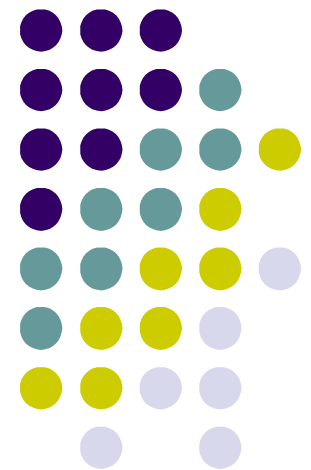


High Precision Calibration of Wide Bandwidth Observations with the GBT

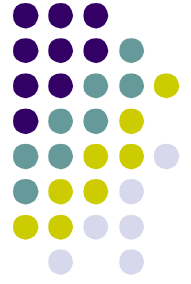


Ronald J Maddalena
NRAO, Green Bank
July 26 2007



Shelly Hynes (Louisiana School for Math, Science and the Arts)
Charles Figura (Wartburg College)
Chelen Johnson (Breck School)
NRAO-GB Scientific and Engineering staff

Typical Position-Switched Calibration Equation



$$S(\nu) = \left(\frac{2k}{\eta_A(\nu, Elev) \cdot Area_p} \right) \cdot T_A(\nu) \cdot e^{\tau(\nu, t) \cdot A(Elev, t)}$$

$$T_A(\nu) = \left(\frac{Sig(\nu) - Ref(\nu)}{Ref(\nu)} \right) \cdot T_{Sys}^{Ref}$$

$$T_{Sys}^{Ref} = \left\langle \frac{Ref(\nu)}{Ref_{On}(\nu) - Ref_{Off}(\nu)} T_{Cal}(\nu) \right\rangle_{BW}$$

$A(Elev, t)$ = Air Mass

$\tau(\nu, t)$ = Atmospheric Zenith Opacity

T_{cal} = Noise Diode Temperature

Area = Physical area of the telescope

$\eta_A(\nu, Elev)$ = Aperture efficiency (point sources)

$T_A(\nu)$ = Source Antenna Temperature

$S(\nu)$ = Source Flux Density

$Sig(\nu)$, $Ref(\nu)$ = Data taken on source and on blank sky (in units backend counts)

On, Off = Data taken with the noise diode on and off

T_{sys} = System Temperature averaged over bandwidth

Position-Switched Calibration Equation



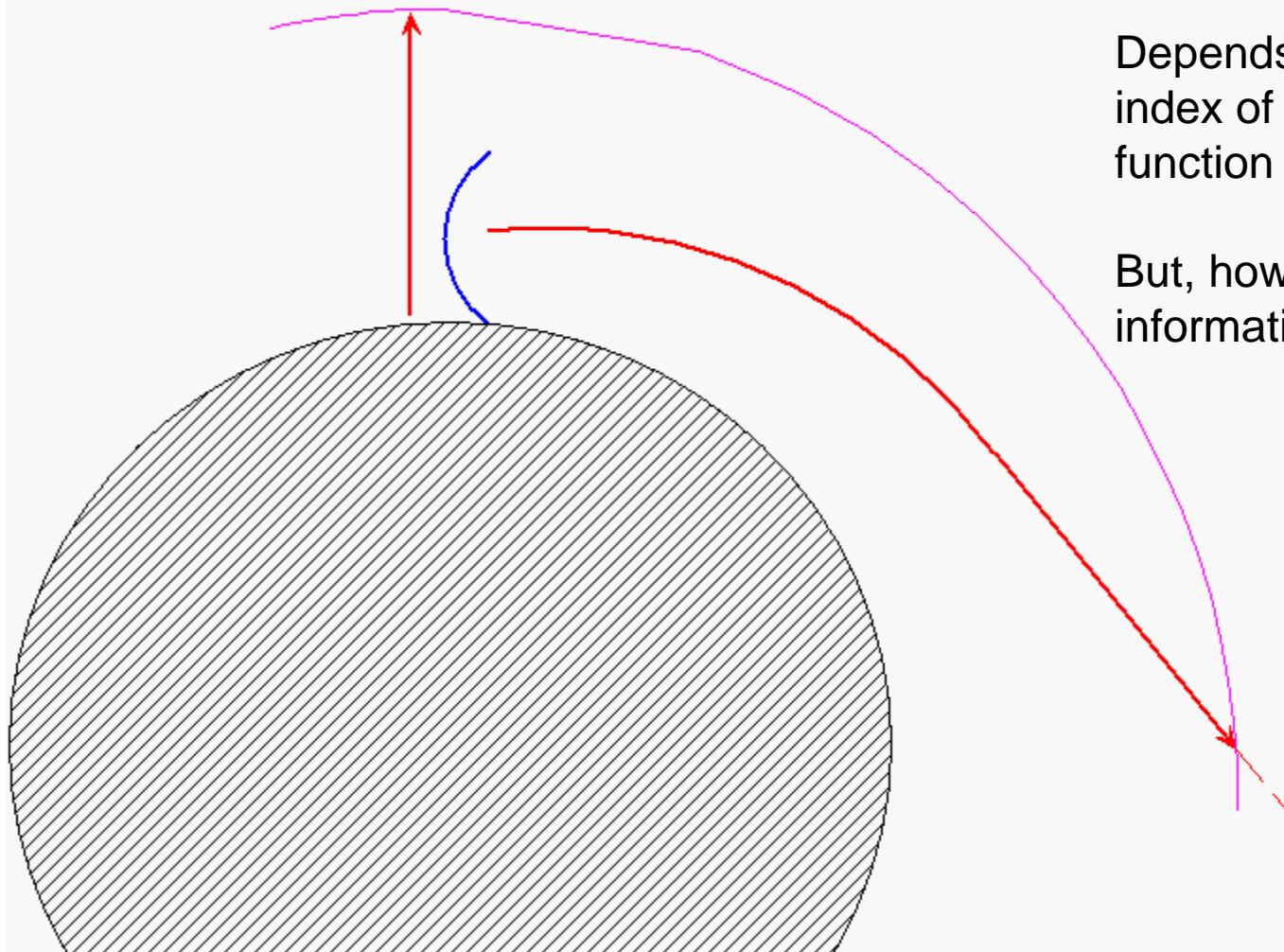
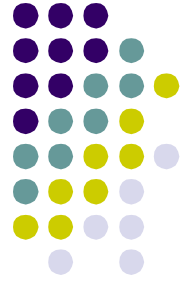
$$S(\nu) = \left(\frac{2k}{\eta_A(\nu, Elev) \cdot Area_p} \right) \cdot \left(\frac{Sig(\nu) - Ref(\nu)}{Ref(\nu)} \right) \left\langle \frac{Ref(\nu)}{Ref_{on}(\nu) - Ref_{off}(\nu)} T_{cal}(\nu) \right\rangle \cdot e^{\tau(\nu) \cdot A(Elev)}$$

Sources of uncertainties

$$\left(\frac{\Delta S}{S} \right)^2 = (\tau \cdot \Delta A)^2 + (A \cdot \Delta \tau)^2 + \left(\frac{\Delta T_{cal}}{T_{cal}} \right)^2 + \left(\frac{\Delta \eta}{\eta} \right)^2$$

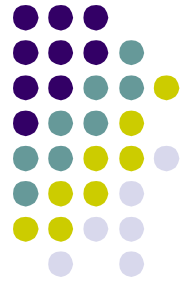
- 10-15% accuracy have been the 'standard'
- Usually, errors in T_{cal} dominate
- Goal: To achieve 5% calibration accuracy without a significant observing overhead.

Air Mass Estimates



Depends upon density and index of refraction as a function of height

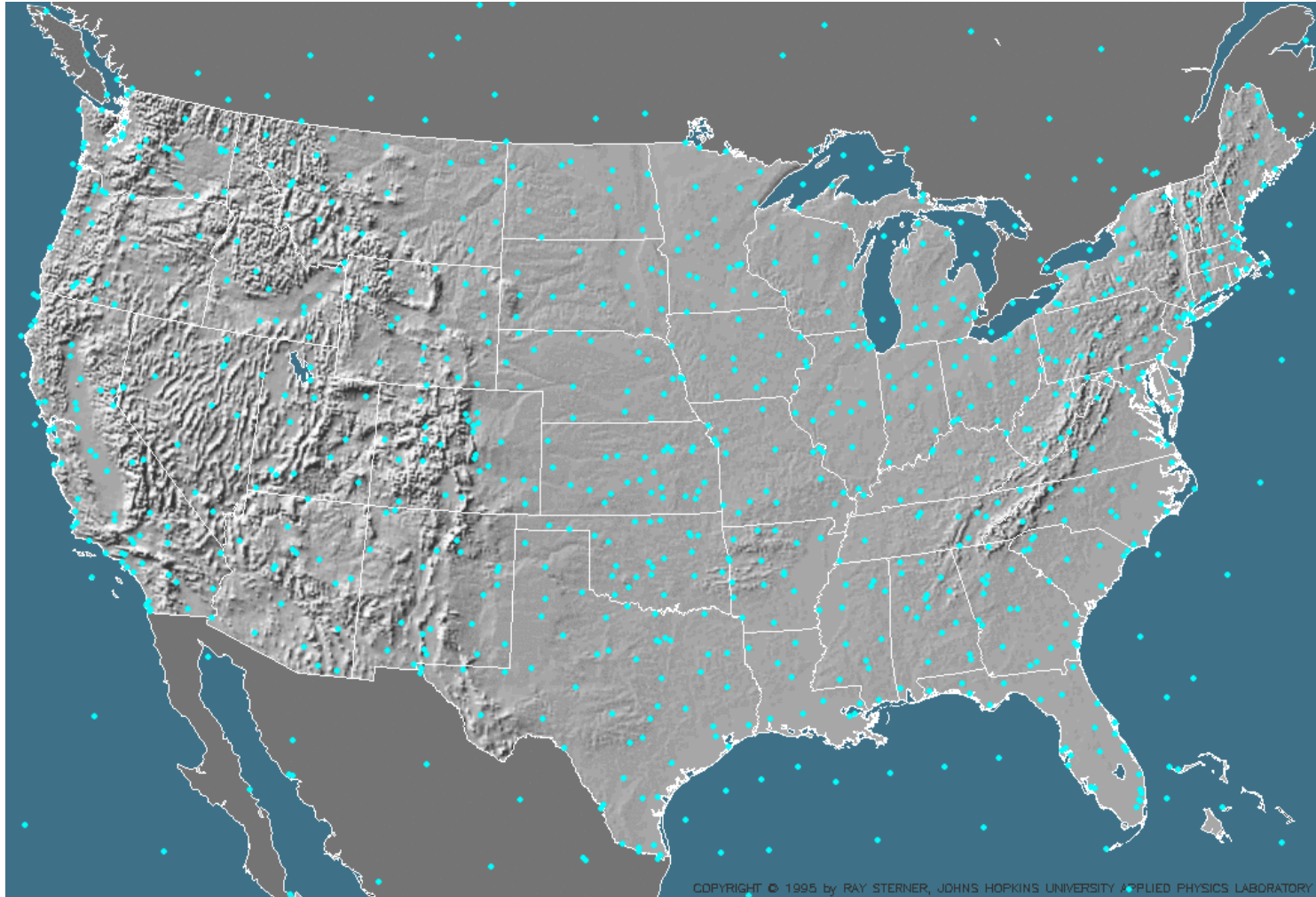
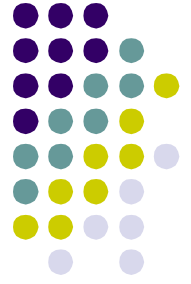
But, how can one get this information?



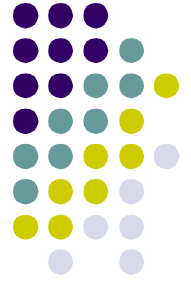
Vertical Weather Data

- Provided by the national weather services via FTP
- 60 hr forecasts (ETA model), updated every 12 hrs
- For each hour, provides as a function of ***height above the ground***
 - Temperature, Pressure, Dew Point, Cloud Cover,
- ~40 heights that extend well into the stratosphere
- One can derive ***as a function of height***:
 - Density
 - Index of refraction
 - Absorption coefficient (dry air, water – continuum & line, oxygen line, hydrosols) (Liebe model)

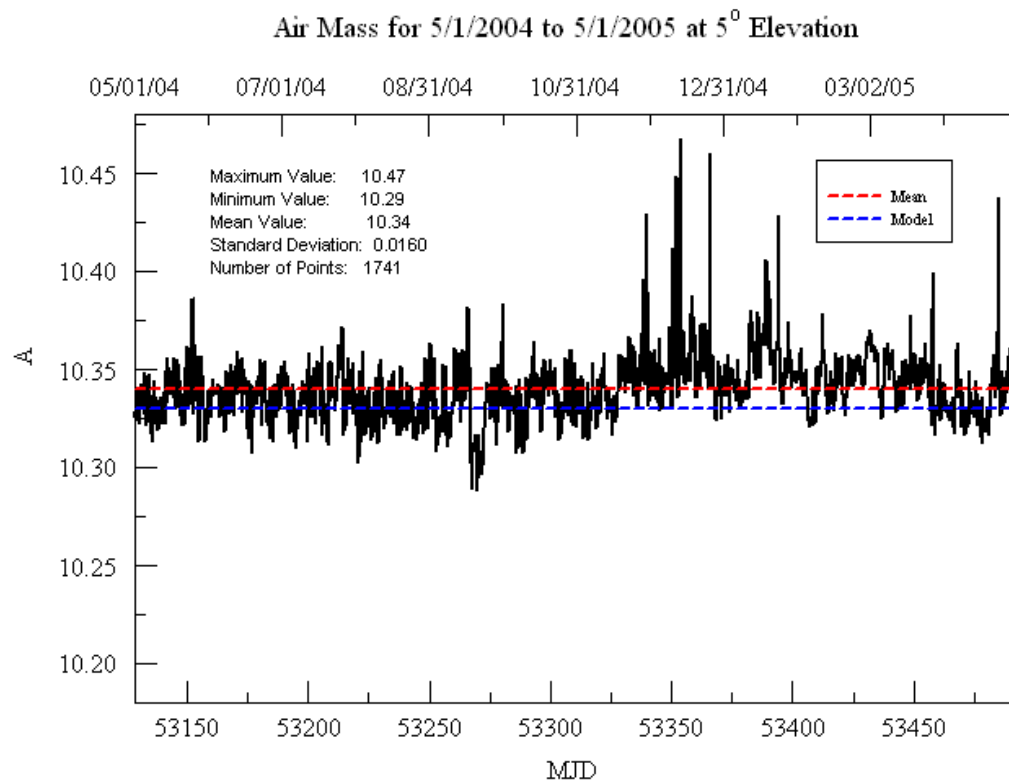
Vertical Weather Data



Air Mass Estimates



- Air Mass derived from the vertical values of density & index of refraction.
- For 1% calibration error, require A to ~ 0.1

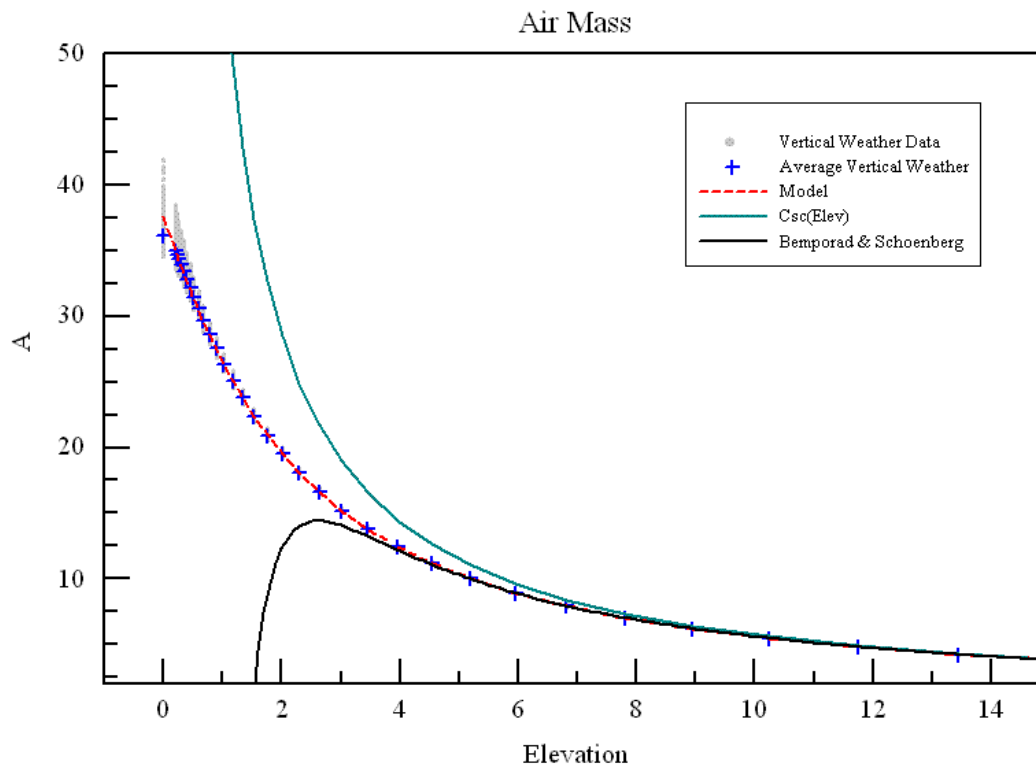


- Probably should use weather dependent Air Mass for elevations below 5 deg
- Probably can ignore weather dependency above 5 deg.



Air Mass Estimate

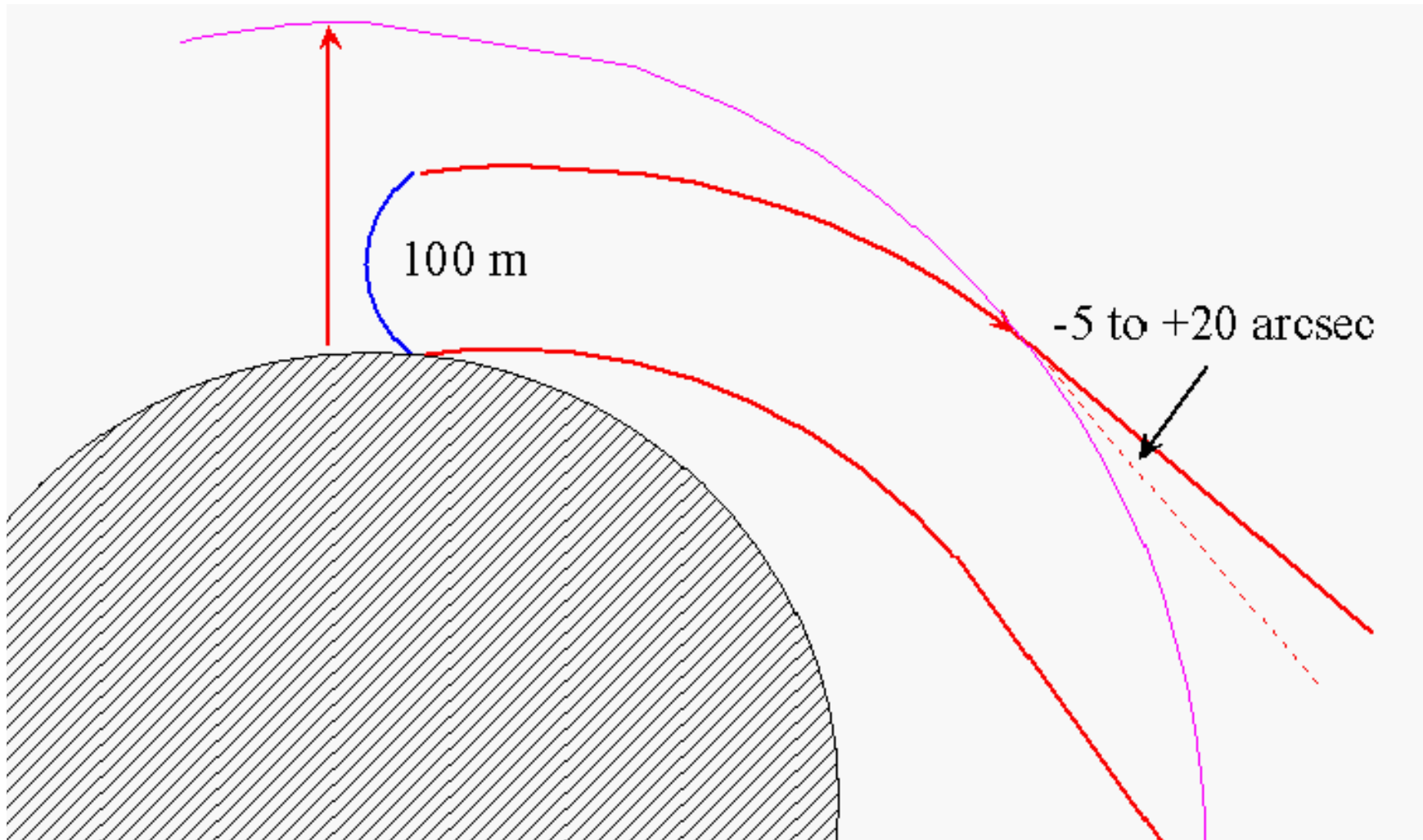
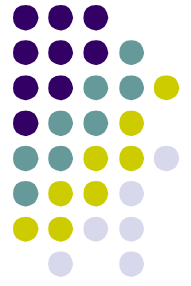
- Air Mass traditionally modeled as $1/\sin(\text{Elev})$
- For 1% calibration accuracy, must use a better model below 15 deg.



$$A = -0.0234 + \frac{1.014}{\sin\left(\text{Elev} + \frac{5.18}{\text{Elev} + 3.35}\right)}$$

- Good to 1 deg
- Use $1/\sin(\text{Elev})$ above 60 deg
- Coefficients are site specific, at some low level

Air Mass Estimates



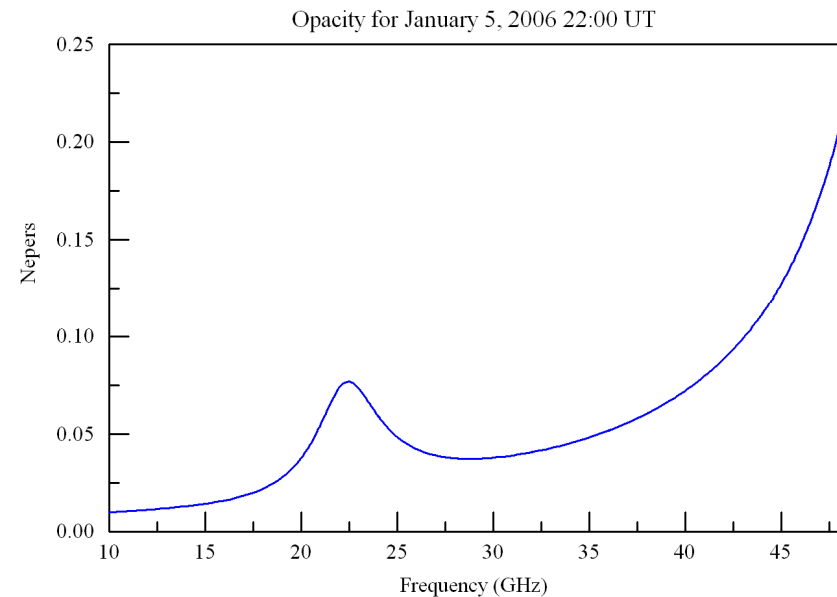
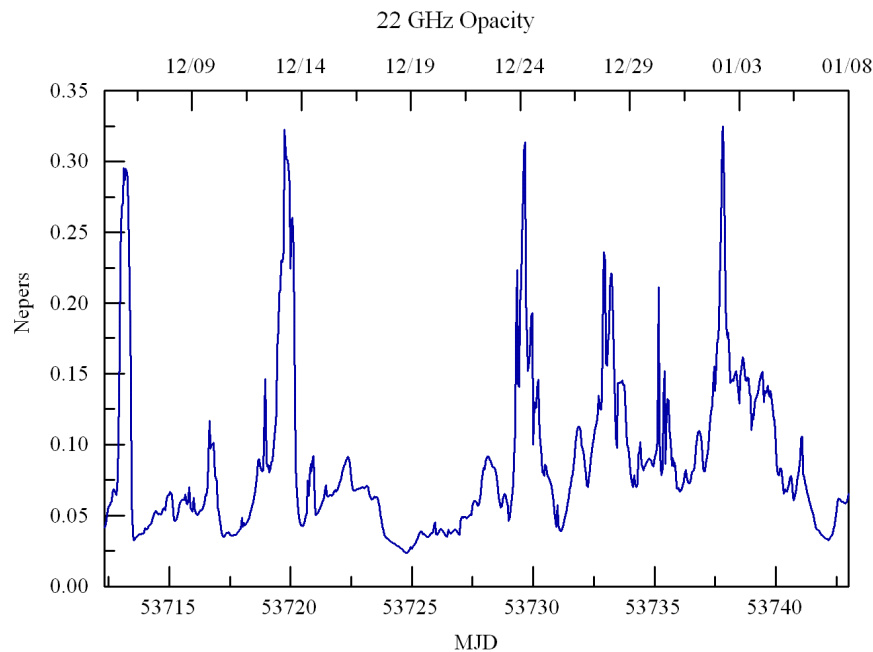
Opacity Estimates

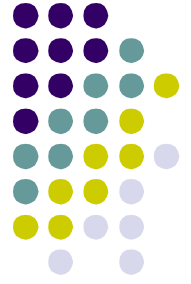


- Vertical weather data provides absorption as a function of height

$$\tau(\nu, t) = \int_0^{\infty} \left(\kappa_{Dry}(\nu, t) + \kappa_{O_2}(\nu, t) + \kappa_{Water_cont}(\nu, t) + \kappa_{water_line}(\nu, t) + \kappa_{hydrosols}(\nu, t) \right) dH$$

$$T_{Sys}(Elev, \nu, t) \cong T_{rcvr}(\nu) + T_{spill}(Elev) + T_{cmb} e^{-\tau(\nu, t) \cdot A(Elev)} + \int_0^{\tau} T(H, t) \cdot e^{\tau(h, \nu, t)} d\tau$$



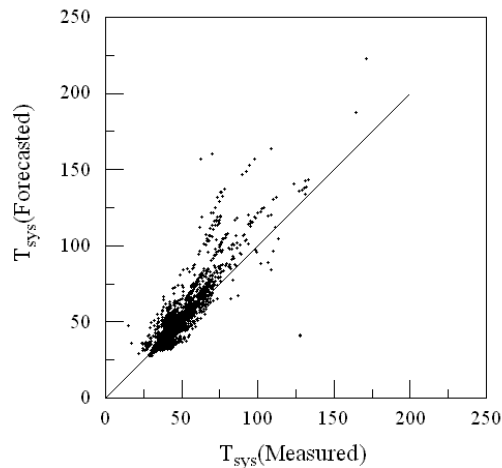
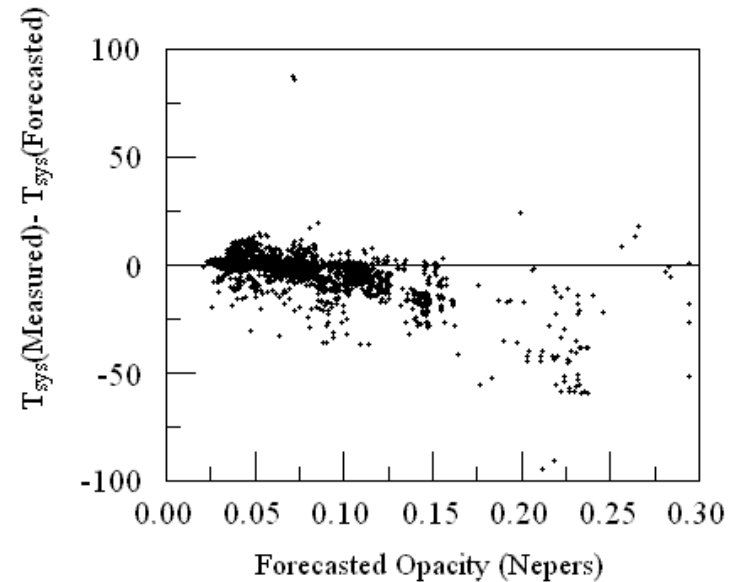
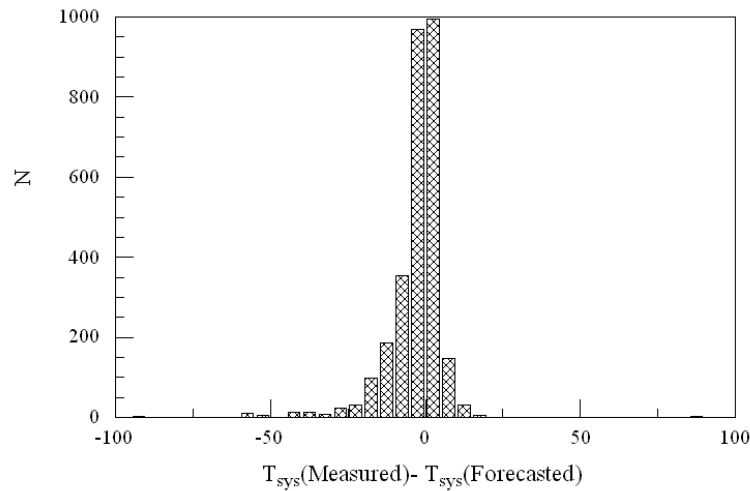
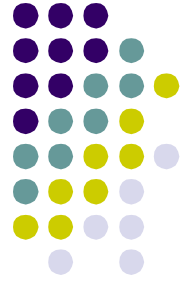


Opacity Estimates

- Are derived opacities accurate? Comparisons using tipping radiometers have difficulties
 - Must do multiple tips for wideband observations
 - Tips take up telescope time
 - Requires knowing T_{cal} to high accuracy, which requires knowing τ .
 - Some dedicated tippers do not provide enough information to estimate τ near the 22 GHz water line
 - Requires a representative T_{Atm} that is good to ~ 5 K

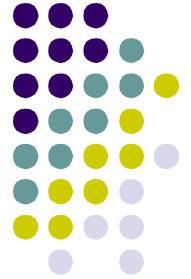
$$T_{\text{Atm}} \approx \frac{\int \kappa(H) \cdot T(H) \cdot dH}{\int \kappa(H) \cdot dH}$$

Comparison of measured and estimated 22 GHz Tsys

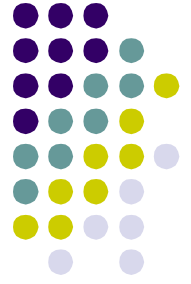


- For 1% calibration accuracy, requires τ to 0.01
- Implies forecasted and actual Tsys should be within 3 K
- Current model sufficient at $\tau < 0.1$
- Overestimates contribution from hydrosols. Not unexpected.

Noise Diode Estimates



- Traditionally used hot-cold load measurements
 - Provide ~10% accuracy
 - Frequency resolution sometimes wider than frequency structure in T_{rcvr} or T_{cal}
 - Time consuming
 - Systematics/Difficulties
 - Loads must be opaque
 - Frost forming on LN₂ loads
 - Linearity ($T_{Hot} \gg T_{Cold}$)
 - Observers can't do their own Hot-Cold tests



Noise Diode Estimates

- Instead, we recommend an On-Off observation
 - Use a point source with known flux -- polarization should be low or understood
 - Use the same exact hardware, exact setup as your observation. (i.e., don't use your continuum pointing data to calibrate your line observations.)
 - Observations take ~5 minutes per observing run
 - Staff take about 2 hrs to measure the complete band of a high-frequency, multi-beam receiver.
 - Resolution sufficient: 1 MHz, sometimes better
 - Accuracy of ~ 1%, mostly systematics.



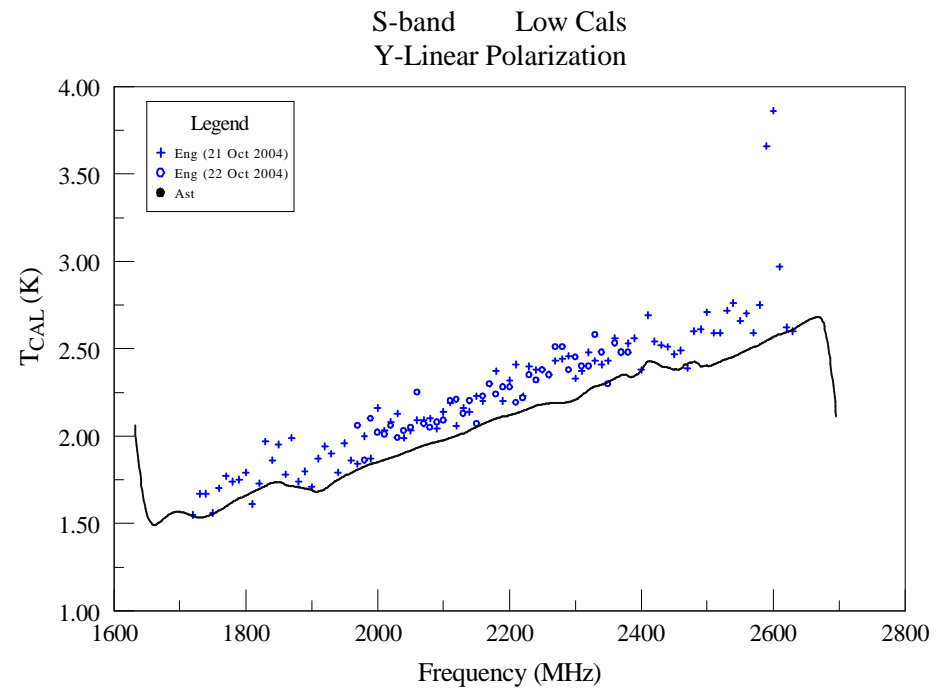
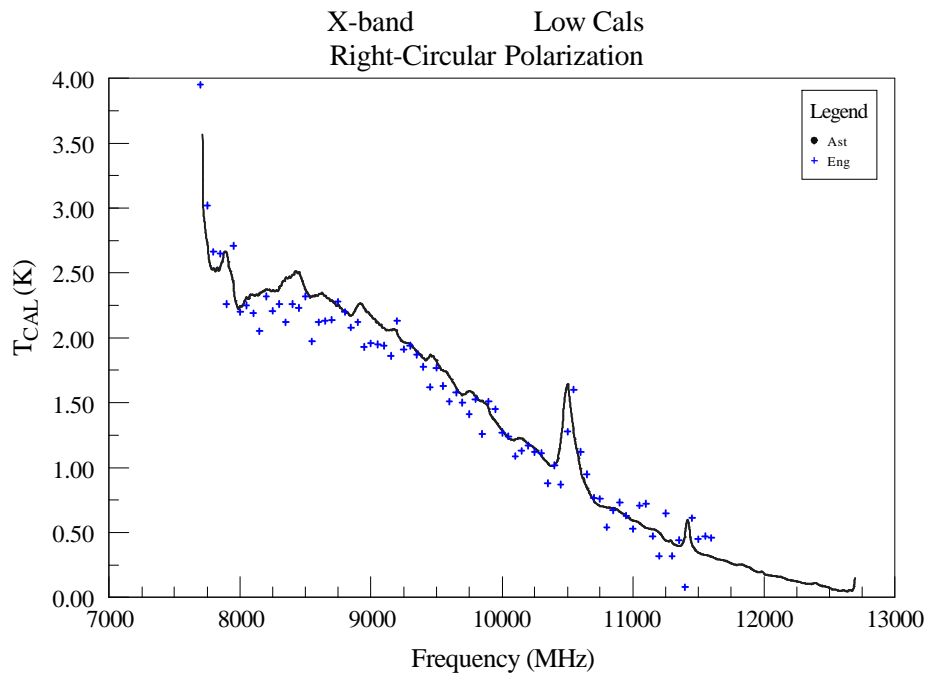
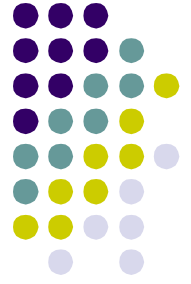
Noise Diode Estimates

$$S(\nu) = \left(\frac{2k}{\eta_A(\nu, Elev) \cdot A_p} \right) \cdot \left(\frac{Sig(\nu) - Ref(\nu)}{Ref(\nu)} \right) \left\langle \frac{Ref(\nu)}{Ref_{on}(\nu) - Ref_{off}(\nu)} T_{Cal}(\nu) \right\rangle \cdot e^{\tau(\nu) \cdot A}$$

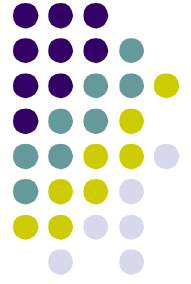
Remove Averaging, Solve for Tcal

$$T_{Cal}(\nu) = \frac{\eta_A(\nu, Elev) \cdot Area_p}{2k \cdot e^{\tau(\nu) \cdot A}} \left(\frac{Ref_{on}(\nu) - Ref_{off}(\nu)}{Sig(\nu) - Ref(\nu)} \right) \cdot S(\nu)$$

Noise Diode Estimates



Position-Switched Calibration Equation



$$S(\nu) = \left(\frac{2k}{\eta_A(\nu, Elev) \cdot Area_p} \right) \cdot \left(\frac{Sig(\nu) - Ref(\nu)}{Ref(\nu)} \right) \left\langle \frac{Ref(\nu)}{Ref_{on}(\nu) - Ref_{off}(\nu)} T_{cal}(\nu) \right\rangle \cdot e^{\tau(\nu) \cdot A(Elev)}$$

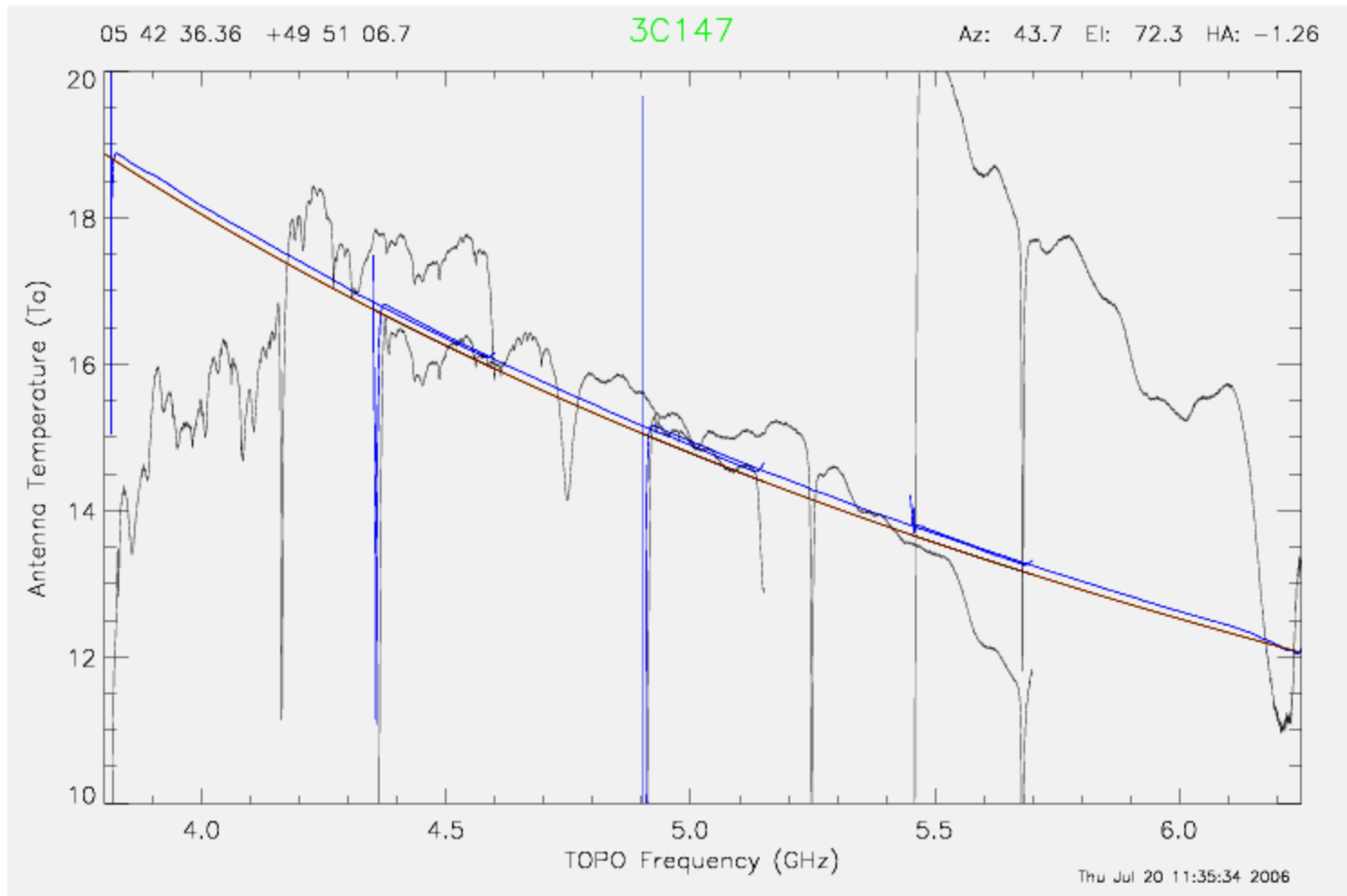
Baseline structure

$$Baselines(\nu) = \frac{T_A + T_{atm}^{Sig} - T_{atm}^{Ref}}{\langle T_{sys} \rangle_{BW}} \cdot \left\{ \left\langle T_{rcvr} + \frac{T_{cal}}{2} \right\rangle_{BW} - T_{rcvr}(\nu) - \frac{T_{cal}(\nu)}{2} \right\}$$

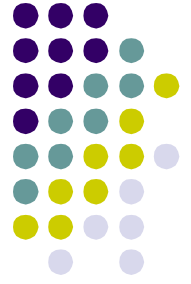
Assumption of linearity

$$S \propto Sig - Ref$$

Baseline Structure



Baseline Shapes



$$S(\nu) = \left(\frac{2k}{\eta_A(\nu, Elev) \cdot A_p} \right) \cdot \left(\frac{Sig(\nu) - Ref(\nu)}{Ref(\nu)} \right) \left\langle \frac{Ref(\nu)}{Ref_{on}(\nu) - Ref_{off}(\nu)} T_{Cal}(\nu) \right\rangle \cdot e^{\tau(\nu) \cdot A}$$

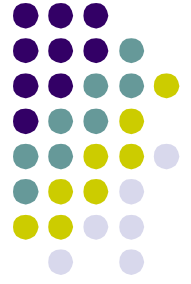
Remove Averaging – Vector Calibration

$$S(\nu) = \left(\frac{2k}{\eta_A(\nu, Elev) \cdot A_p} \right) \cdot \left(\frac{Sig(\nu) - Ref(\nu)}{Ref_{on}(\nu) - Ref_{off}(\nu)} \right) \cdot T_{Cal}(\nu) \cdot e^{\tau(\nu) \cdot A}$$

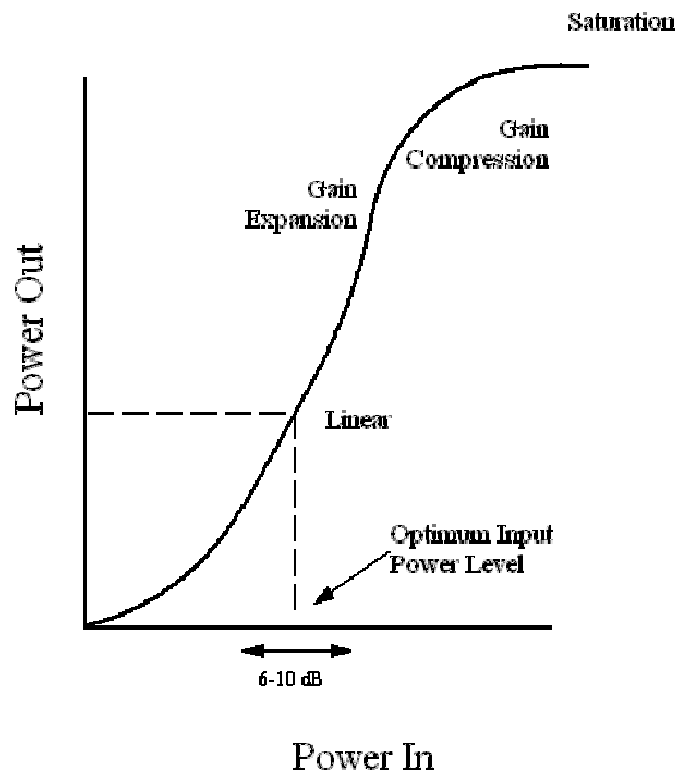
- Traditional equation OK for narrow bandwidth observations
- Traditional provides good calibration only at band center
- Vector algorithm provides good calibration across wide bandwidths
- Vector algorithm is substantially noisier when $T_A \neq 0$

$$\sigma^2 \approx \frac{1}{BW \cdot t} \left(T_A^2 + T_{Sys}^2 + \frac{T_{Sys}^2 T_A^2}{T_{Cal}^2} \right)$$

- Smooth Ref_{on} - Ref_{off} – use Savitzky-Golay smoothing

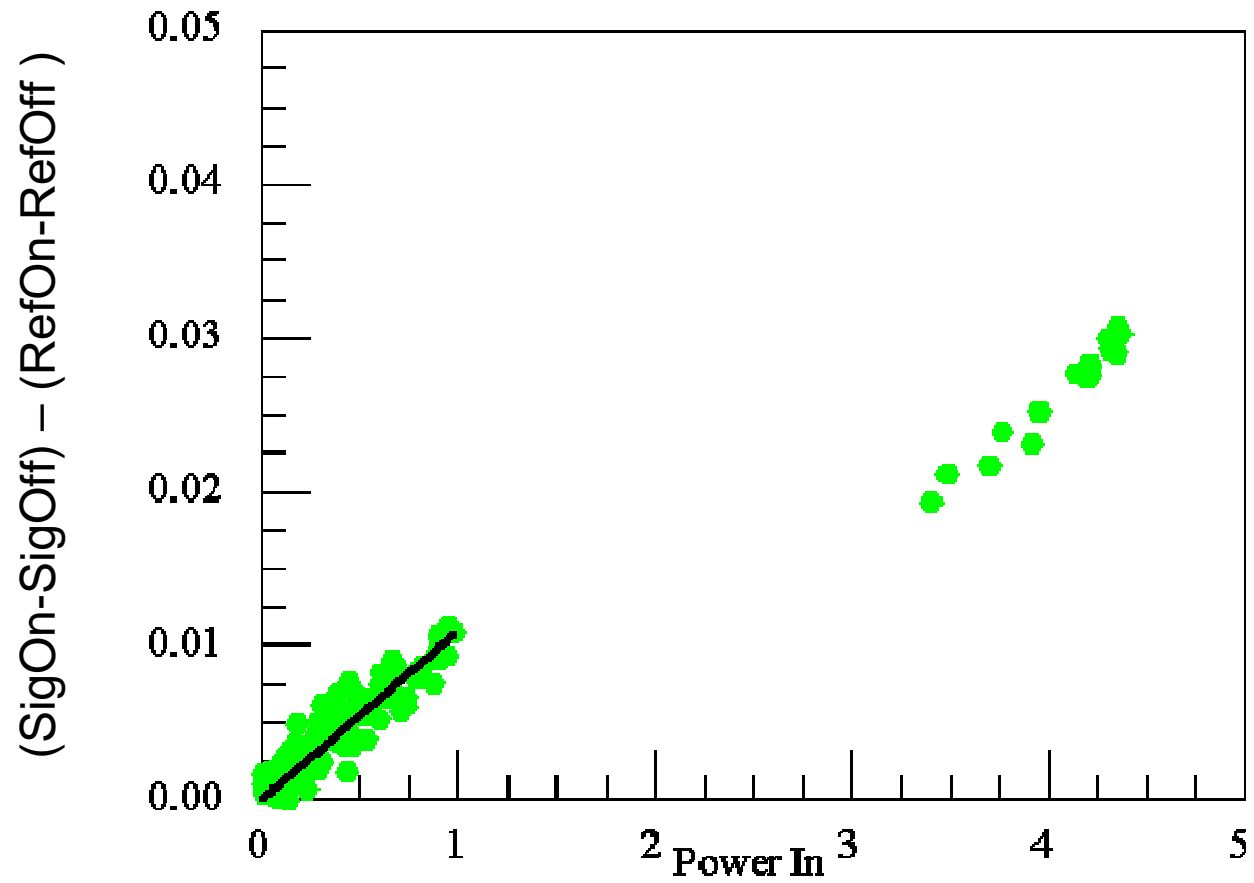
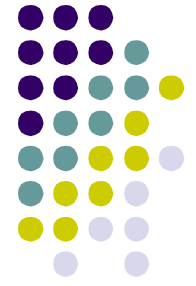


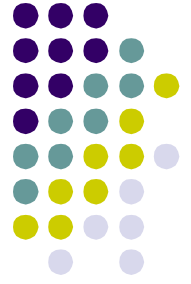
Non-linearity



- If system is linear,
 - $P_{out} = B * P_{in}$
 - $(Sig_{On} - Sig_{Off}) - (Ref_{On} - Ref_{Off}) = 0$
- Model the response curve to 2nd order:
 - $P_{out} = B * P_{in} + C * P_{in}^2$
- Our 'On-Off' observations of a calibrator provide:
 - Four measured quantities: Ref_{off} , Ref_{on} , Sig_{off} , Sig_{on}
 - T_A From catalog
 - Four desired quantities: B , C , T_{cal} , T_{sys}
- It's easy to show that:
 - $C = [(Sig_{on} - Sig_{off}) - (Ref_{on} - Ref_{off})] / (2T_A T_{cal})$
- Thus:
 - Can determine if system is sufficiently linear
 - Can correct to 2nd order if it is not

Non-linearity

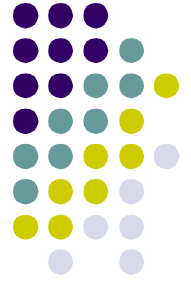




Summary

- To obtain few percent calibration accuracy
 - Wide bandwidths require frequency dependent opacities, efficiencies, T_{sys} , and T_{cal} .
 - New weather-independent model for air mass is usually sufficient
 - Opacities from vertical, forecasted weather data sufficient for medium to low-opacity conditions
 - Simple observation of calibrator provides high accuracy, high frequency resolution T_{cal} .
 - Assumptions of traditional on-off calibration algorithm introduces baseline shapes for wide bandwidth observations. Should use 'vector' algorithms.
 - 2nd order non-linearity -- measurable (by-product of T_{cal} observation) and correctable. Might be significant.

Typical Position-Switched Calibration Equation



$$S(\nu) = \left(\frac{2k}{\eta_A(\nu, Elev) \cdot Area_p} \right) \cdot T_A(\nu) \cdot e^{\tau(\nu, t) \cdot A(Elev, t)}$$

$$T_A(\nu) = \left(\frac{Sig(\nu) - Ref(\nu)}{Ref(\nu)} \right) \cdot T_{Sys}^{Ref}$$

$$T_{Sys}^{Ref} = \left\langle \frac{Ref(\nu)}{Ref_{on}(\nu) - Ref_{off}(\nu)} T_{Cal}(\nu) \right\rangle_{BW}$$

$$T_{Sys}^{Ref}(Elev, \nu, t) \cong T_{rcvr}(\nu) + T_{spill}(Elev) + T_{cmb} e^{-\tau(\nu, t) \cdot A(Elev)} + T_{Atm}(\nu, t) \cdot (1 - e^{-\tau(\nu, t) \cdot A(Elev, t)})$$

$A(Elev, t)$ = Air Mass

$\tau(\nu, t)$ = Atmospheric Zenith Opacity

Area = Physical area of the telescope

$\eta_A(\nu, Elev)$ = Aperture efficiency (point sources)

$T_A(\nu)$ = Source Antenna Temperature

$S(\nu)$ = Source Flux Density

$Sig(\nu), Ref(\nu)$ = Data taken on source and on blank sky (in units backend counts)

On, Off = Data taken with the noise diode on and off

$T_{sys}(\nu, Elev, t)$ = System Temperature

T_{CMB} = Cosmic Microwave Background

$T_{rcvr}(\nu)$ = Receiver Temperature

$T_{spill}(Elev)$ = Antenna Spillover

$T_{Atm}(\nu, t)$ = Representative temperature of the atmosphere