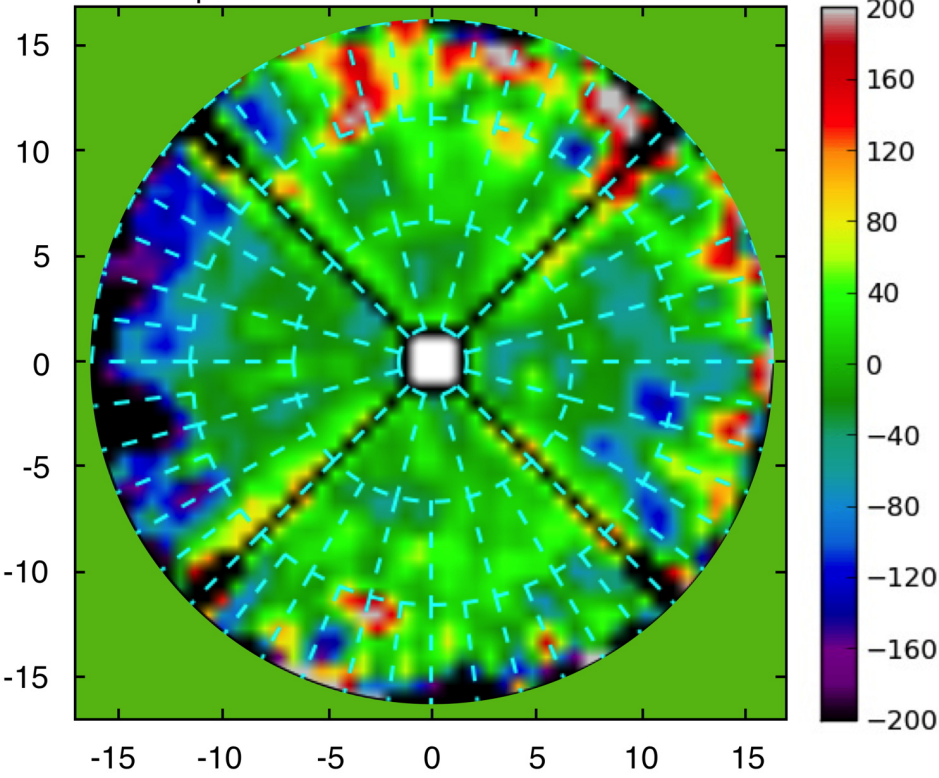


Laser tracker El 90° (5 hrs)

Photogrammetry El 90° (1 hr)

RMS: 52um

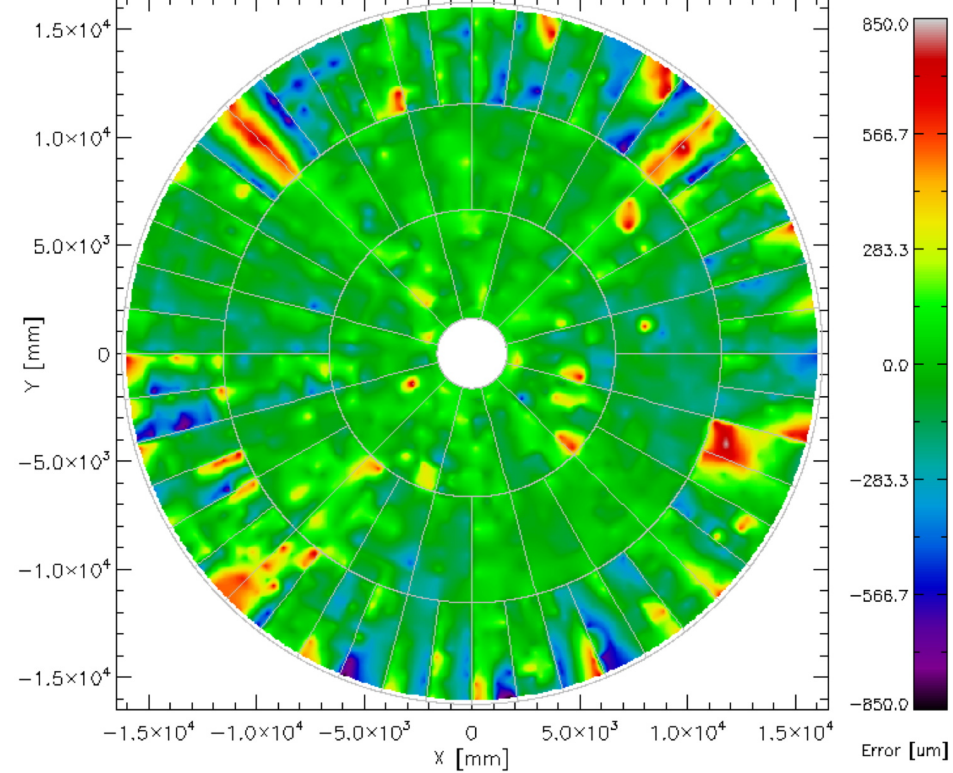
Aperture Phase Diff ObsNum=46066



FOTO_GLOBAL_M19_031015

RMS: 192 um

Outliers removed



Holography El 61° (1 hr)

Photogrammetry El 60° (1 hr)

LMT Measurements

Subreflector (M2) Surface

- Laser metrology

- Photogrammetry

M2 Relative Position

M2 Absolute Position

- Focus

- Beam Shape

- Laser metrology

LMT Actuation

Main Axis

16 Azimuth Motors (all wheels)

4 Elevation Motors

M1 Figure

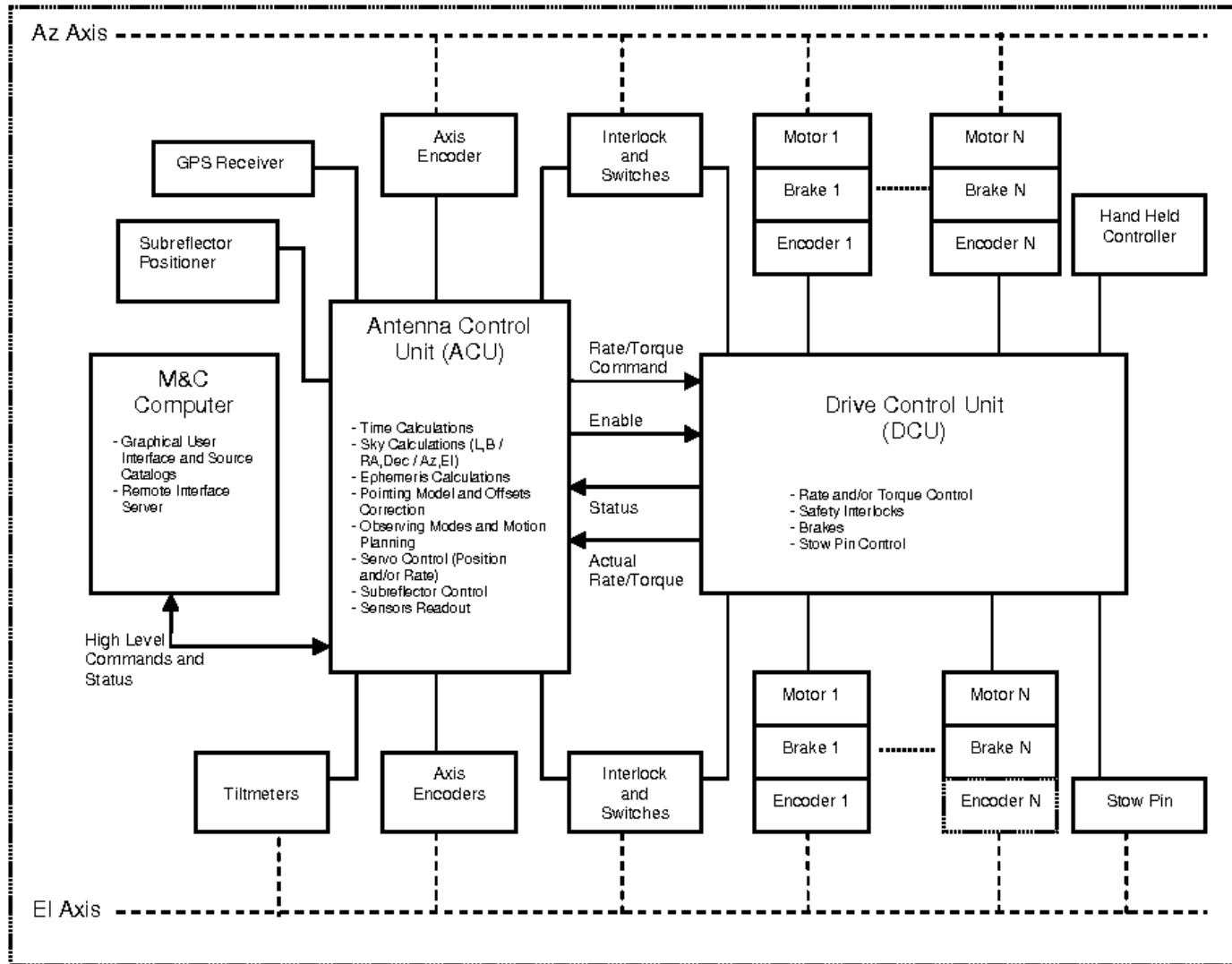
336 Segment Actuators (32.5m)

720 Segment Actuators (50m)

M2

Hexapod

LMT Servo System



Main Axis Control

Trajectory Generator

P-PI with Feedforward

Friction compensation

Future work (when we get access again!)

- State space (even elevation dependent)

- Adaptive

- Look-up table corrections

Active Surface Control

Current Implementation

- Elevation look-up table (holography)

- Additional astigmatism correction

 - Out of focus (OOF) beam maps

Future work

- Corrections based on temperature sensors

- Largest effects look like focus/astigmatism

 - Corrected with focus and OOF checks

M2 Position Control

Pointing and collimation

Position only (tilt is equivalent)

Elevation look-up

Focus

Checks during observation

LMT Performance

Main axis accuracy:

Tracking: $<0.5''$ RMS

Mapping scans: $<1''$ RMS

All-sky pointing accuracy (RMS):

Az: 3-4''

El: 6-8'' (known hysteresis in encoder mount)

Surface accuracy (all optics): $\sim 80-85\mu\text{m}$

M1: $\sim 60\mu\text{m}$

M2: $\sim 40\mu\text{m}$ (now being replaced)

Setting/other: $\sim 40\mu\text{m}$

Hexapod accuracy: $20\mu\text{m}$ pos., 2'' tip/tilt

KVN Telescopes

Korean VLBI Network (KVN) Telescopes

Three fully-steerable
telescopes in South Korea

Ulsan

Yonsei

Tamna

21m Az/El design (Antedo)

2-7mm



KVN Challenges

Surfaces and hexapod are conventional

Measurement by photogrammetry

Main design challenge is fast switching

High axis rates

High axis accelerations

Smooth motions to avoid vibration

Fast Switching Requirements

2.5° on sky

2.5s on source

2.5s on reference

Total cycle time, T

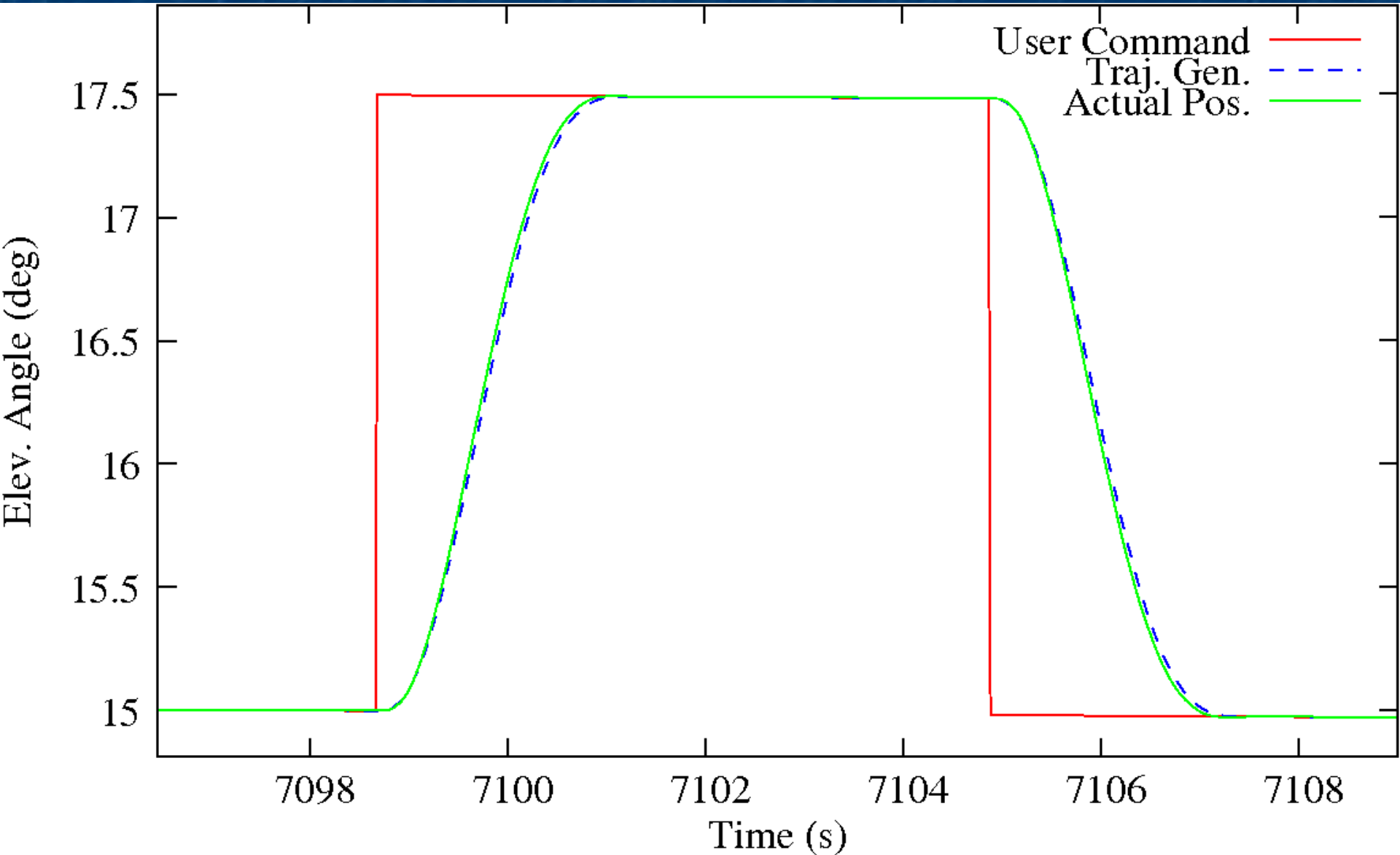
Elevation switching: $T \leq 13s$

Azimuth switching:

$EI \leq 60^\circ$, $T \leq 13s$

$60^\circ < EI \leq 80^\circ$, $T \leq 26s$

Fast Switching



KVN Main Axis Specifications

Maximum velocity (Az/EI): $3^\circ/\text{s}$

Maximum acceleration (Az/EI): $3^\circ/\text{s}^2$

Minimum smooth rate: 1 arcsec/s

On-sky Pointing accuracy: 4 arcsec RMS

Servo pointing goal: 2 arcsec RMS

LRF (Az/EI): 2.5 Hz

KVN Drive System

Fast slew/slow tracking

Brushless servo motors

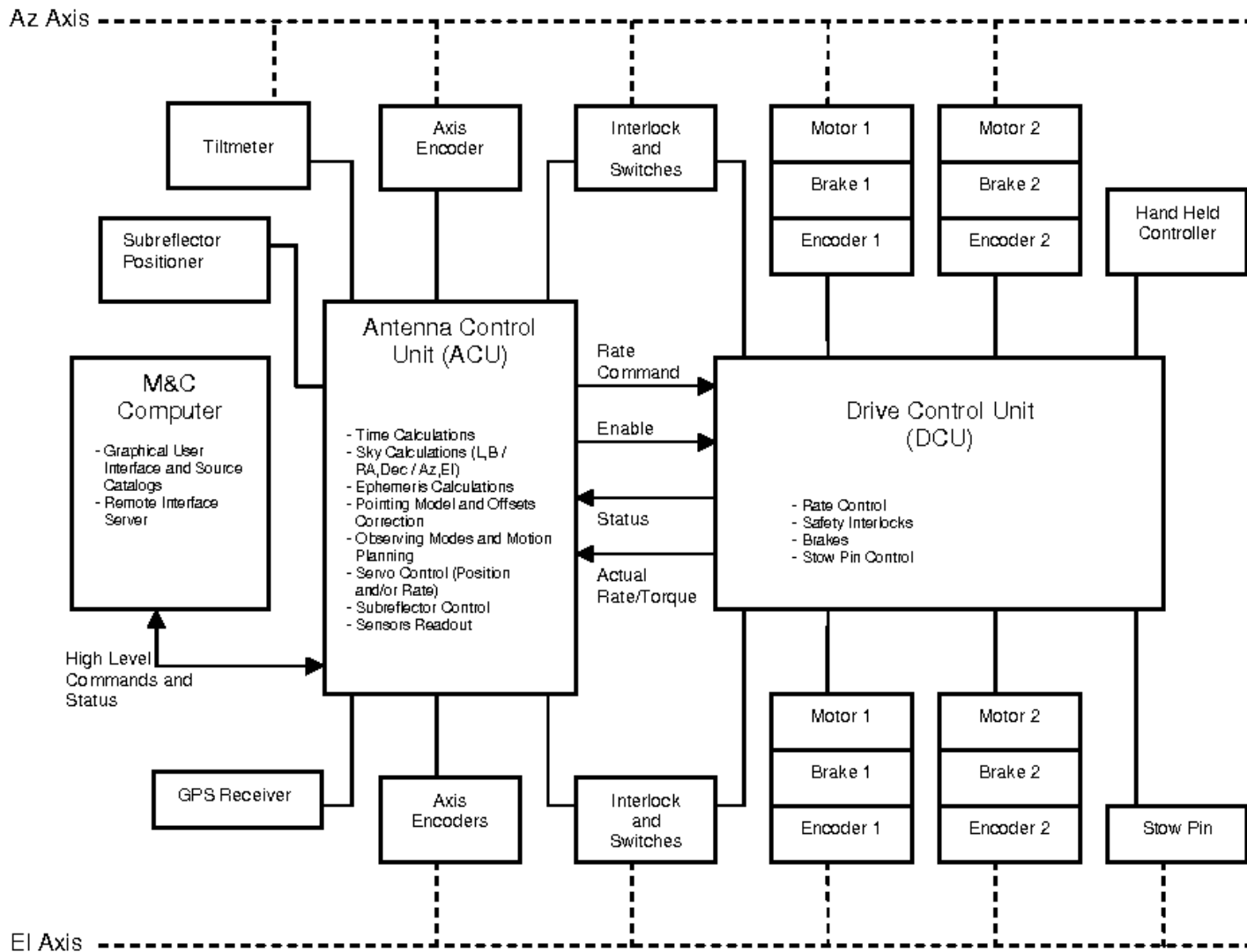
2 million counts/rev motor feedback

Manufacturer proved low speed operation

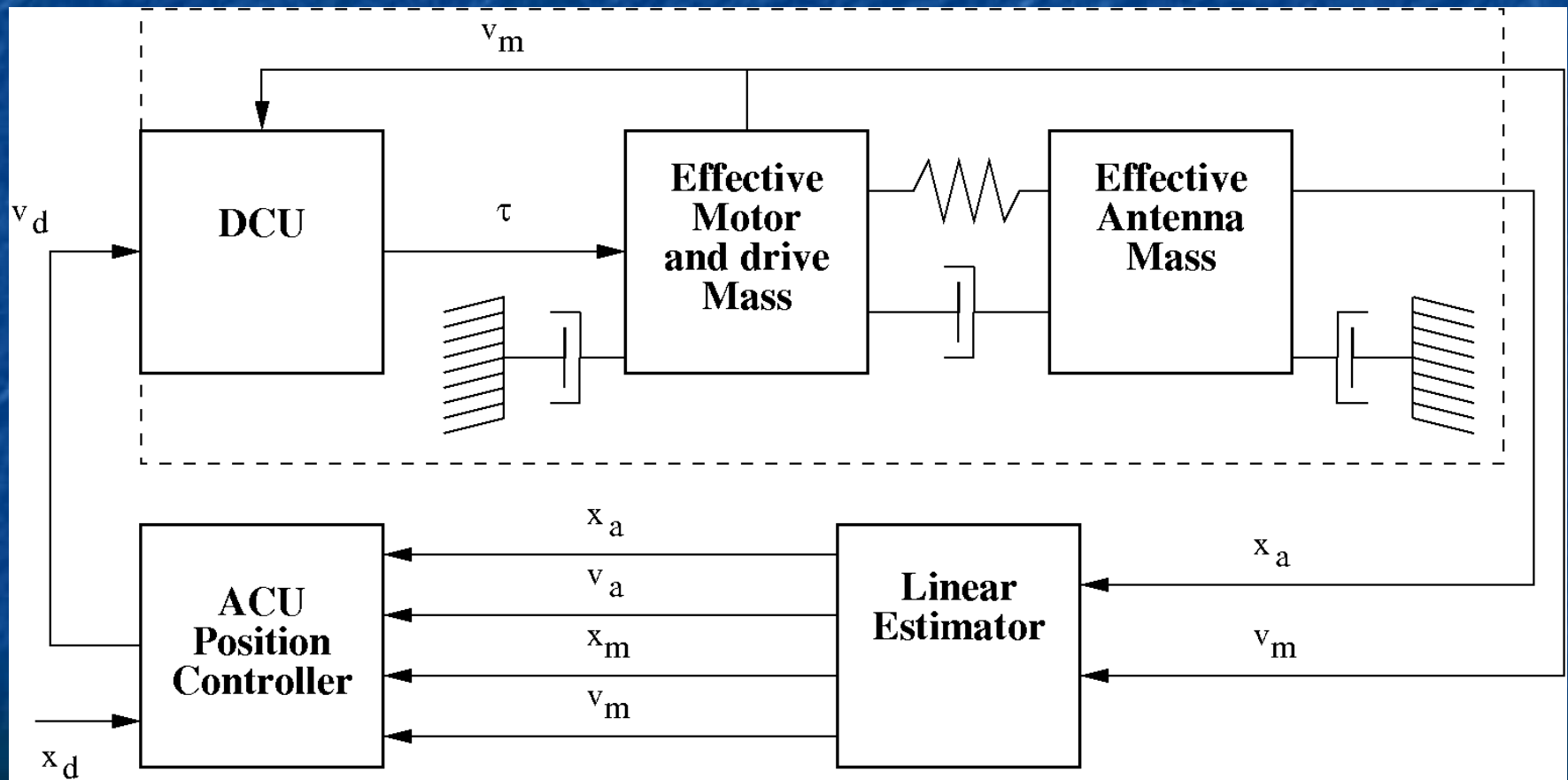
Amplifier issues

Rate loop performance superior to torque mode performance for this manufacturer

KVN System Architecture



KVN Controller Algorithm



KVN Controller

DCU Rate Loop is nominally PI... but

Better than when closed with PI in torque

Full-state optimal estimator to provide state vector

Model-based control on two-mass model for the position loop. Also includes:

Integrator

Trajectory generator

Feed forward control

KVN Performance

Usual metrics (estimates):

Tracking accuracy: $<0.5''$ RMS

Fast switching

Elevation: $<13s$

Azimuth: depends on elevation angle

KVN Fast Switching in Az

Az $+2.5^\circ$ on sky
($2.5^\circ/\cos(\text{El})$)

Large motions at high
elevation angles

Settling defined as
peak error <4 arcsec.

RMS <2 arcsec

| Elevation Angle | Cycle Time (s) |
|--------------------|-------------------|
| 15° | 12.6 |
| 30° | 12.0 |
| 50° | 12.2 |
| 70° | 13.7 |
| 80° | 18.0 |

Digital System Advantages

For both LMT and KVN, the digital system provides excellent performance in main axis control

Architecture means we have access:

Torque, Rate, Position, Trajectory, Astro.

LMT also has an active surface, improving the performance of the primary

It's OK to be jealous...

Digital System Advantages

What do we get with the digital system?

Usual (but improved):

- Astronomical tracking

- Position switching

- Pointing offsets

- Instrument holds

- Raster scanning

New capabilities also emerge

Digital System Capabilities

Easier measurements

- Balance

- Controller frequency response

New observing modes

New surface corrections

Design paradigm

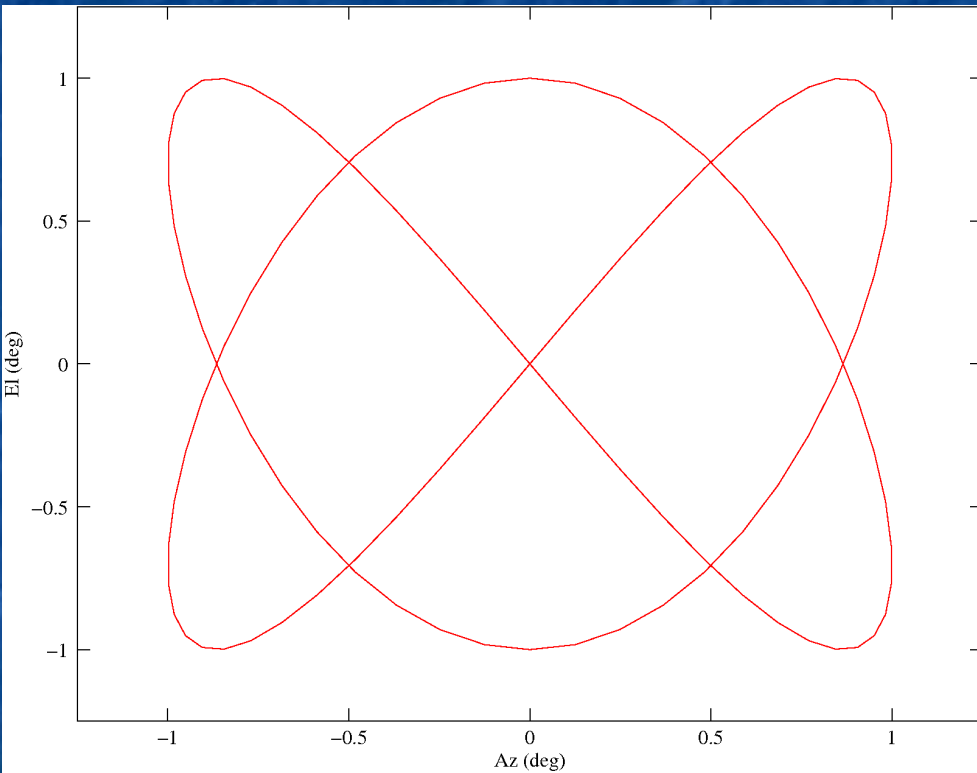
Remember: the goal is time on sky

New Observing Modes

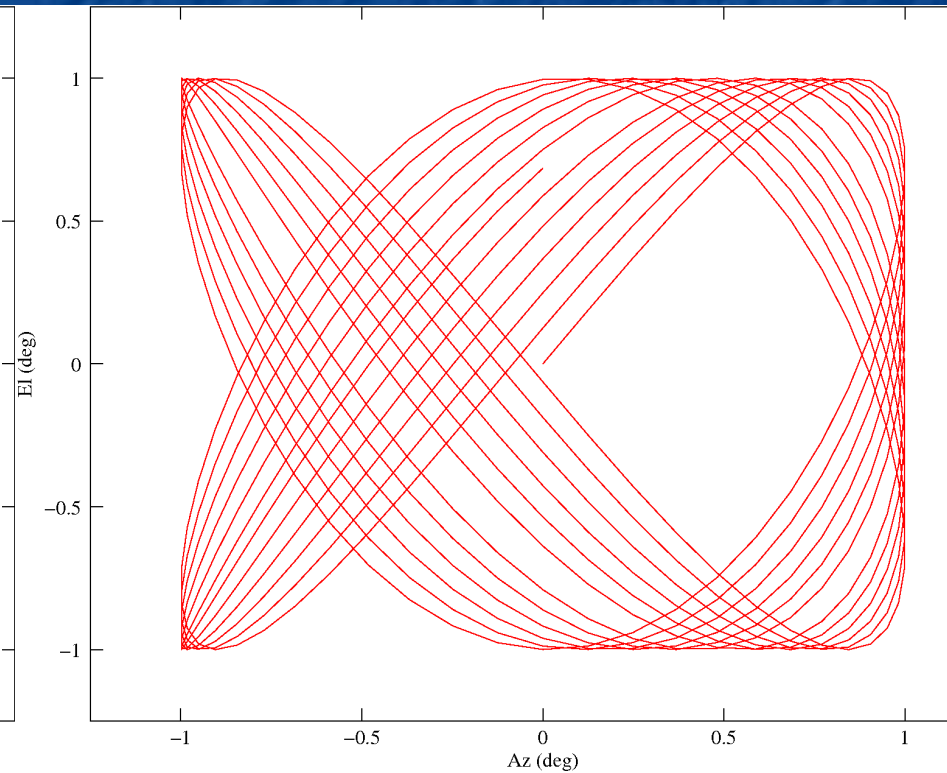
Time parametric scan patterns

Position parametric scan patterns

Time Parametric Scans

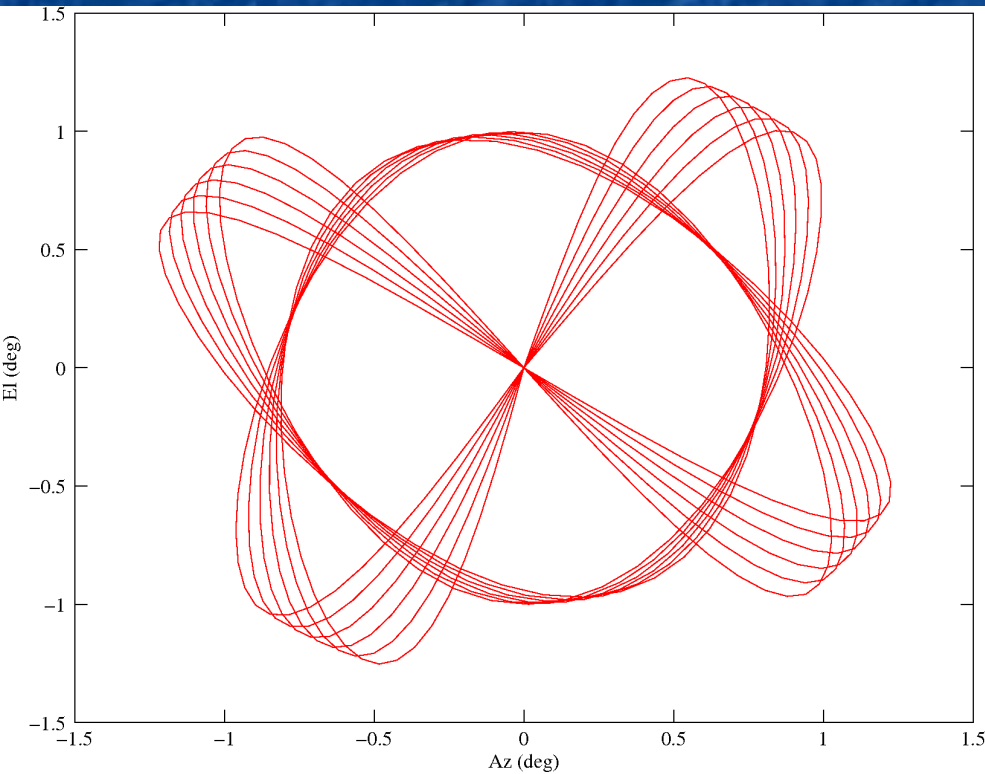


2:3 Lissajous

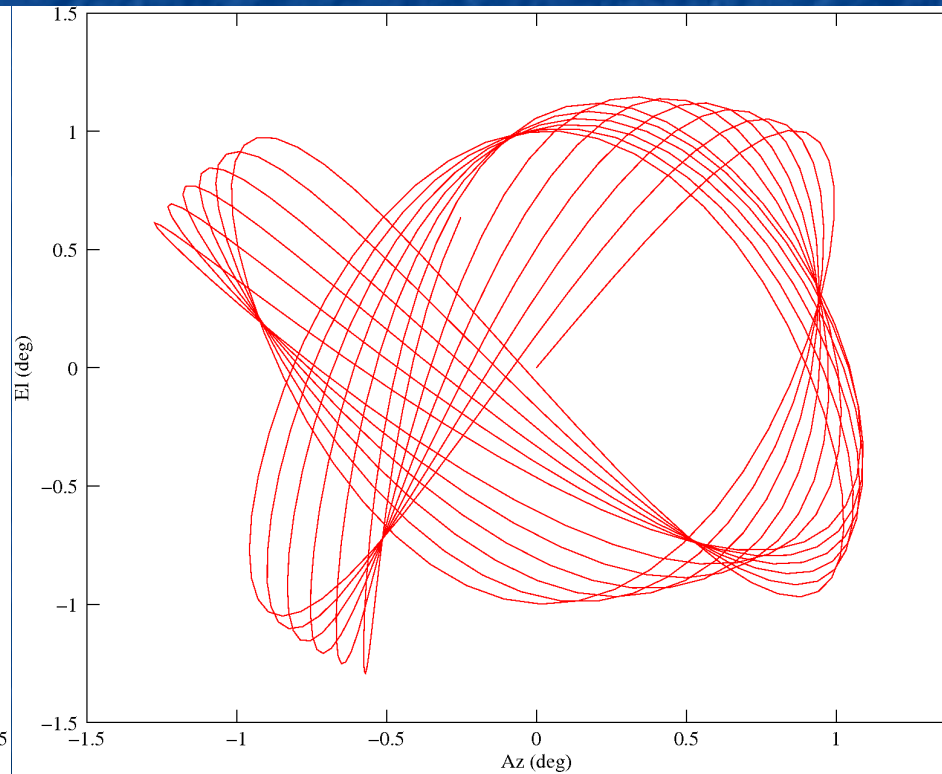


2:(3.02) Lissajous

Rotating Time Parametric Scans



2:3 Lissajous



2:(3.02) Lissajous

Position Parametric Scans

For Holography maps

Frequent return to center of the map

Improved approach

Make radial cuts through the center

Rotate to a new starting point and repeat

Problem: Non-uniform time coverage

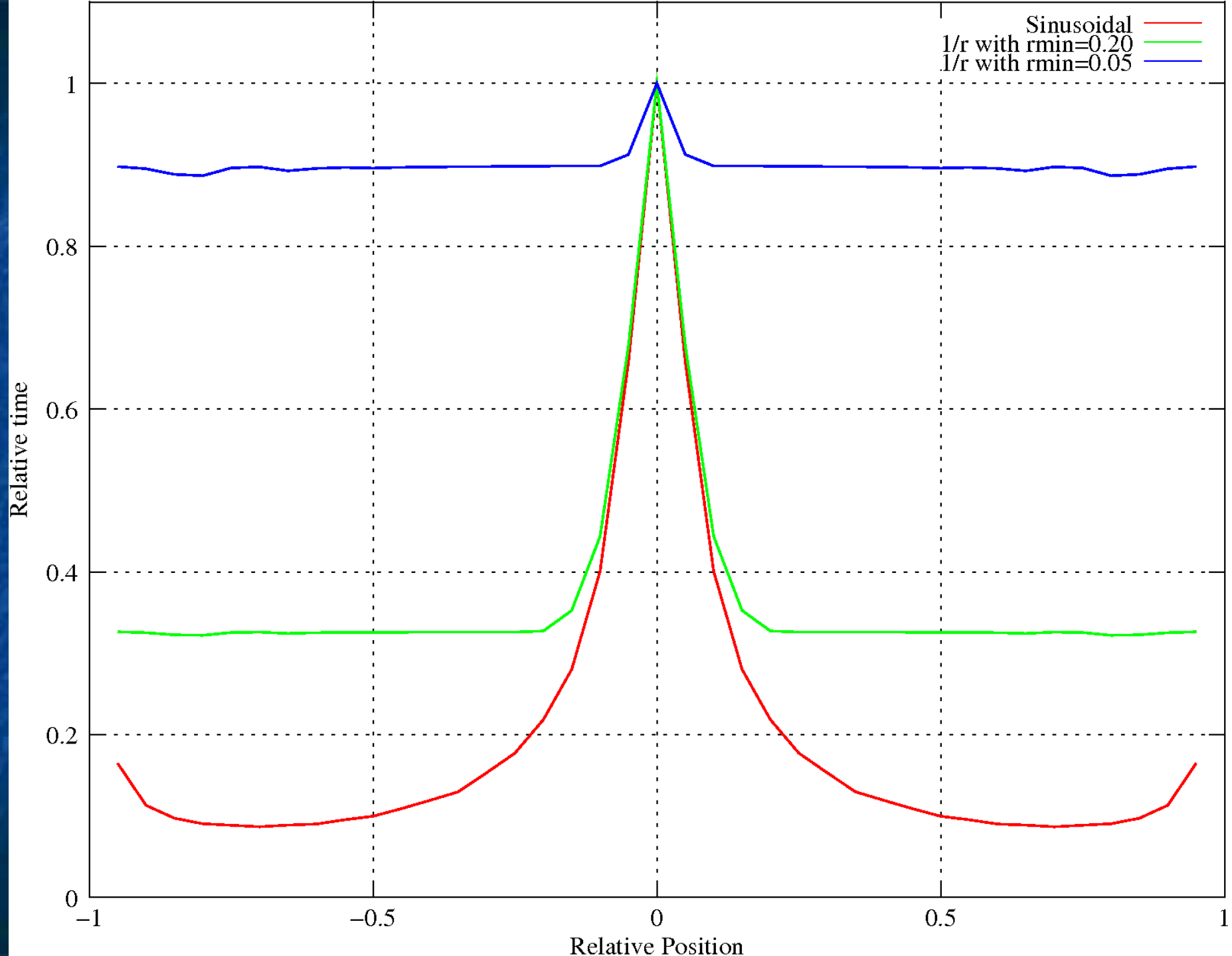
Radial Map Definition

Parametric Scan Map, $1/r$ velocity profile

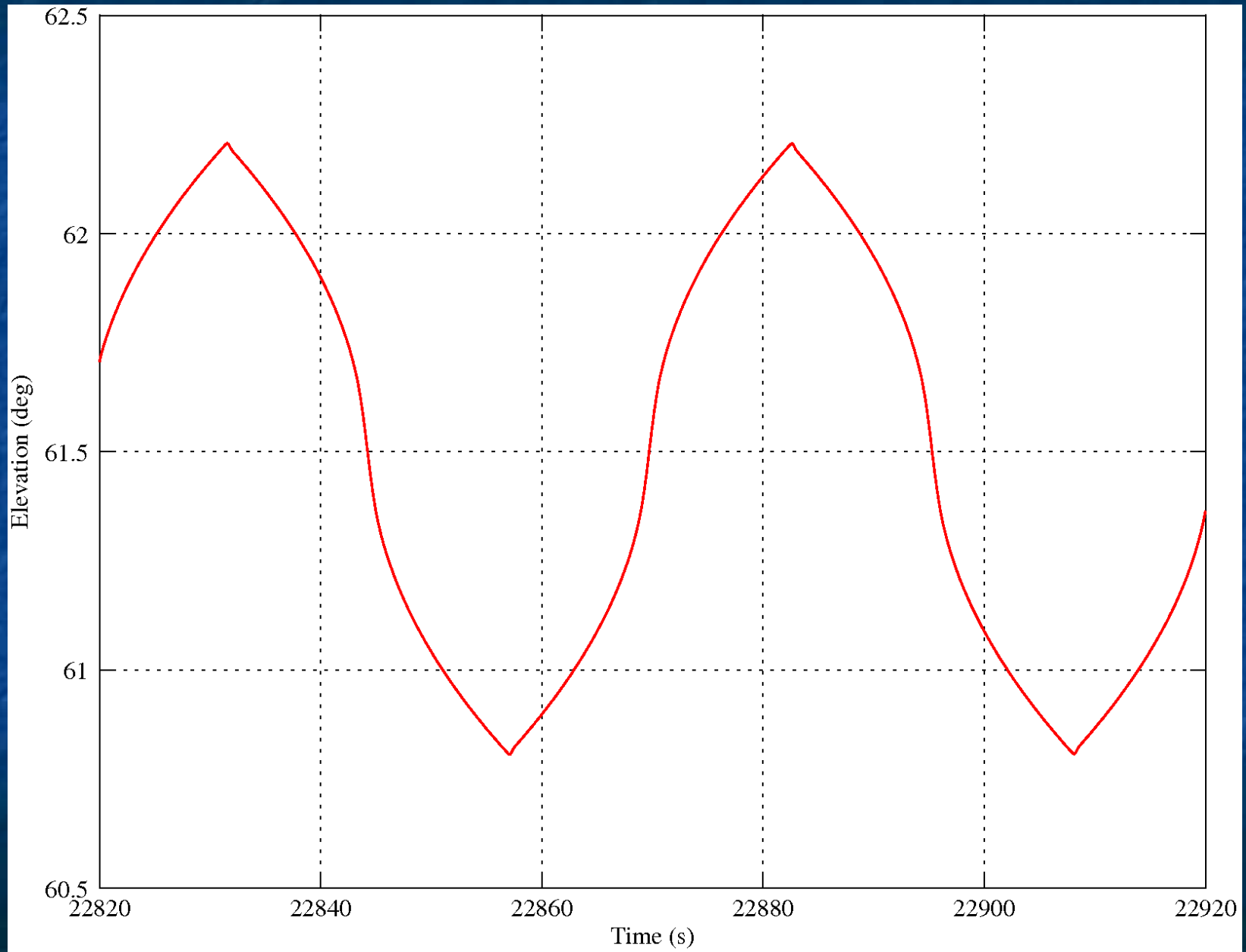
High velocity near map center

Smooth the transition to limited region

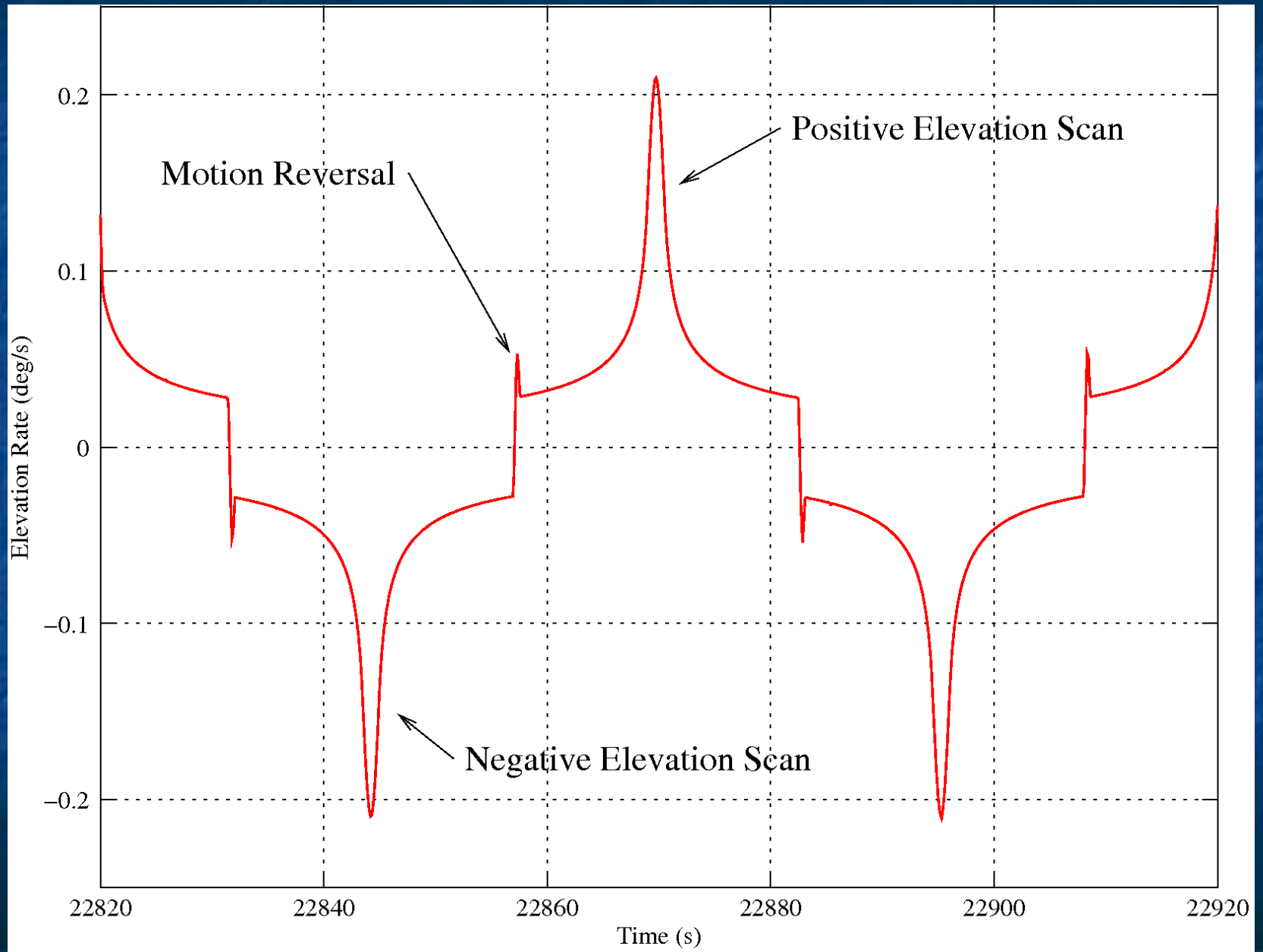
Time coverage is increasingly uniform as the limited region is smaller



Typical Results (Position)



Typical Results (Velocity)



New Surface Corrections

Standard practice: Check pointing/focus

OOF to measure astigmatic effects

Additional Zernike's can also be measured

Time trade-off

Astigmatism is most important

Immediate active surface correction

New Design Paradigm

Fully Digital Main Axis Control

Improves with higher LRF

Active surface

Can remove gravity effects

Dominant residual effects

Pointing, focus, astigmatism, trefoil

Homology is not only not necessary, it can reduce performance for active telescopes

Questions?