

WE ARE ALL MADE OF STARS

WITHOUT STARS, WE WOULDN'T BE HERE. STARS
PRODUCE BASIC ELEMENTS, LIKE CARBON, THE
FOUNDATION OF THE COMPLEX MOLECULES WHICH
MAKE UP LIFE.

AS • TRO • CHEM • IS • TRY

/ astrō'keməstrē /

(n) the study of molecules in space - where they are, how they got there, and what they are doing



THE BIG QUESTION

Where do we come from?



What is the interstellar molecular inventory?

How does chemistry evolve toward complexity?

How does physical environment impact chemical evolution?

What is the interstellar molecular inventory?

How does chemistry evolve toward complexity?

How does physical environment impact chemical evolution?

What molecules are good tracers of astrophysical processes and properties?

What is the interstellar molecular inventory?

How does chemistry evolve toward complexity?

How does physical environment impact chemical evolution?

~~What molecules are good tracers of astrophysical processes and properties?~~

GBT facilities are, and always have been, at the forefront
of astrochemical discovery

'Minor' technical upgrades will enable significant new
science

The upgrade we really need is political, not technical

VOLUME 22, NUMBER 13

PHYSICAL REVIEW LETTERS

31 MARCH 1969

MICROWAVE DETECTION OF INTERSTELLAR FORMALDEHYDE**Lewis E. Snyder and David Buhl**

National Radio Astronomy Observatory,* Green Bank, West Virginia 22901

and

B. Zuckerman

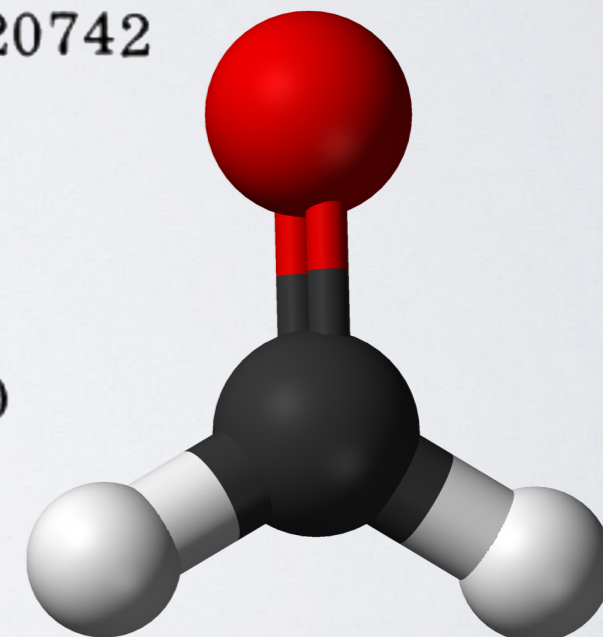
University of Maryland, College Park, Maryland 20742

and

Patrick Palmer

University of Chicago, Chicago, Illinois 60680

(Received 17 March 1969)



THE ASTROPHYSICAL JOURNAL, 162:L203–L210, December 1970

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DETECTION OF METHYL ALCOHOL IN SAGITTARIUS

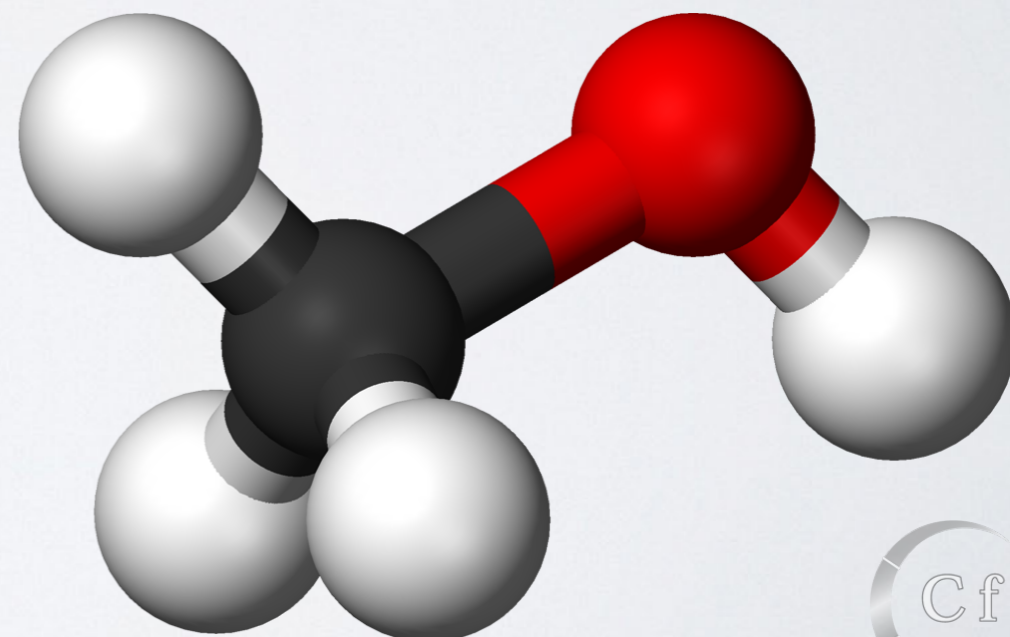
JOHN A. BALL, CARL A. GOTTLIEB, AND A. E. LILLEY
Harvard College Observatory, Cambridge, Massachusetts

AND

H. E. RADFORD

Smithsonian Astrophysical Observatory, Cambridge, Massachusetts

Received 1970 October 31



THE ASTROPHYSICAL JOURNAL, 163:L41-L45, 1971 January 15

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MICROWAVE DETECTION OF INTERSTELLAR FORMIC ACID

B. ZUCKERMAN

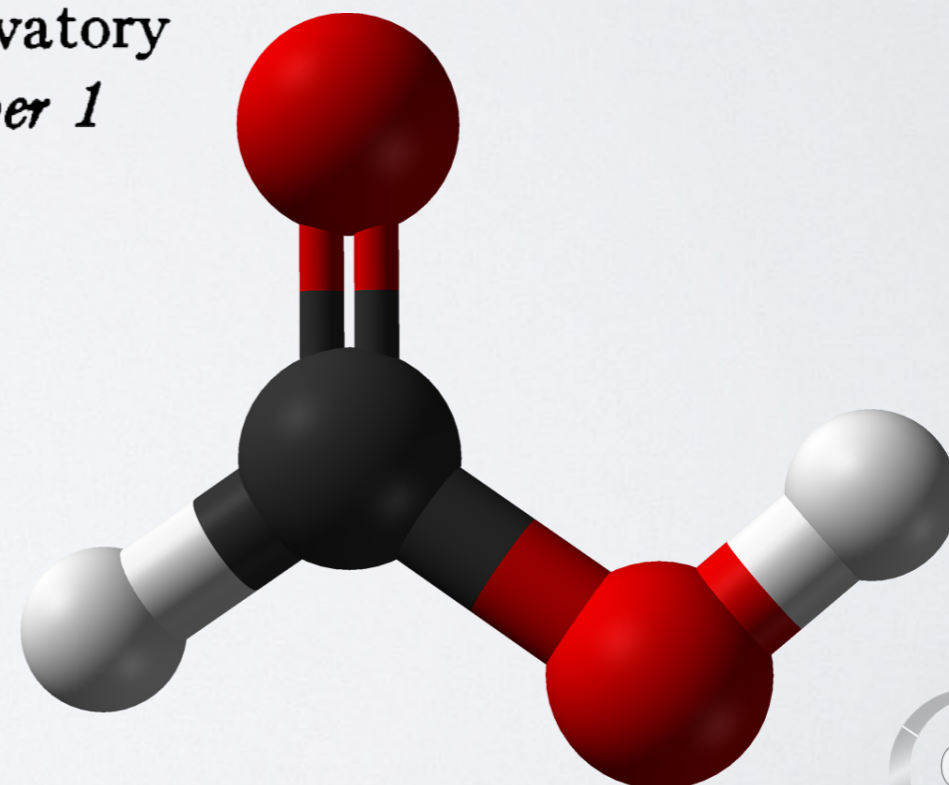
University of Maryland

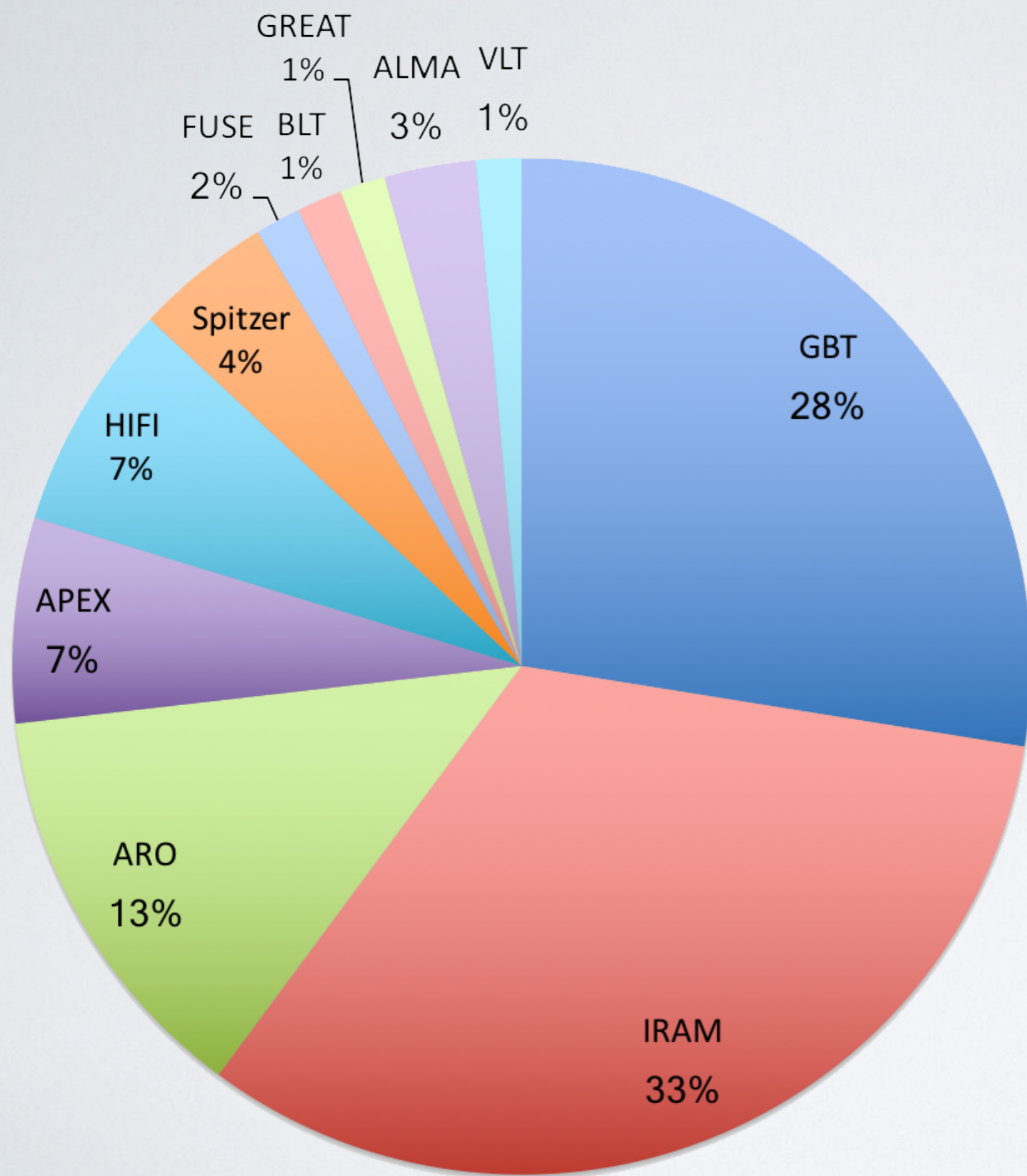
AND

JOHN A. BALL AND CARL A. GOTTLIEB

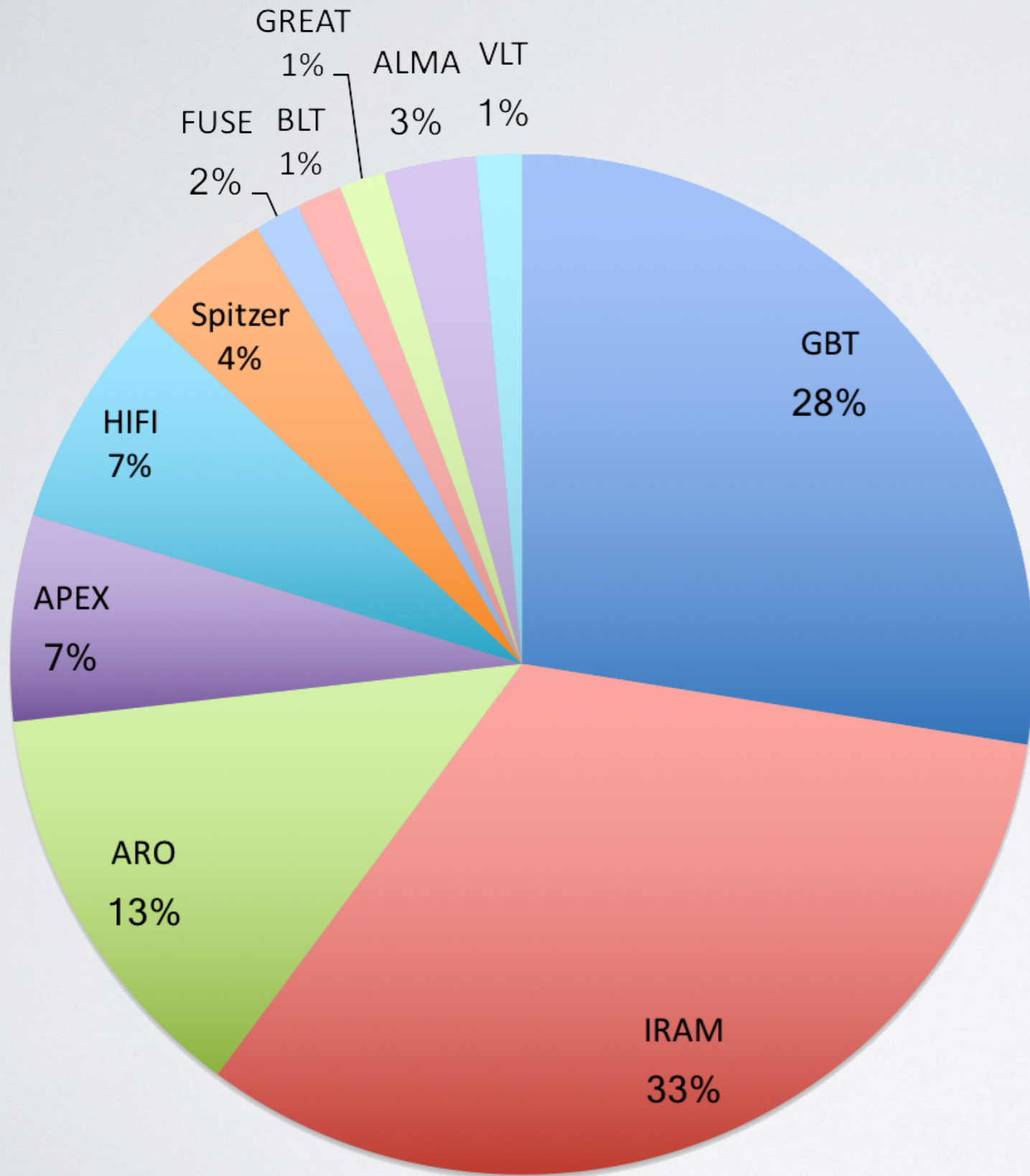
Harvard College Observatory

Received 1970 December 1



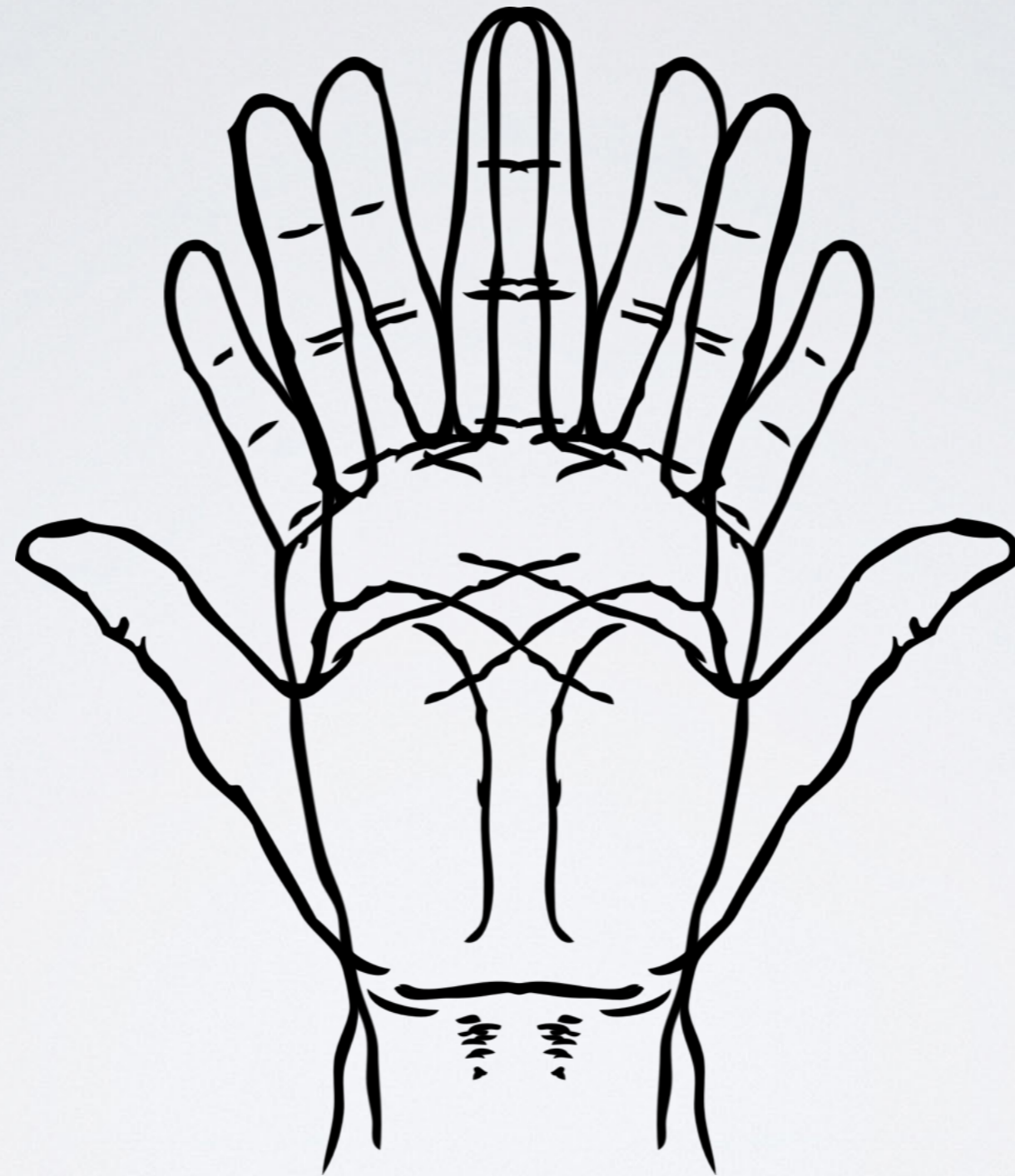


- | | |
|-------------------------------------|-------------------------------------|
| CH ₂ CHCHO | C ₈ H ⁻ |
| CH ₃ CH ₂ CHO | CNCHO |
| CH ₂ CCHCN | HNCNH |
| c-H ₂ C ₃ O | CH ₃ CNHN |
| CH ₂ CONH ₂ | <i>E</i> -HNCHCN |
| CH ₃ C ₆ H | CH ₃ CHCH ₂ O |
| CH ₂ CNH | HC ₅ O |
| CH ₃ C ₅ N | HC ₇ O |
| C ₆ H ⁻ | |



- | | |
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| c-H ₂ C ₃ O | CH ₃ CNHN |
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| CH ₂ CNH | HC ₅ O |
| CH ₃ C ₅ N | HC ₇ O |
| C ₆ H ⁻ | |

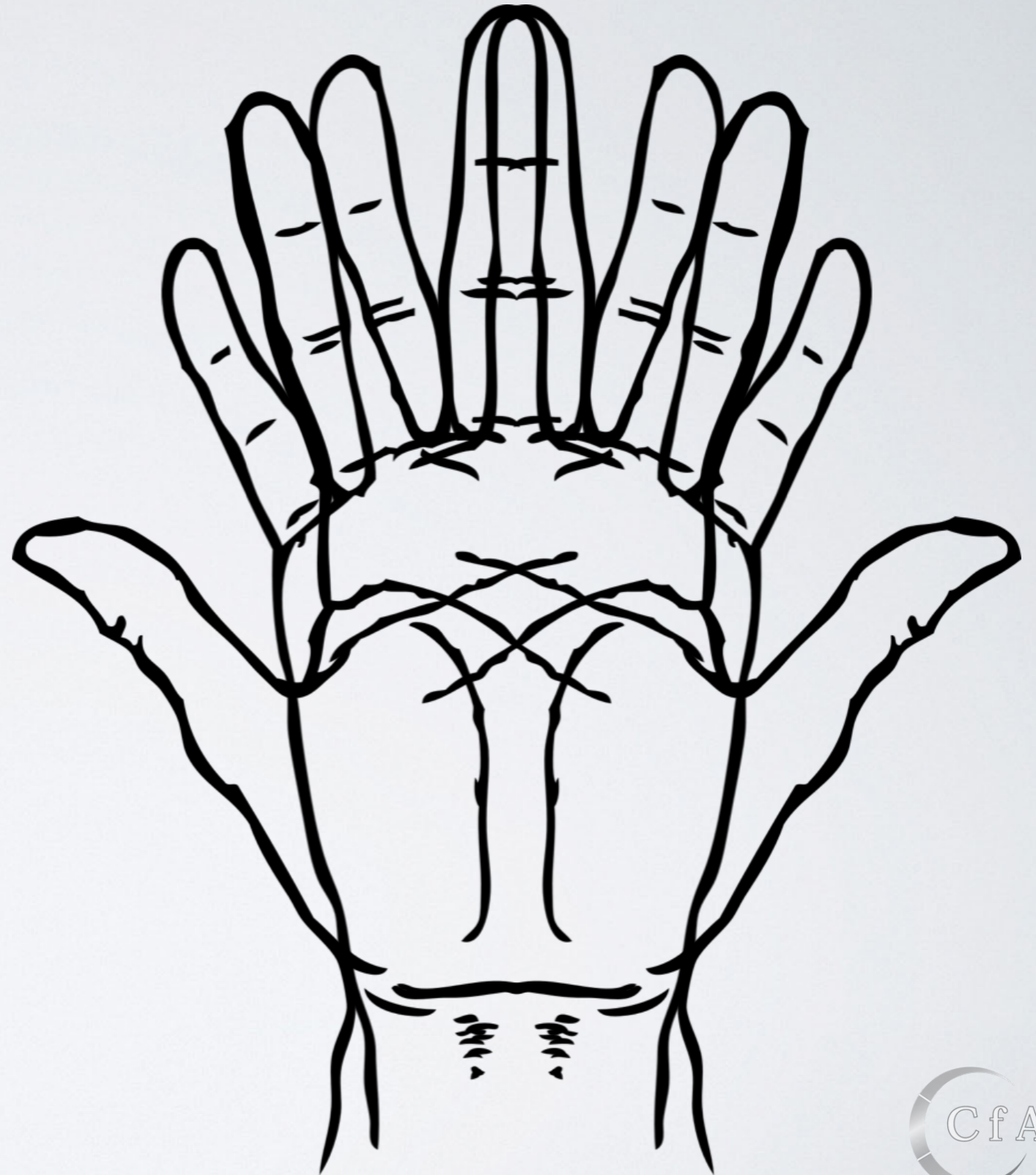




Chirality

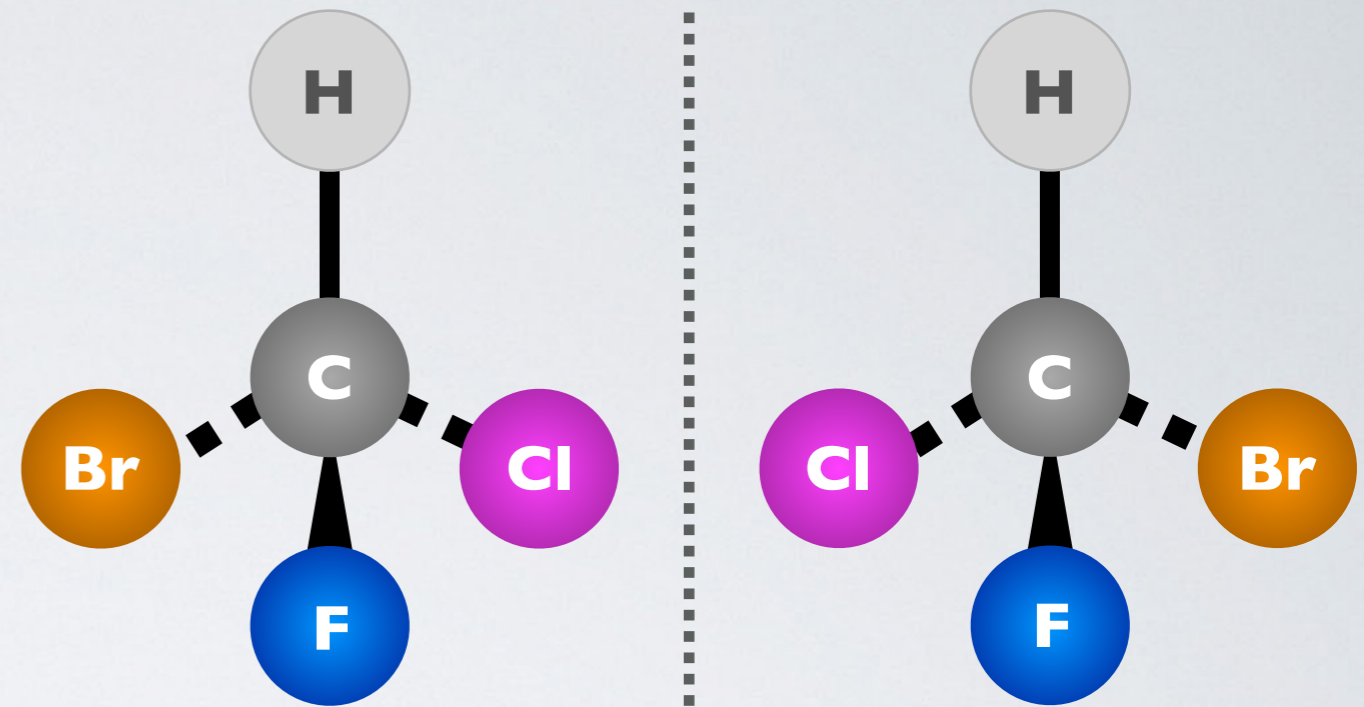
καί'ράλιτι

χείρ (kheir) - 'hand'



Enantiomer

ἐναντιόμορμη

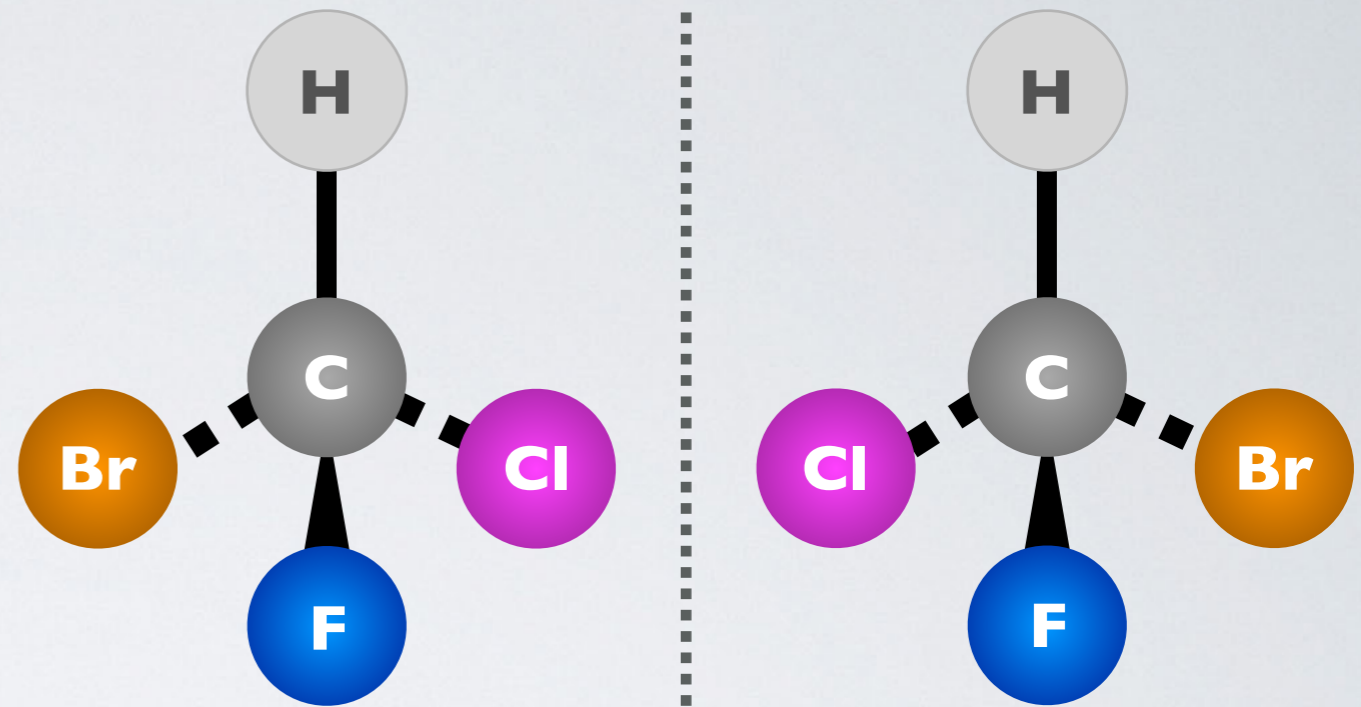


ἐνάντιος (enántios) - 'opposite'

μέρος (méros) - 'part'

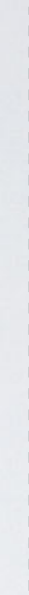
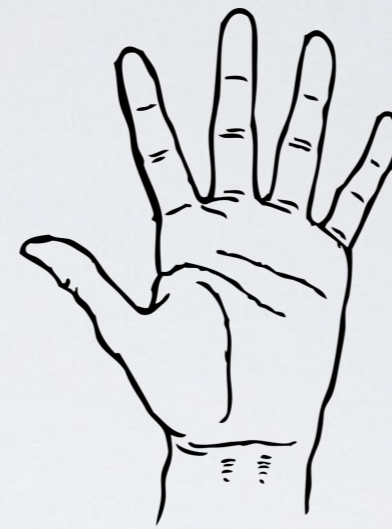
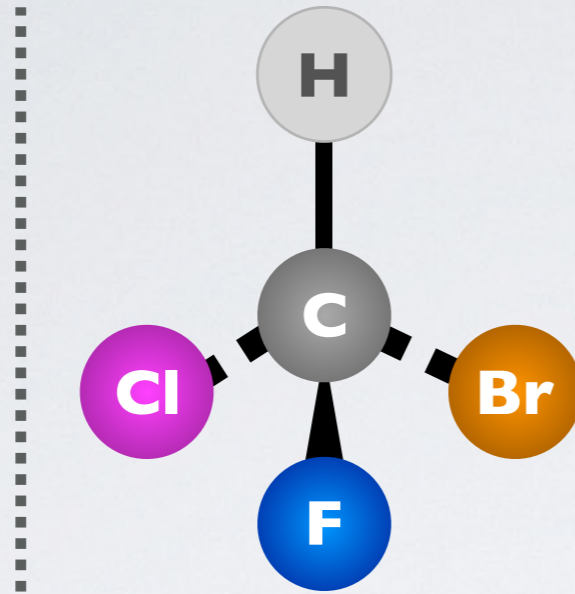
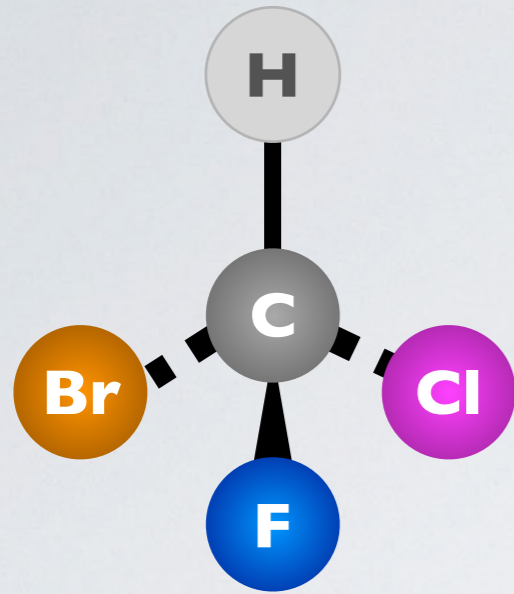
Enantiomer

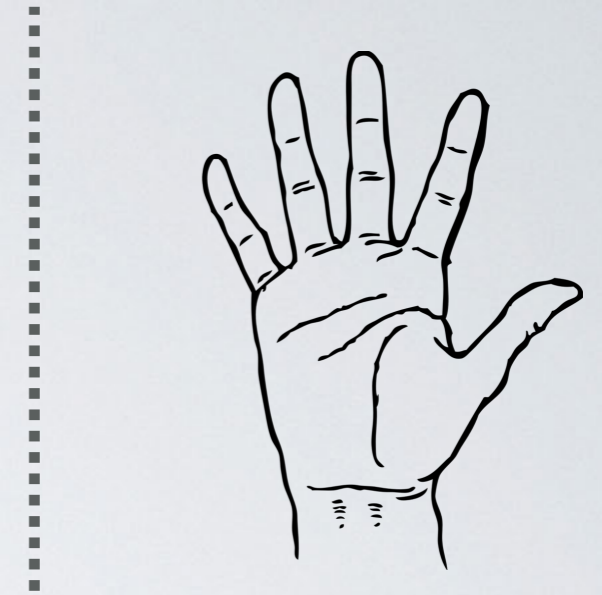
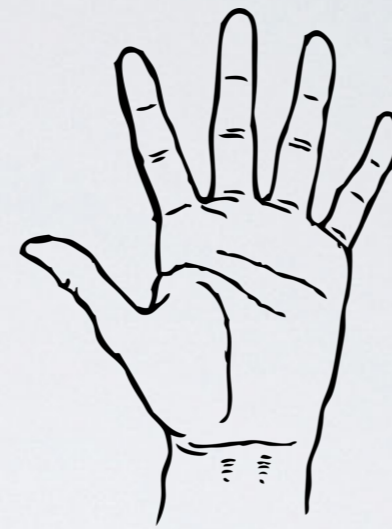
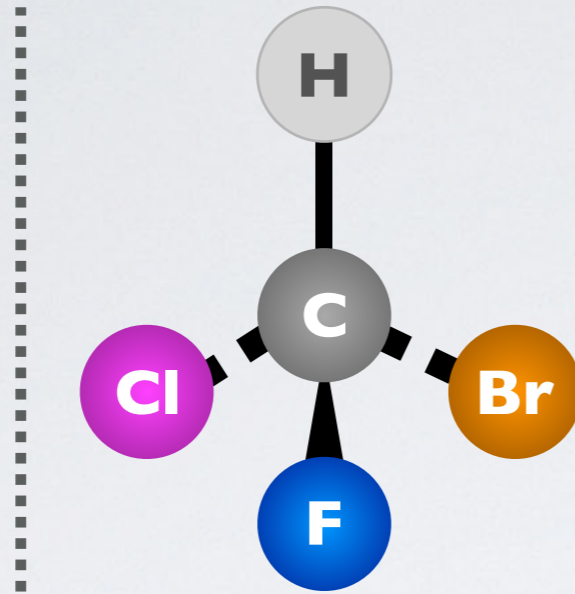
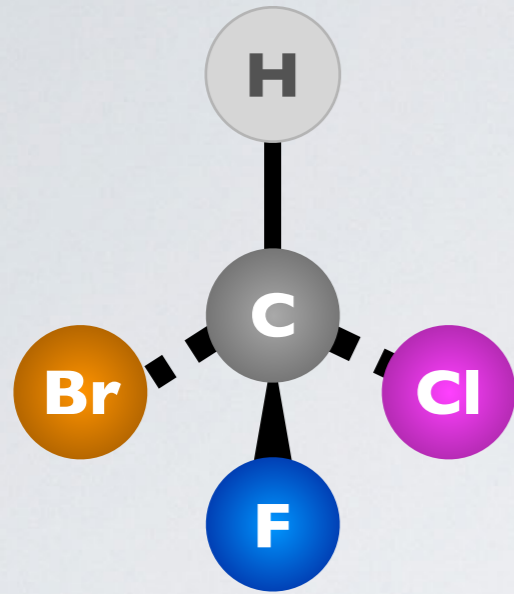
ἐναντιώμορ

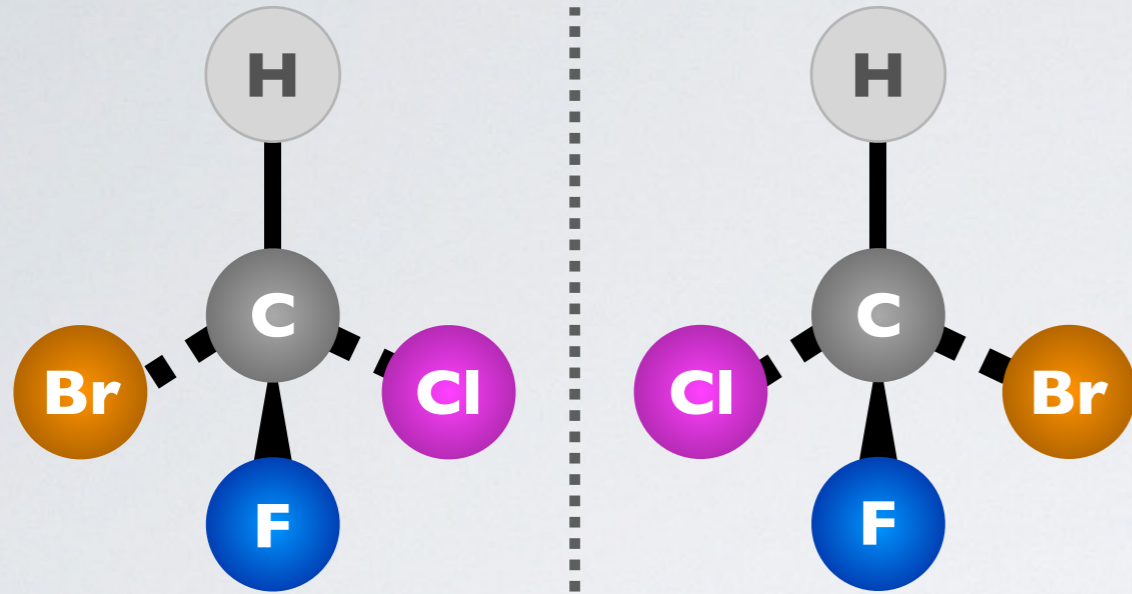


ἐνάντιος (enántios) - 'opposite'
μέρος (méros) - 'part'

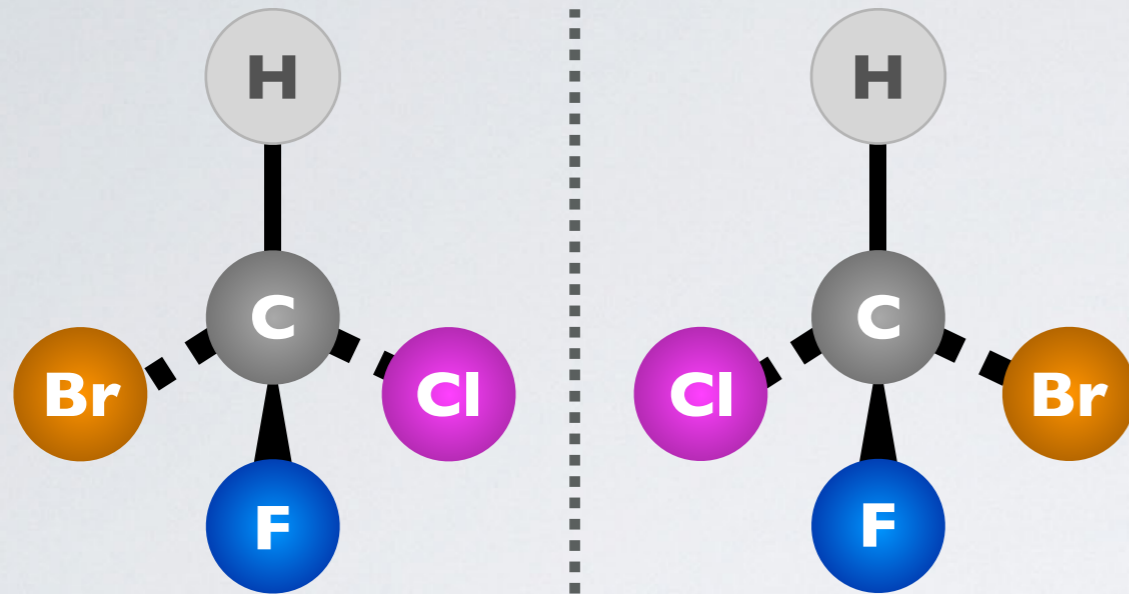
	Left		Right	
(<i>levorotory</i>)	L		D	(<i>dextrorotory</i>)
(<i>sinister</i>)	S		R	(<i>rectus</i>)



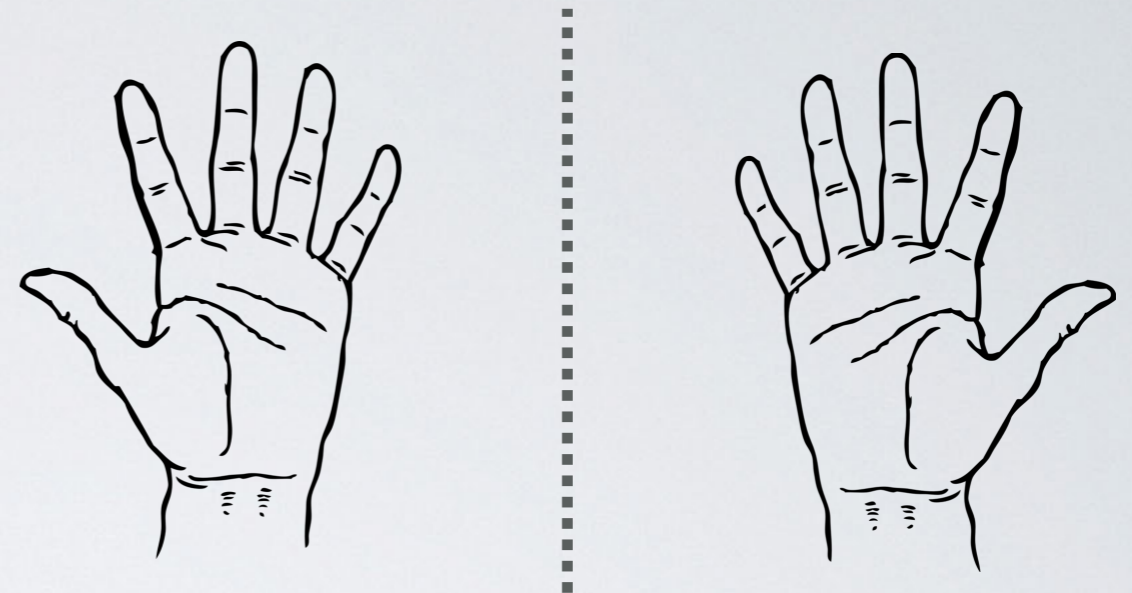




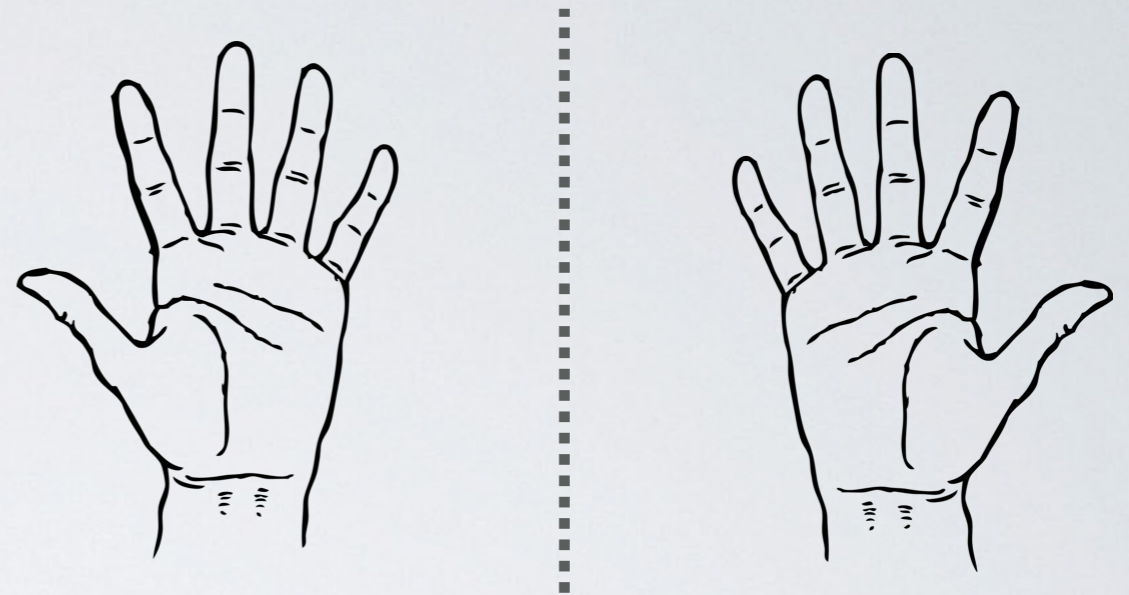
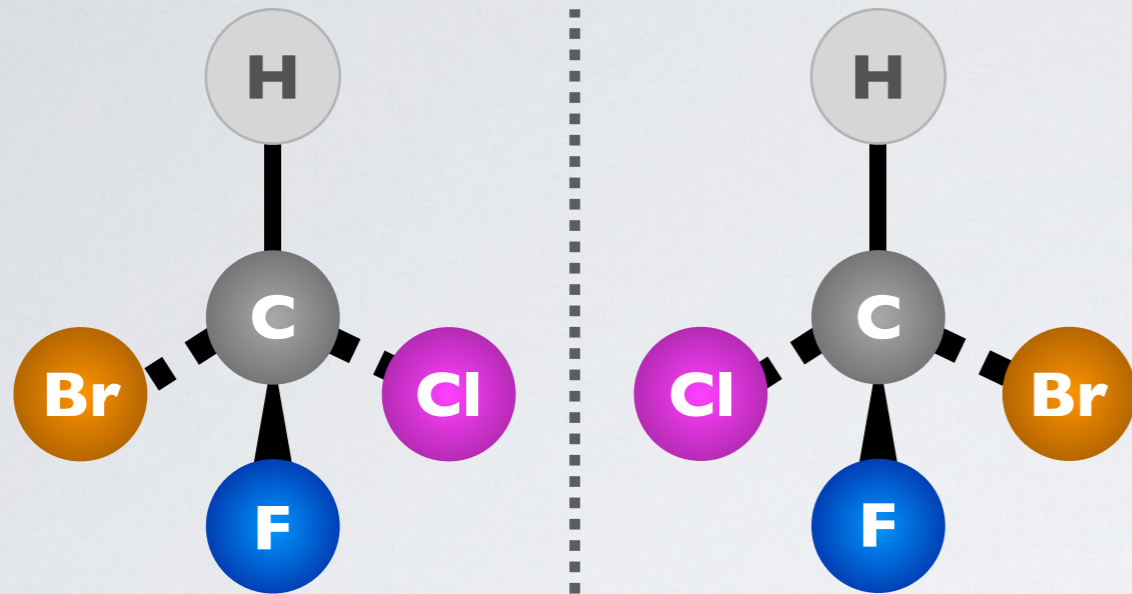
Same atoms, bonds



Same atoms, bonds



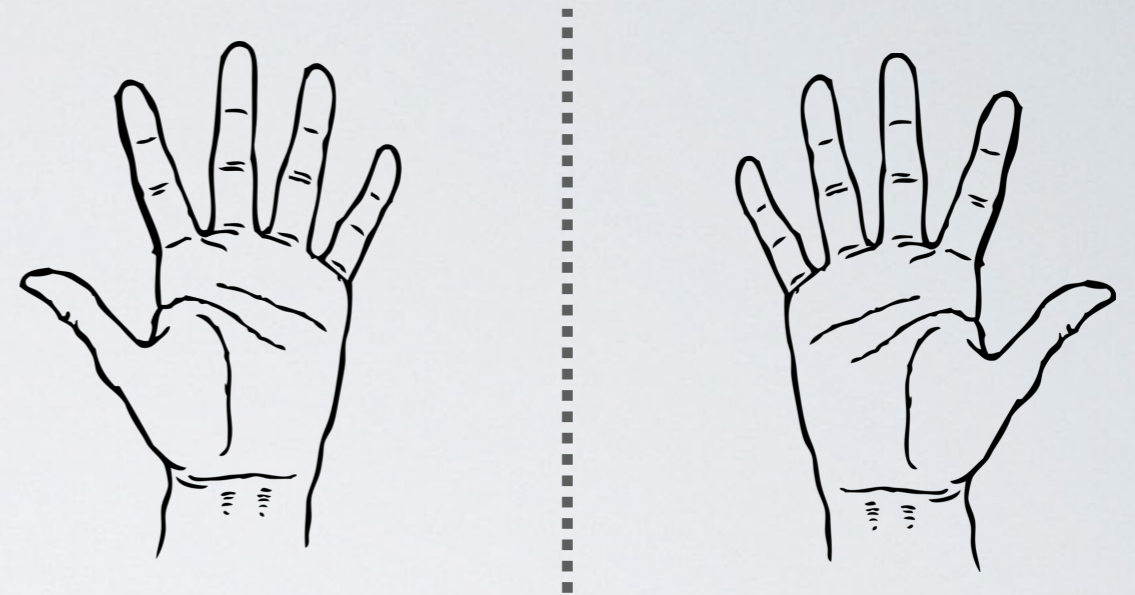
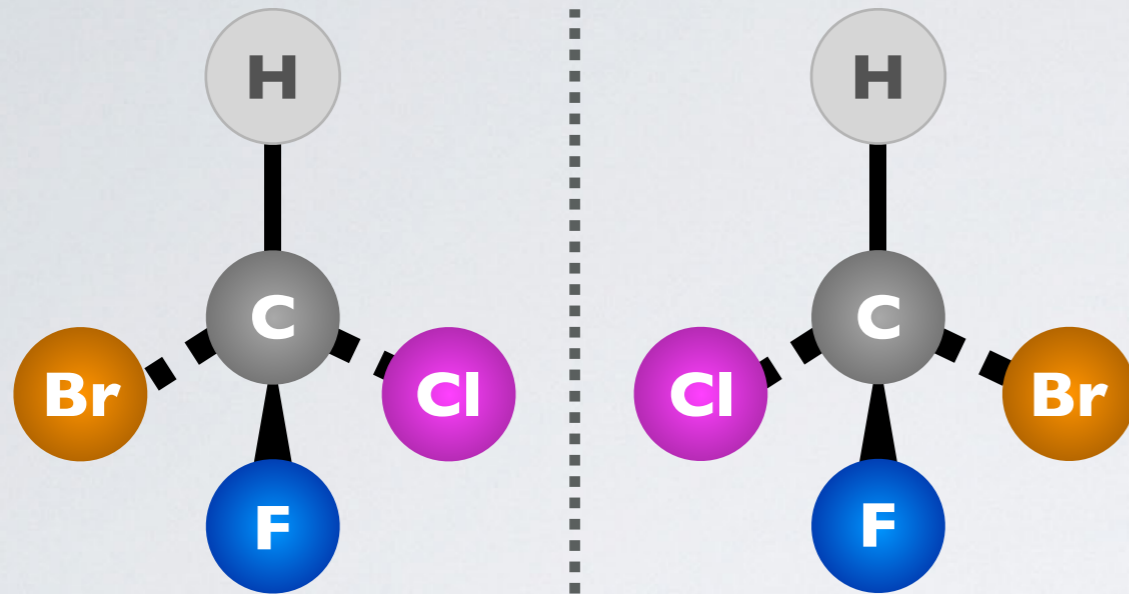
Same bones, tendons



Same atoms, bonds

Same bones, tendons

Same melting/boiling/freezing points

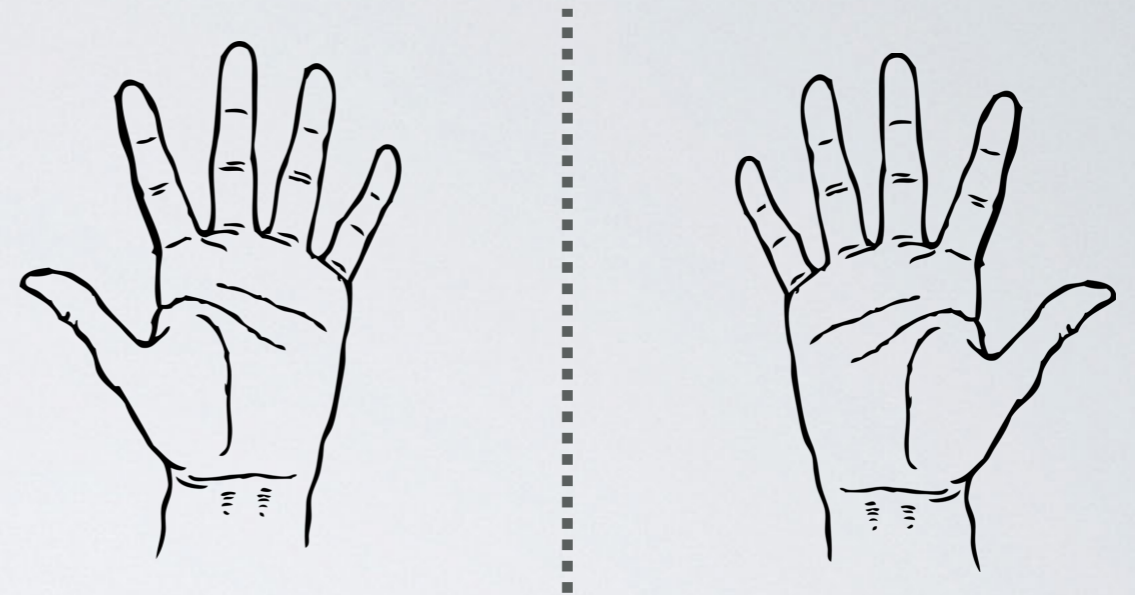
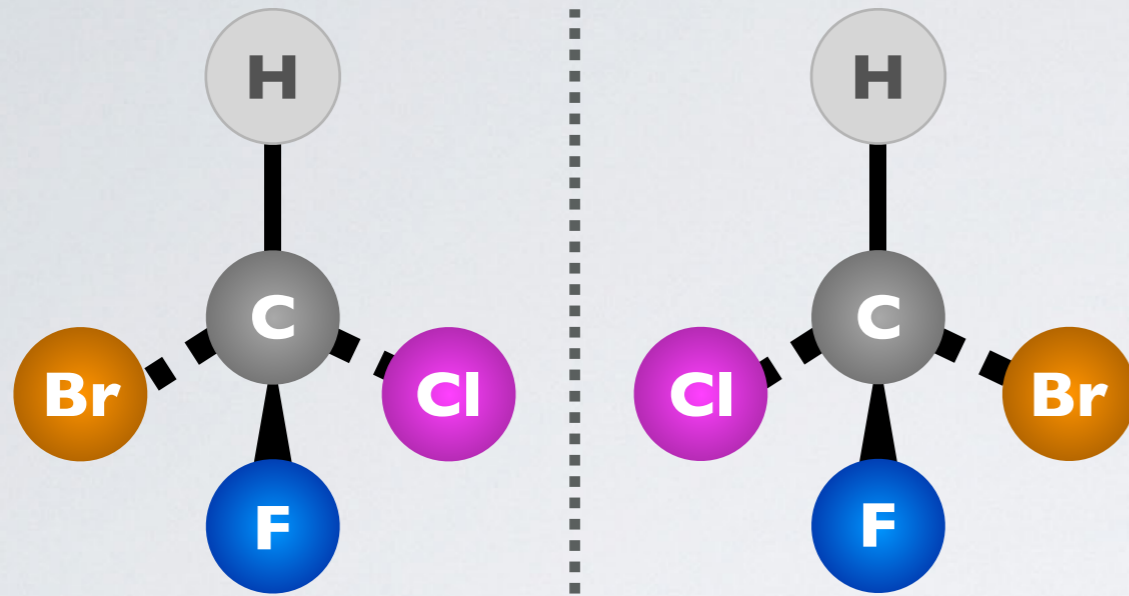


Same atoms, bonds

Same bones, tendons

Same melting/boiling/freezing points

Same melting/boiling/freezing points



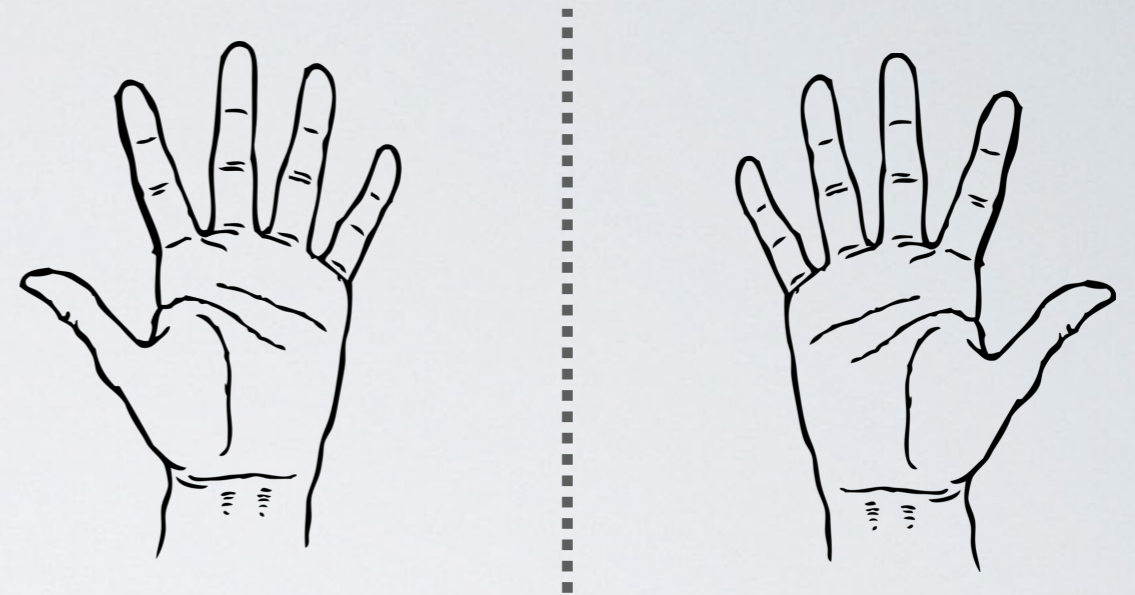
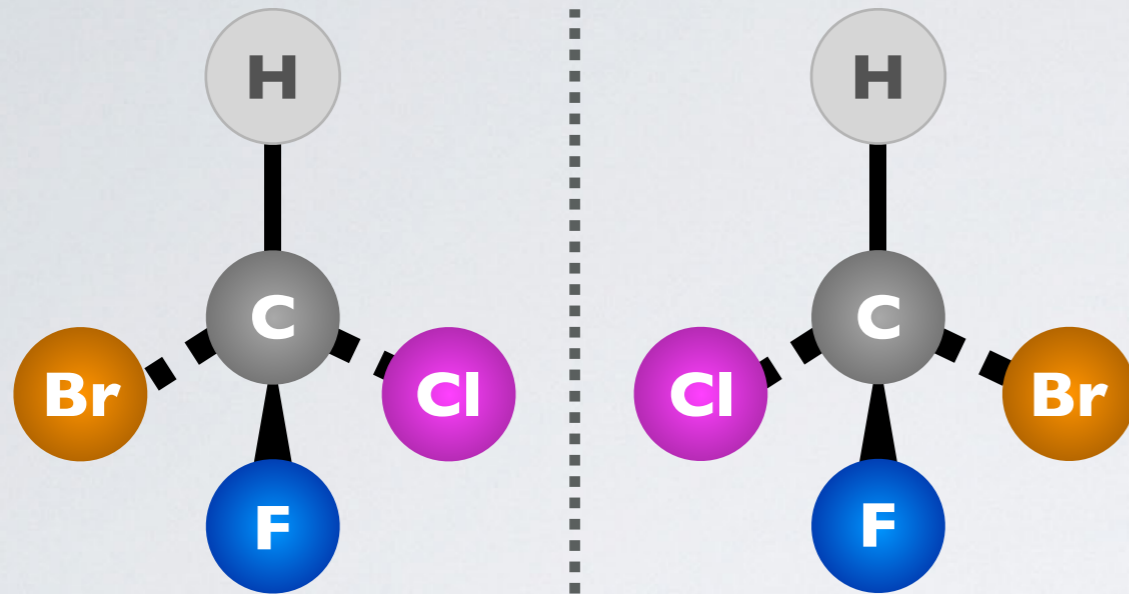
Same atoms, bonds

Same bones, tendons

Same melting/boiling/freezing points

Same melting/boiling/freezing points

Same spectra



Same atoms, bonds

Same bones, tendons

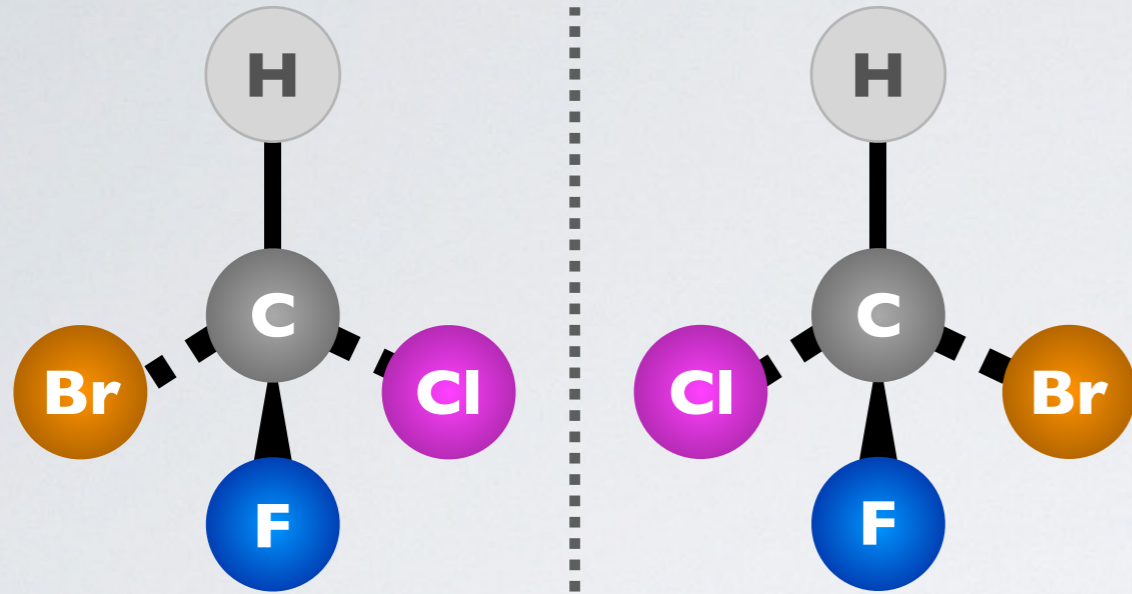
Same melting/boiling/freezing points

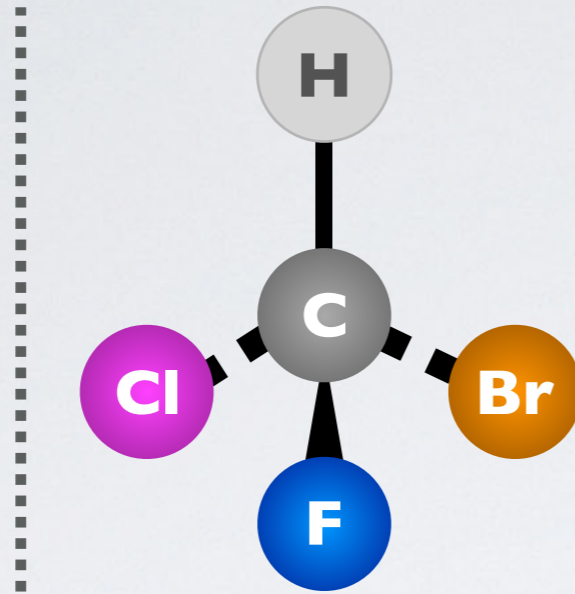
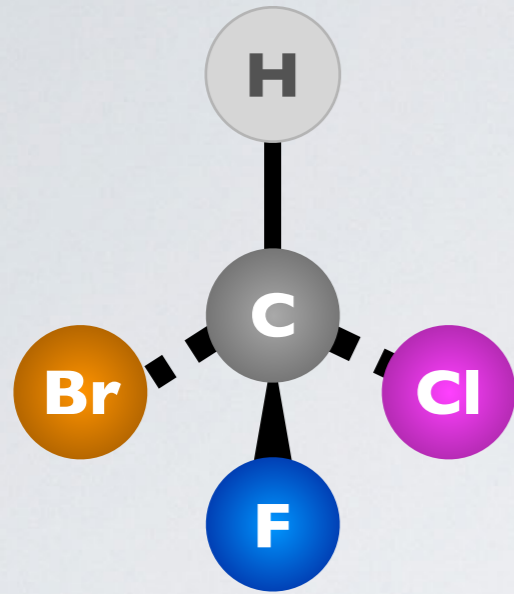
Same melting/boiling/freezing points

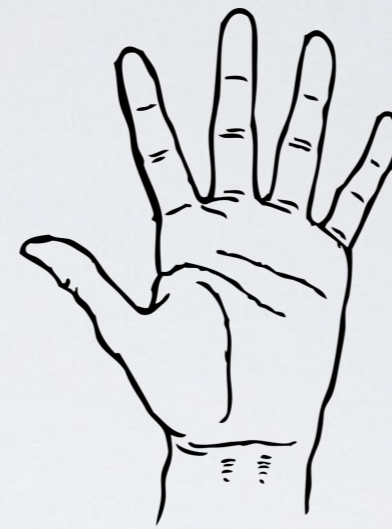
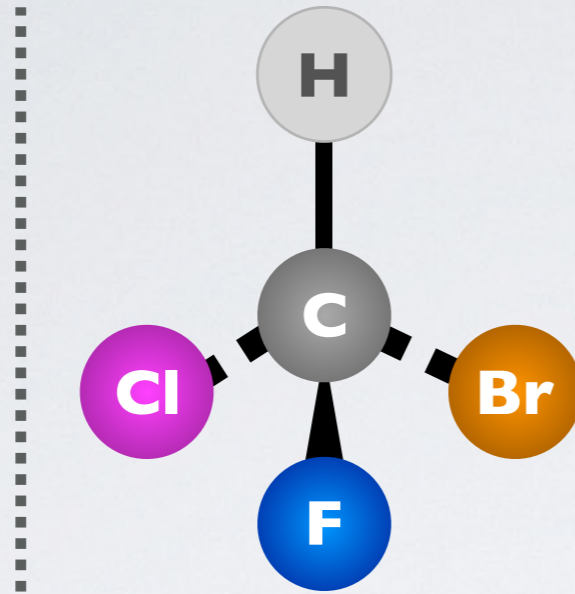
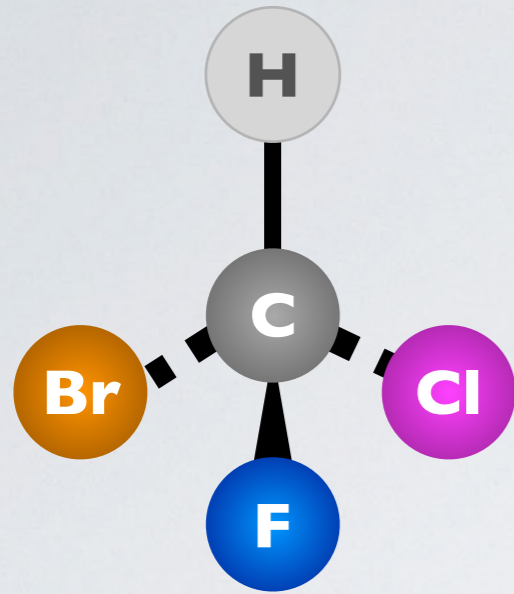
Same spectra

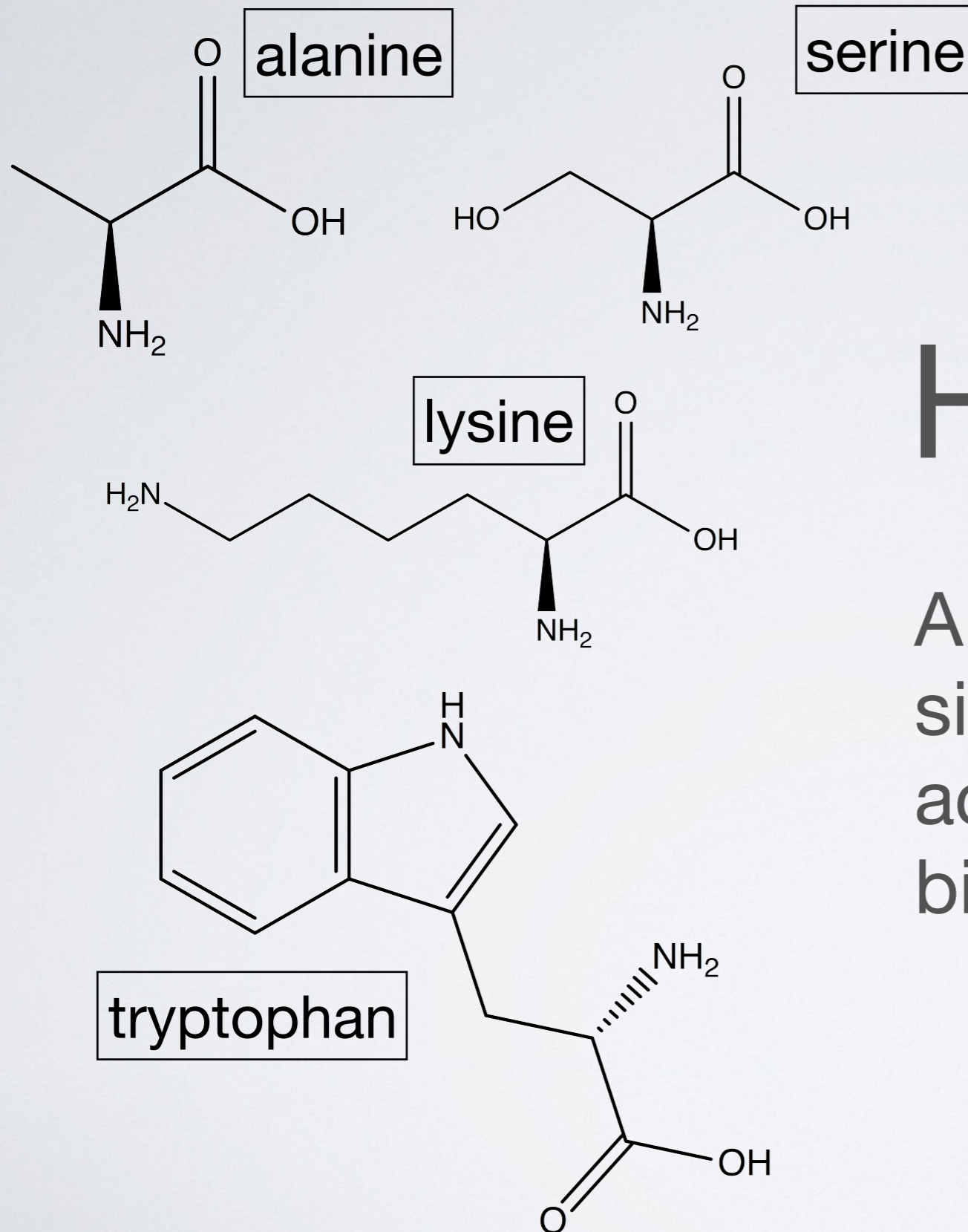
Same shadow

BUT DIFFERENT









Homochirality

All life on Earth uses only a single enantiomer of amino acids, sugars, and other biomolecules

Enantiomeric Excesses in Meteoritic Amino Acids

John R. Cronin and Sandra Pizzarello

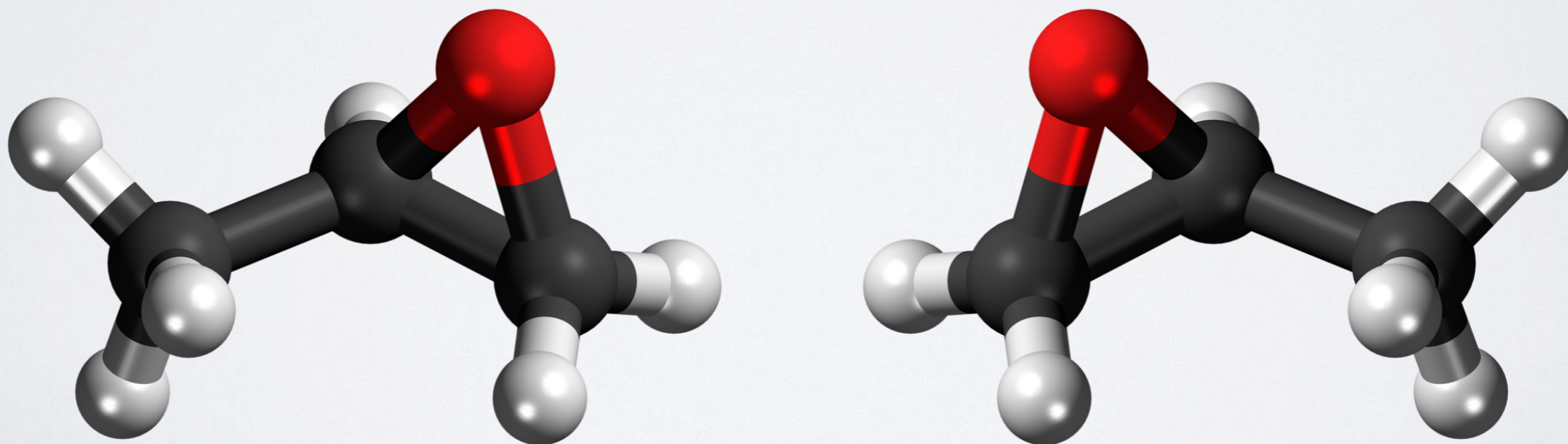


A few amino acids show excess of *L* by almost 10%

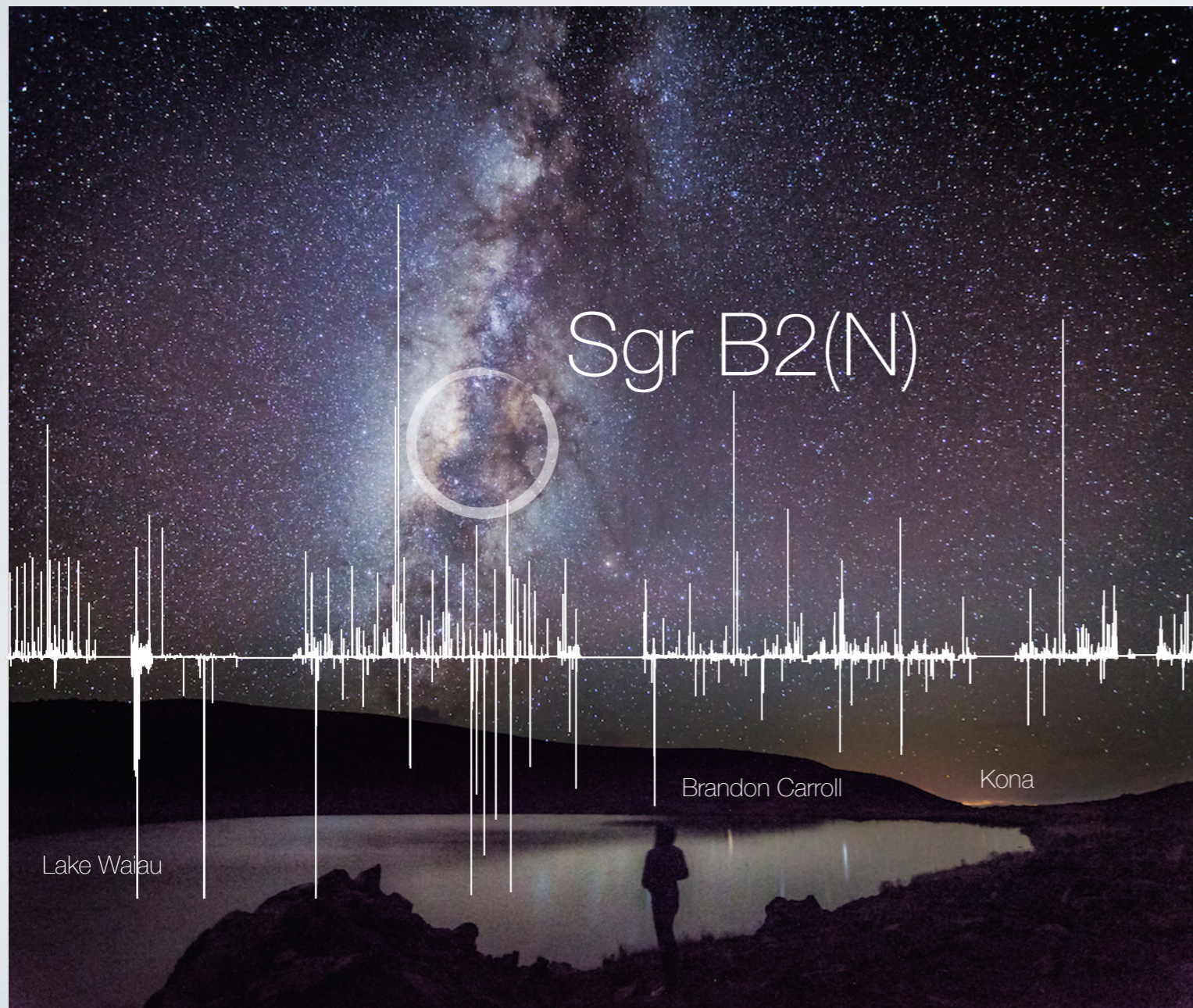
But where did it come from!?

Discovery of the interstellar chiral molecule propylene oxide ($\text{CH}_3\text{CHCH}_2\text{O}$)

Brett A. McGuire,^{1,2*}† P. Brandon Carroll,^{2*}† Ryan A. Loomis,³ Ian A. Finneran,² Philip R. Jewell,¹ Anthony J. Remijan,¹ Geoffrey A. Blake^{2,4}



PRebiotic Interstellar MOlecular Survey



Coverage

0.5 - 50 GHz

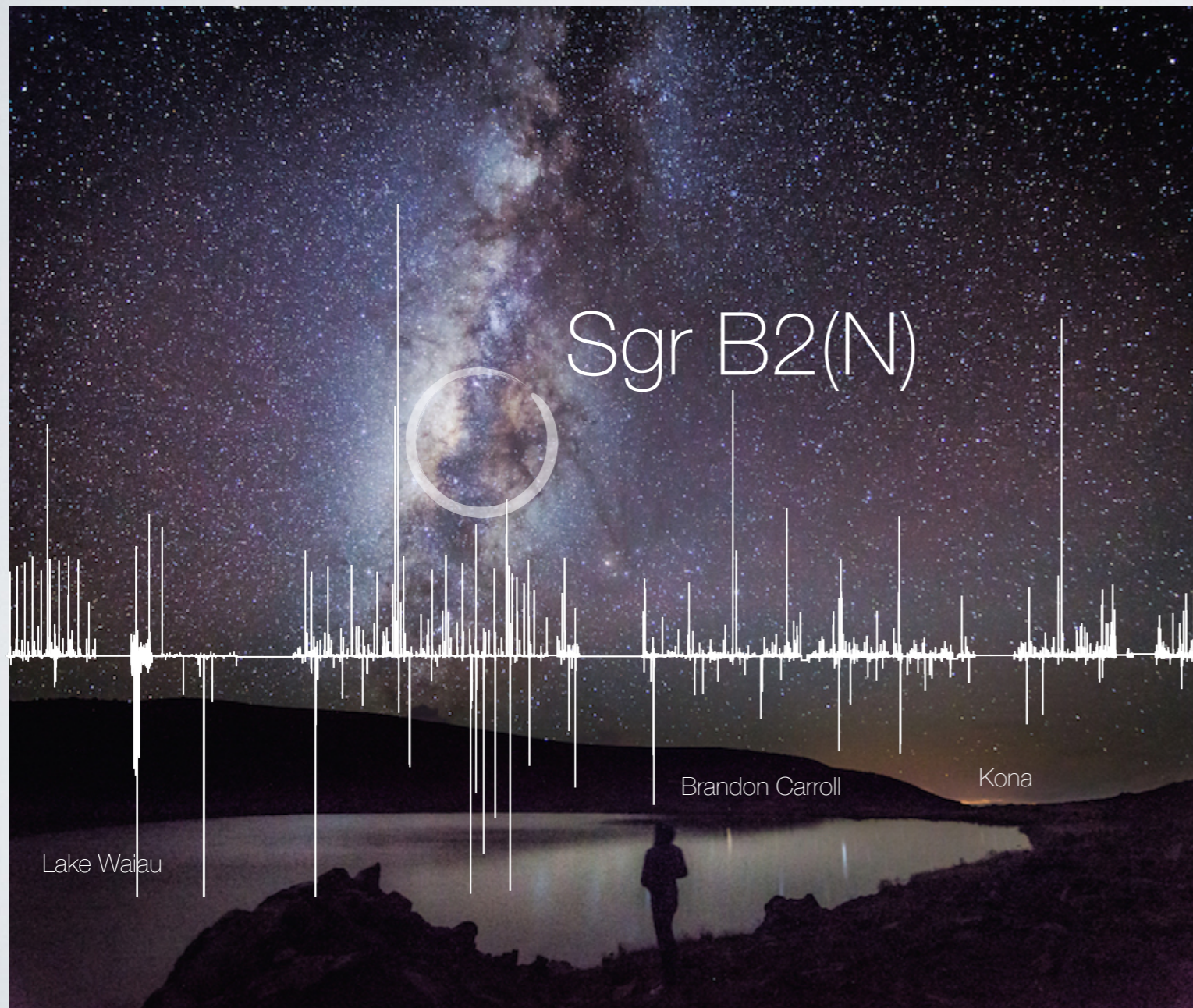
Resolution

24.4 kHz

15 km/s @ 0.5 GHz

0.15 km/s @ 50 GHz

PRebiotic Interstellar MOlecular Survey



Sensitivity @ 15 GHz

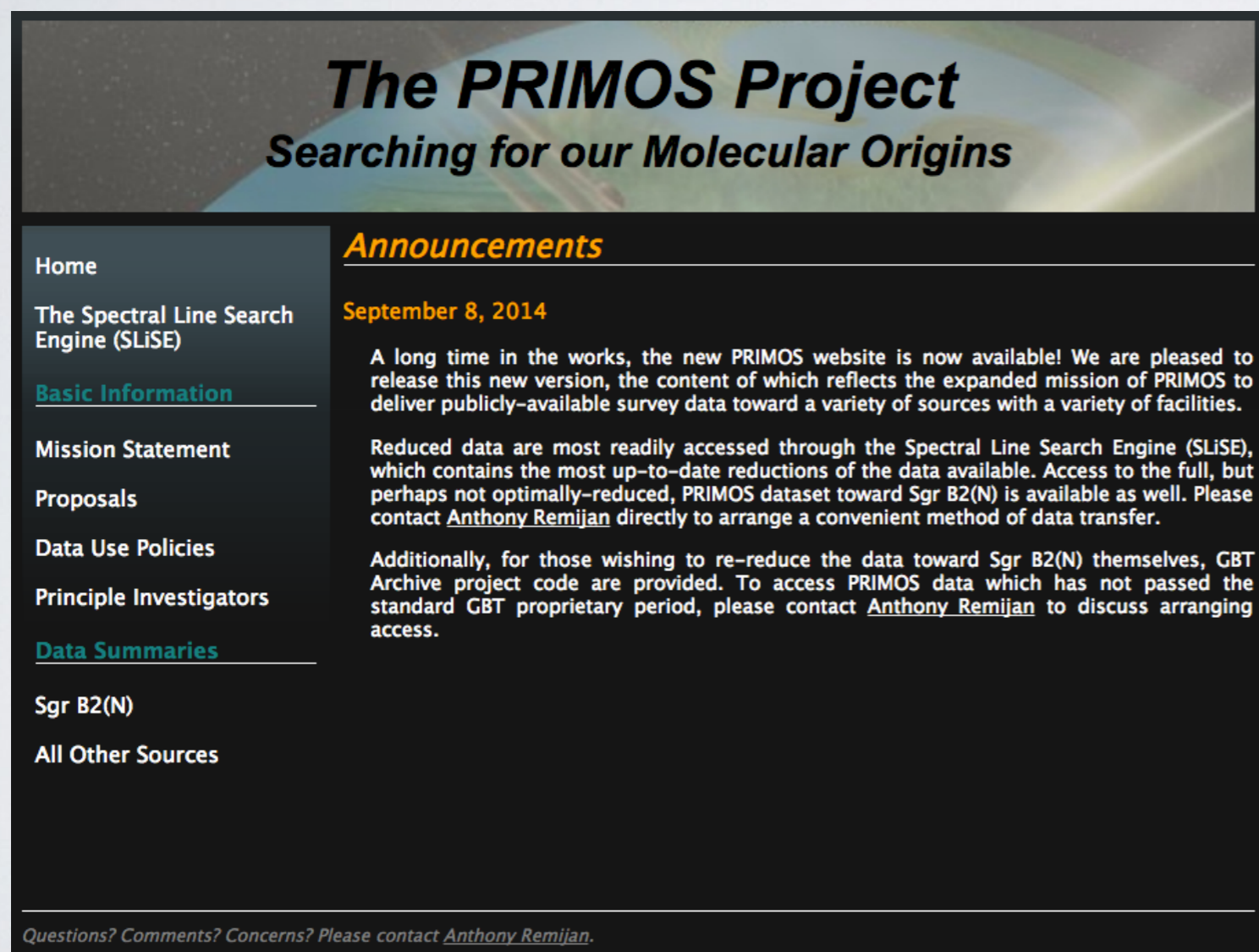
~ 3 mK RMS

Sensitivity @ 50 GHz

~ 10 mK RMS

PRebiotic Interstellar MOlecular Survey

<http://www.cv.nrao.edu/~aremijan/PRIMOS/>



The PRIMOS Project
Searching for our Molecular Origins

Announcements

September 8, 2014

A long time in the works, the new PRIMOS website is now available! We are pleased to release this new version, the content of which reflects the expanded mission of PRIMOS to deliver publicly-available survey data toward a variety of sources with a variety of facilities.

Reduced data are most readily accessed through the Spectral Line Search Engine (SLiSE), which contains the most up-to-date reductions of the data available. Access to the full, but perhaps not optimally-reduced, PRIMOS dataset toward Sgr B2(N) is available as well. Please contact [Anthony Remijan](#) directly to arrange a convenient method of data transfer.

Additionally, for those wishing to re-reduce the data toward Sgr B2(N) themselves, GBT Archive project code are provided. To access PRIMOS data which has not passed the standard GBT proprietary period, please contact [Anthony Remijan](#) to discuss arranging access.

Home
The Spectral Line Search Engine (SLiSE)
[Basic Information](#)
Mission Statement
Proposals
Data Use Policies
Principle Investigators
[Data Summaries](#)
Sgr B2(N)
All Other Sources

Questions? Comments? Concerns? Please contact [Anthony Remijan](#).



Publicly Available



Fully reduced



Raw data

PRebiotic Interstellar MOlecular Survey

<http://www.cv.nrao.edu/~aremijan/PRIMOS/>

The PRIMOS Project

Searching for our Molecular Origins

Home

The Spectral Line Search Engine (SLiSE)

[Basic Information](#)

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Data Use Policies

Principle Investigators

[Data Summaries](#)

Sgr B2(N)

All Other Sources

Data Use Policies

Data made available through the PRIMOS project and the SLiSE interface are available for use by the scientific community with "no strings attached." We request, but do not require, that those using the data kindly consider the following:

- Notifications of use (or intent to use) PRIMOS data by contacting [Anthony Remijan](#) are greatly appreciated.
- Suggested citations and/or attributions for the data reduction strategy and original publication of each observational set are given in their respective data summary pages. We would greatly appreciate the citation of these works and/or individuals in publications and presentations.
- When PRIMOS data are used in publications, we would appreciate a footnote in the observational section to the effect of: "Access to the entire PRIMOS data set, specifics on the observing strategy, and overall frequency coverage information is available at <http://www.cv.nrao.edu/~aremijan/PRIMOS/>."
- When data from SLiSE are used in publications, we would appreciate a footnote in the observation section to the effect of: "These observational data are accessible at <http://www.cv.nrao.edu/~aremijan/SLiSE/>."
- The PRIMOS team of PIs and Co-Is includes staff, postdoctoral scholars, graduate students, and undergraduate students in astronomy, astrochemistry, and experimental chemistry who are always interested in pursuing potential collaborations using PRIMOS data. If you are interested in working with a PRIMOS team member, please contact [Anthony Remijan](#).

Data made available through the PRIMOS project and the SLiSE interface are available for use by the scientific community with no strings attached.

Primary Science Result

Reference

Non-detection of propanimine ($\text{CH}_3\text{CNHCH}_3$)	Margulés et al. 2017 (in prep.)
Non-detection of thioacetaldehyde (CH_3CHS)	Margulés et al. 2017 (in prep.)
Detection of ubiquitous complex molecules in diffuse and translucent clouds	Corby et al. 2017 (in prep.)
First discovery of ^{13}C isotopologues of HC_7N	Burkhardt et al. 2017
First discovery of HC_7O	McGuire et al. 2017ba
First discovery of HC_5O	McGuire et al. 2017a
Probing diffuse \rightarrow translucent cloud transitions	Corby et al. 2017
Non-detection of 2-hydroxyacetonitrile (HOCH_2CN)	Margulés et al. (2017)
First discovery of propylene oxide ($\text{CH}_3\text{CH}(\text{O})\text{CH}_2$)	McGuire et al. (2016b)
Search for vinyl formate (CH_2CHOCHO)	Alonso et al. (2016)
Investigations of the [H, N, C, S] isomeric family in the ISM	McGuire et al. (2016a)
Chemical simulations of prebiotic molecules	Quan et al. (2016)
Non-detection of HSO	Cazzoli et al. (2016)
Insights into cyanomethanimine isomers	Vazart et al. (2015)
Insights into kinetic vs thermodynamic-drive chemistry	Loomis et al. (2015)
Refuted detection of <i>trans</i> -ethyl methyl ether ($\text{C}_2\text{H}_5\text{OCH}_3$)	Carroll et al. (2015)
Identification of methyl formate masers in Sgr B2(N)	Faure et al. (2014)
Observational confirmation of identity of B11244 <i>l</i> - C_3H^+	McGuire et al. (2014)
Tentative first discovery of H_2NCO^+	Gupta et al. (2013)
Non-detection of allyl isocyanide ($\text{CH}_2\text{CHCH}_2\text{NC}$)	Haykal et al. (2013)
Detection of B11244 in Sgr B2(N)	McGuire et al. (2013)
First discovery of ethanimine (CH_3CHNH)	Loomis et al. (2013)
First discovery of <i>E</i> -cyanomethanimine (HNCHCN)	Zaleski et al. (2013)
First discovery of carbodiimide (HNCNH)	McGuire et al. (2012)
Tentative first discovery of <i>trans</i> -methyl formate	Neill et al. (2012)
Insights into spatial distributions and interstellar reactions	Neill et al. (2011)
Line profile analysis of hydrogen radio recombination lines	von Prochazka et al. (2010)
Non-detection of 1,2-propanediol ($\text{CH}_2\text{OHCH}_2\text{CH}_2\text{OH}$)	Plusquellic et al. (2009) Lovas et al. (2009)

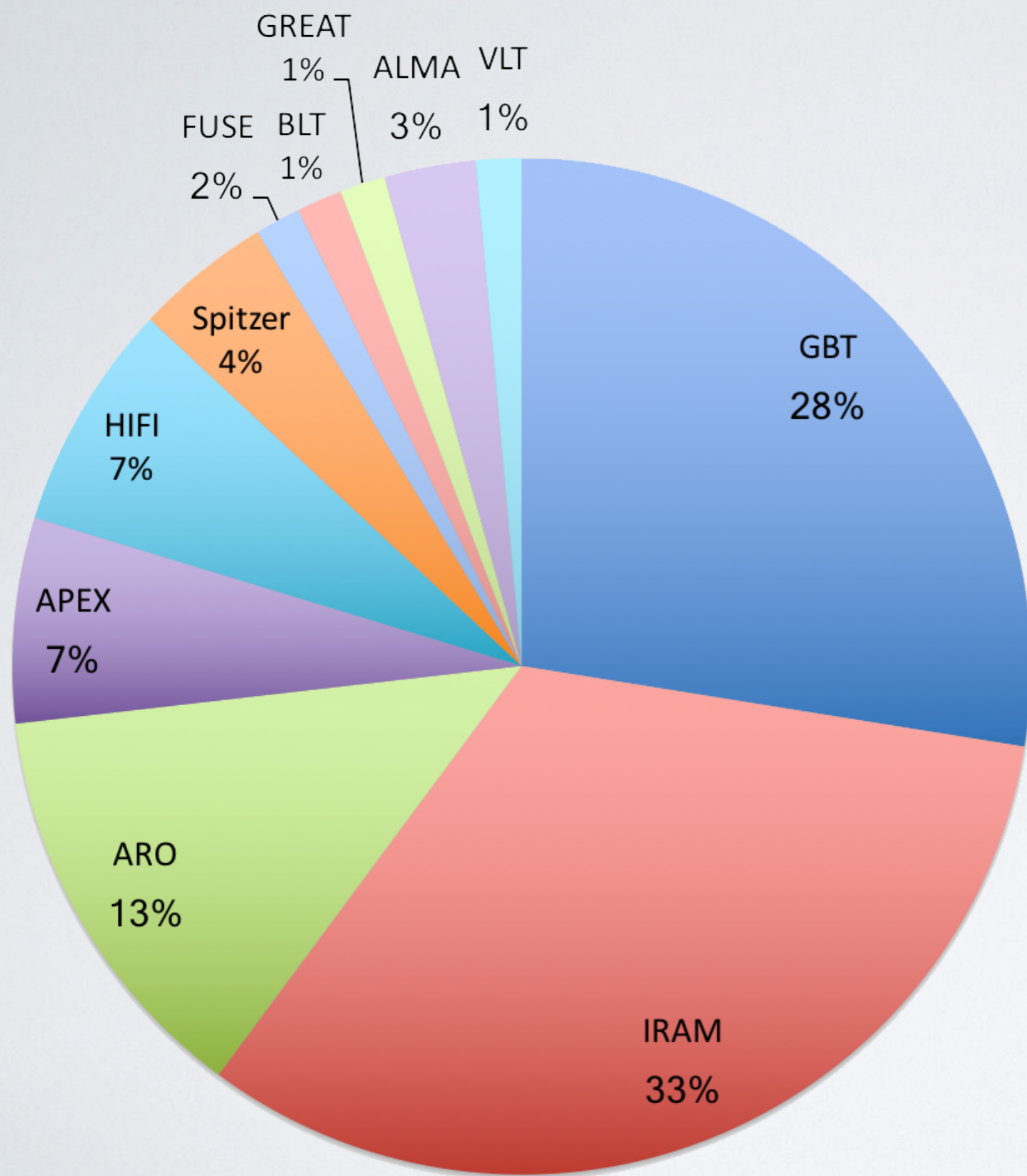
29 Total

ALMA and other mm/sub-mm facilities have the monopoly on 'traditional' sources
(hot cores, outflows, etc.)

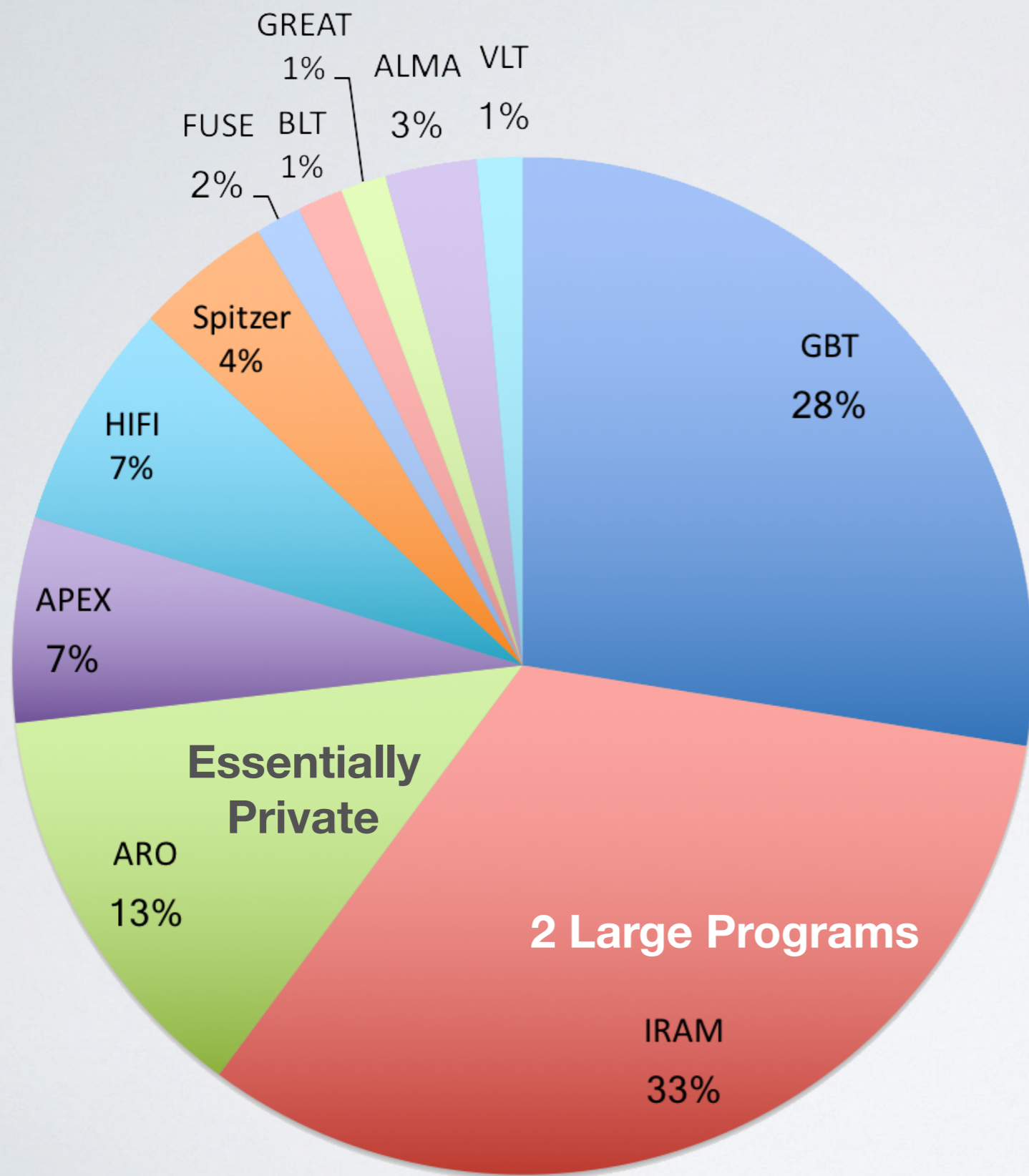
ALMA and other mm/sub-mm facilities have the monopoly on 'traditional' sources
(hot cores, outflows, etc.)

GBT excels at understanding the inventories and chemistry occurring BEFORE these later stages

Absolutely essential; if you don't have $t = 0$ correct, everything after that is hogwash and hand waving



- | | |
|-------------------------------------|-------------------------------------|
| CH ₂ CHCHO | C ₈ H ⁻ |
| CH ₃ CH ₂ CHO | CNCHO |
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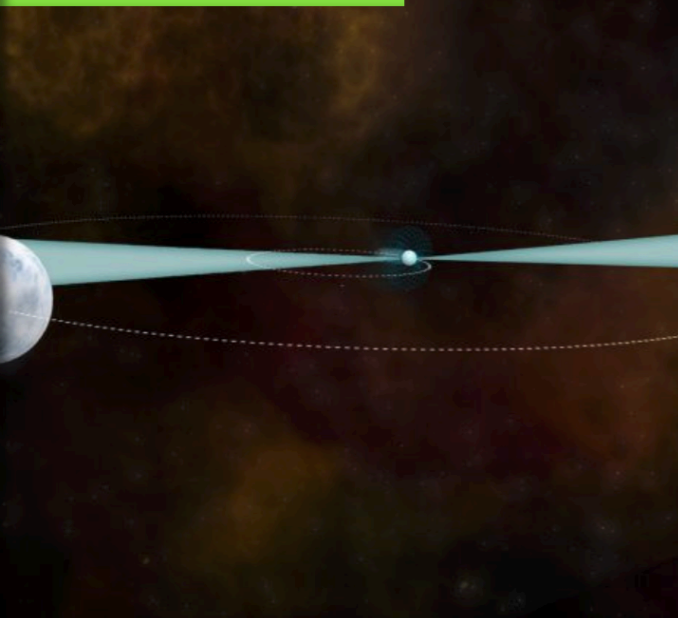
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| CH ₃ C ₅ N | HC ₇ O |
| C ₆ H ⁻ | |

IRAM: Astrochemistry featured as 'key science' in every annual report since 2004

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WITHOUT STARS, WE WOULDN'T BE HERE. STARS
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FOUNDATION OF THE COMPLEX MOLECULES WHICH
MAKE UP LIFE.

FUNDAMENTAL PHYSICS



ORIGIN OF LIFE

Connect chemistry in space with life on Earth.



THE BIRTH OF PLANETS

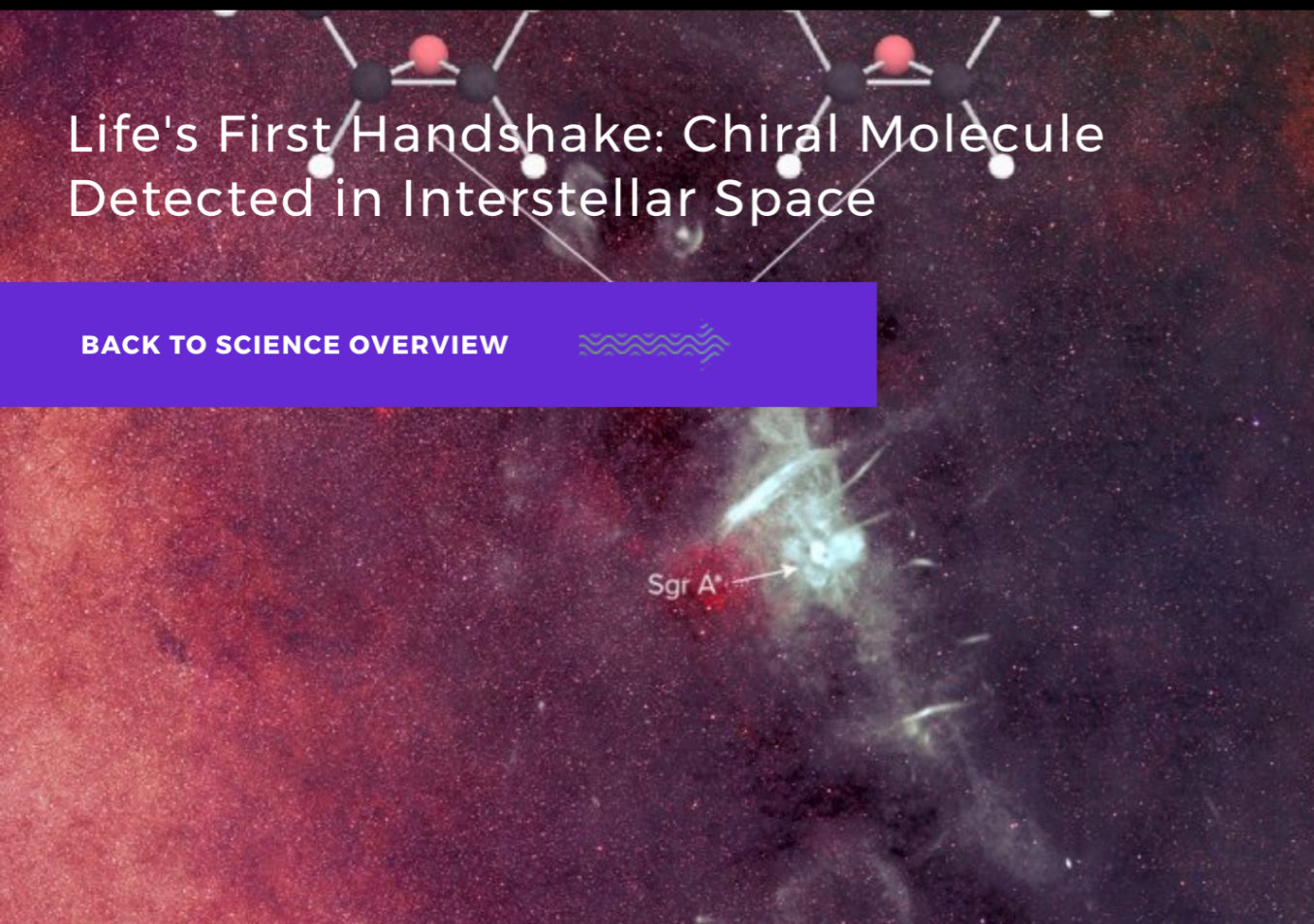


STARS, GAS, AND GALAXIES

PHOTO: GBT
B. Saxton, NRAO/AUI/NSF
from data provided by F. Lockman
Sky photo: A. Mellinger



size of full Moon



Life's First Handshake: Chiral Molecule Detected in Interstellar Space

BACK TO SCIENCE OVERVIEW



CONNECT ORGANIC CHEMISTRY IN SPACE WITH LIFE ON EARTH

HOW COMMON ARE THE BUILDING BLOCKS OF LIFE?

The GBT is used to discover and measure interstellar chemical processes to determine the characteristics of pre-biotic chemistry in star-forming regions

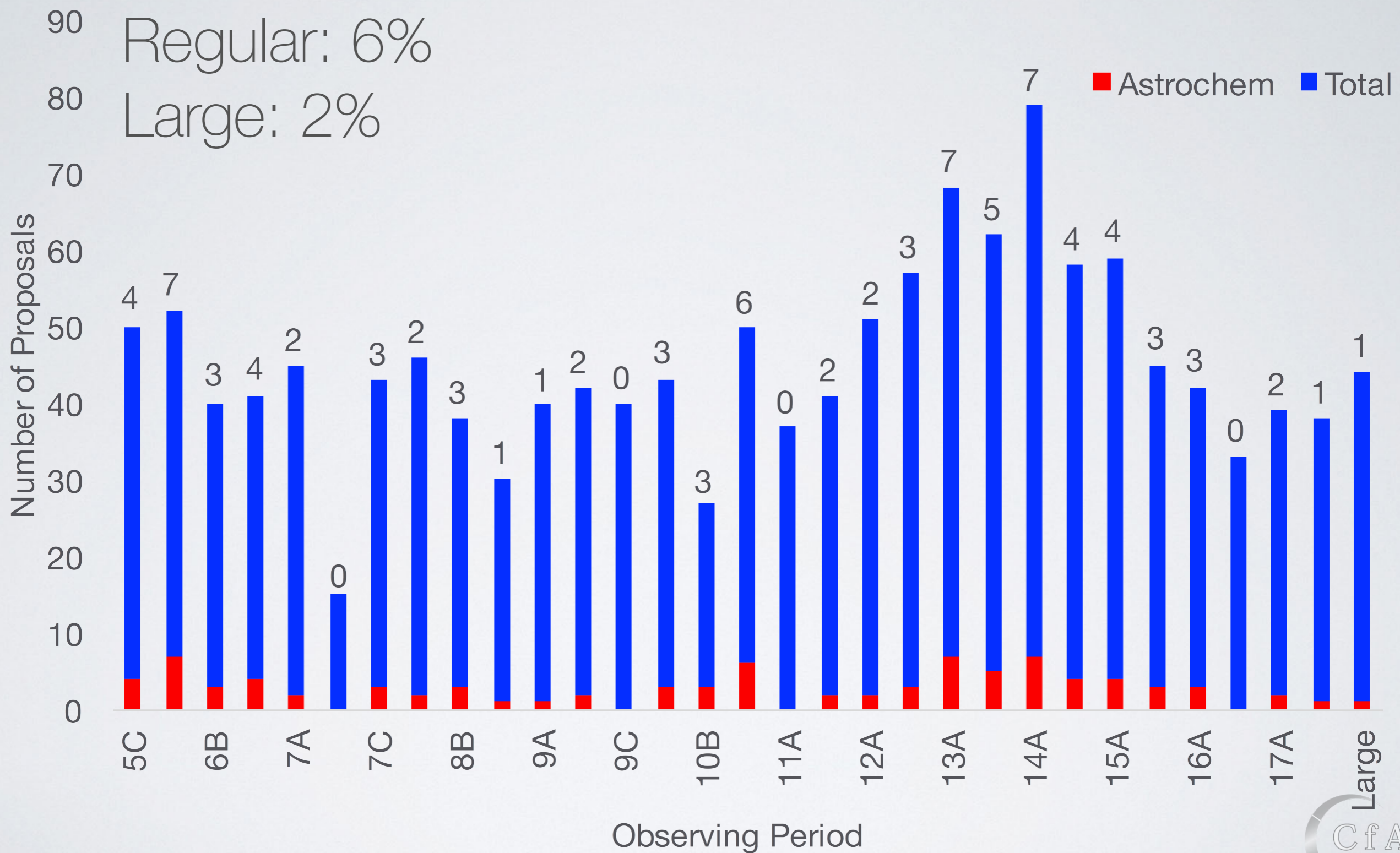
HOW DID ORGANIC CHEMISTRY IN SPACE WITH LIFE ON EARTH? LIFE ARRIVE ON EARTH?

HOW DID LIFE ARRIVE ON EARTH?

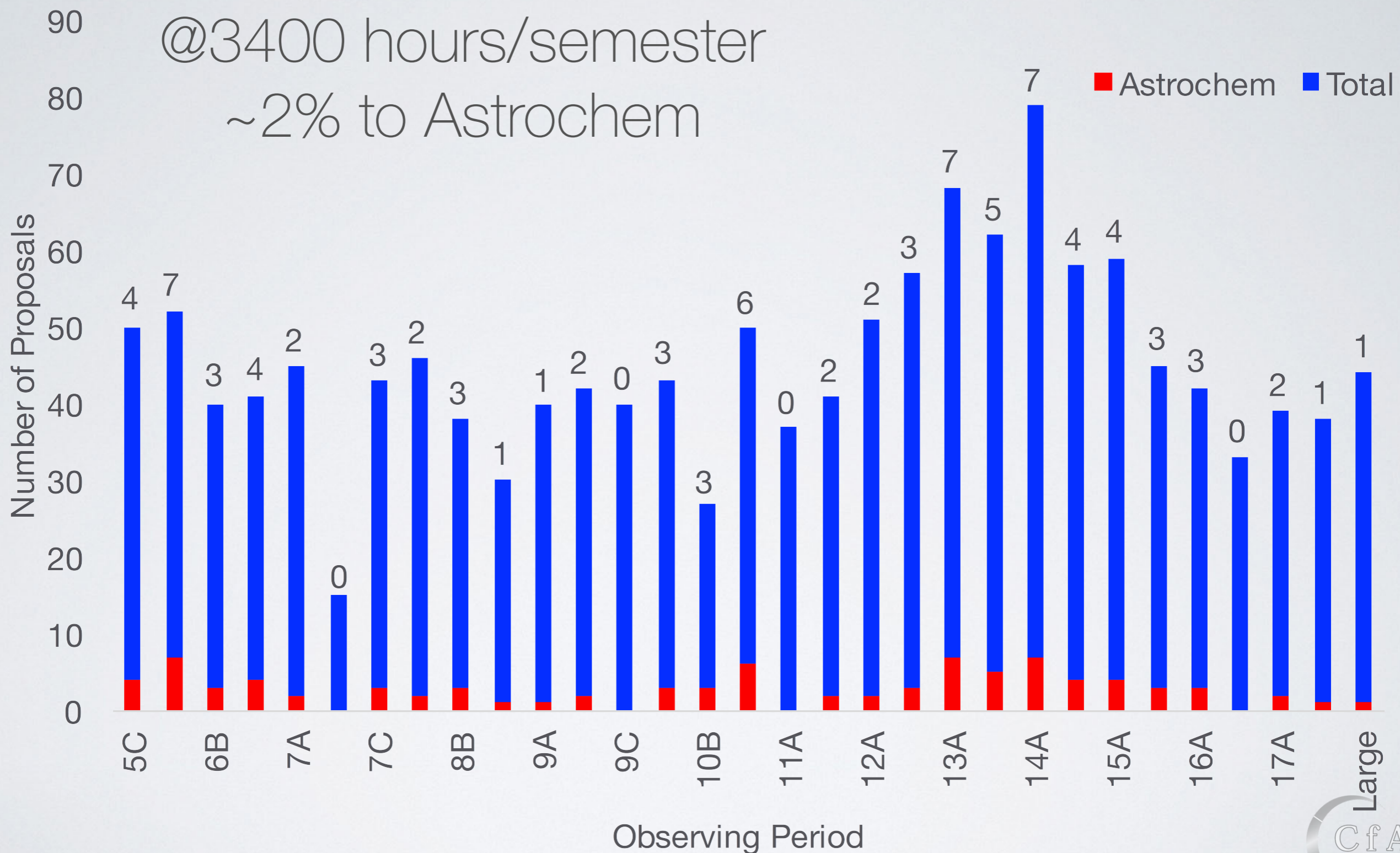
Rapid imaging of cometary molecules sets boundary conditions on solar system chemistry The GBT is the only telescope in the world which can provide the needed instantaneous sensitivity, resolution, and sky coverage

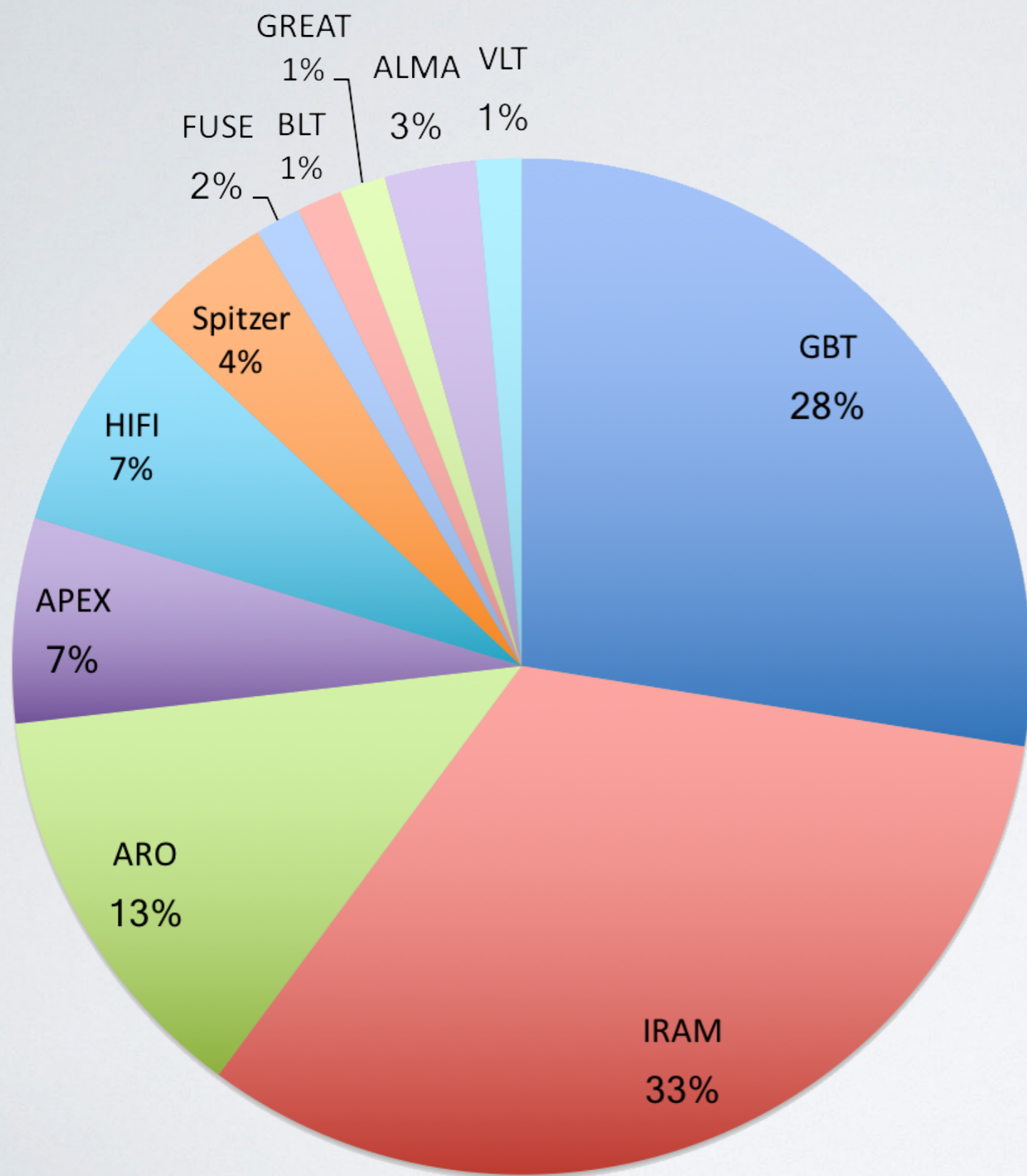
Science Category	▼
Active Galactic Nuclei	39
Astrometry/Geodesy	0
Energetic Transients and Pulsars	234
Extragalactic	263
Extragalactic Structure	139
Galactic	271
High Redshift and Source Surveys	62
Interstellar Medium	187
Normal Galaxies, Groups, and Clusters	49
Solar System	43
Solar System, Stars, Planetary Systems	77
Star Formation	70
Stellar	137

Astrochemistry is not a category in the proposal system...



@3400 hours/semester
 ~2% to Astrochem





- | | |
|-------------------------------------|-------------------------------------|
| CH ₂ CHCHO | C ₈ H ⁻ |
| CH ₃ CH ₂ CHO | CNCHO |
| CH ₂ CCHCN | HNCNH |
| c-H ₂ C ₃ O | CH ₃ CNHN |
| CH ₂ CONH ₂ | <i>E</i> -HNCHCN |
| CH ₃ C ₆ H | CH ₃ CHCH ₂ O |
| CH ₂ CNH | HC ₅ O |
| CH ₃ C ₅ N | HC ₇ O |
| C ₆ H ⁻ | |



WHERE DO WE GO FROM HERE?

What's Next?



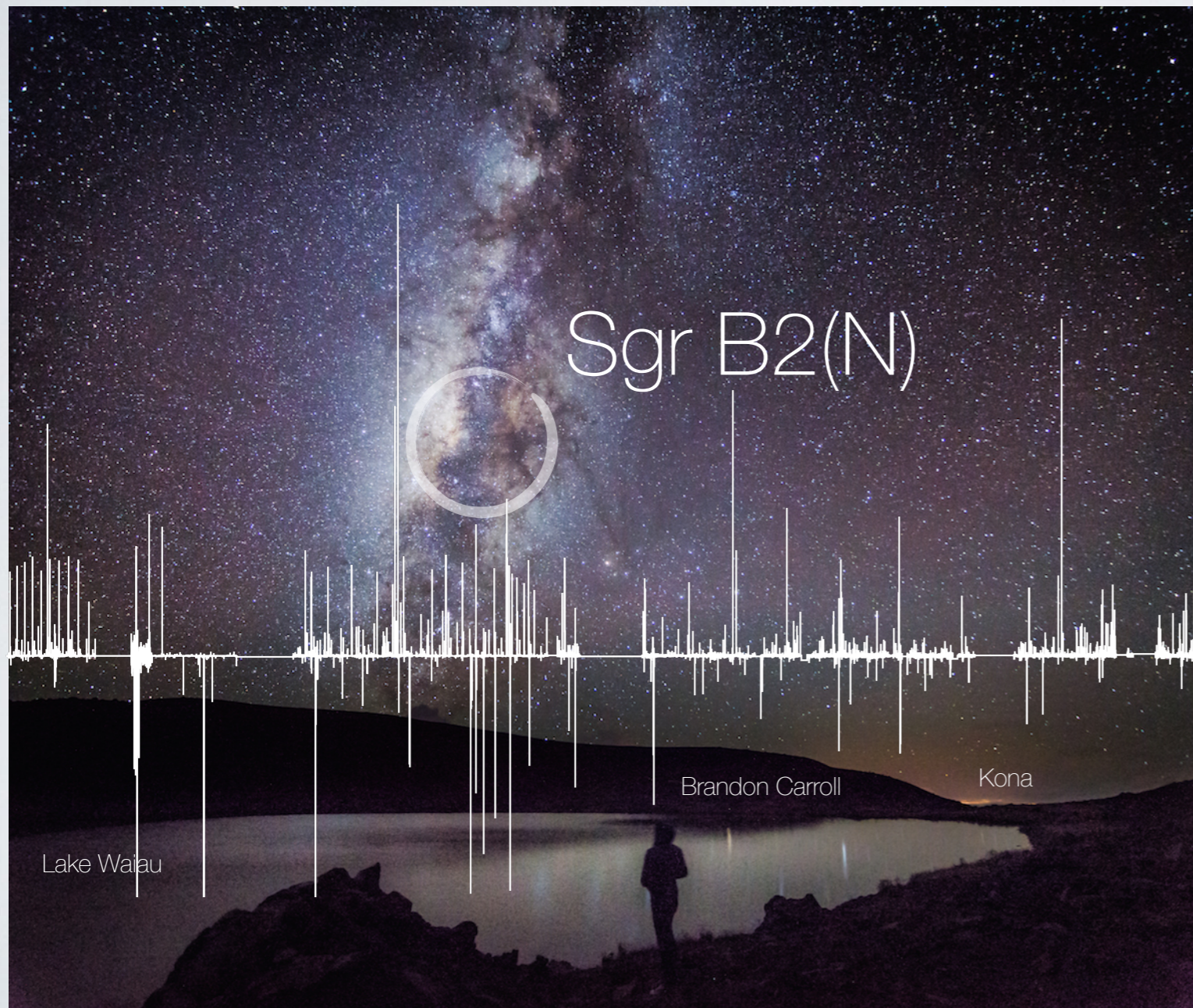
Sensitivity

Spectral Resolution / Bandwidth

~Beam Size

Unique Frequency Range

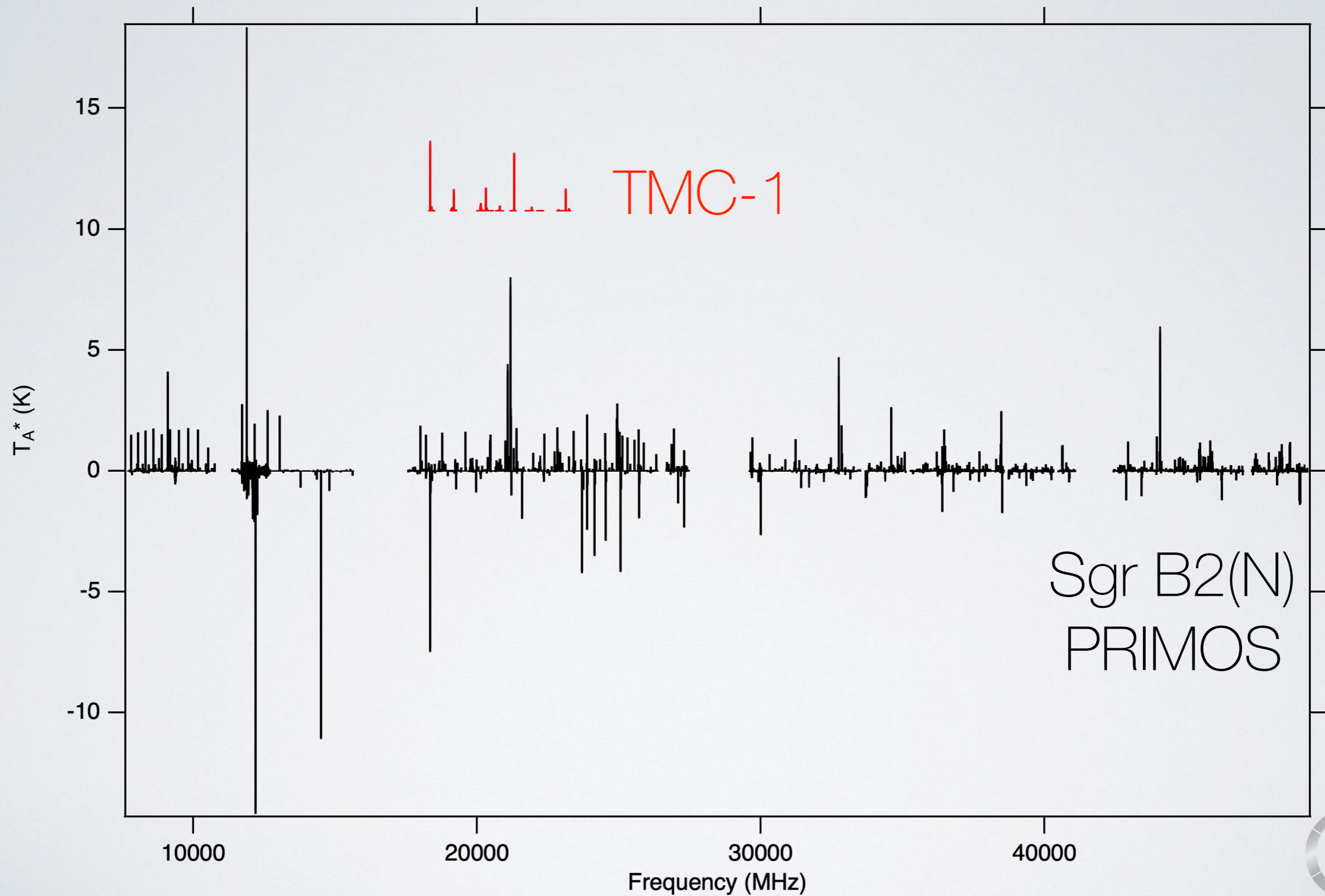
PRebiotic Interstellar MOlecular Survey

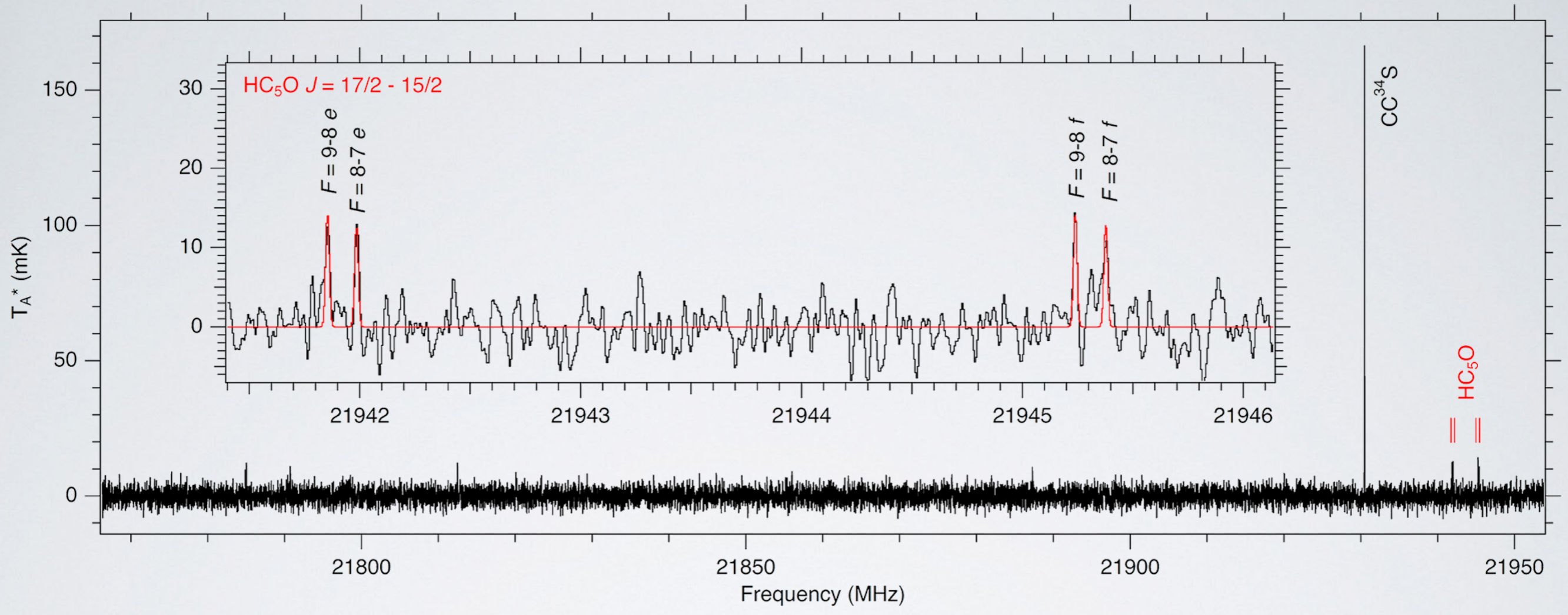


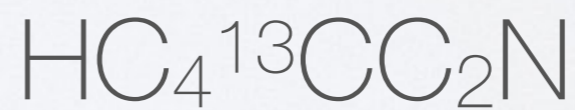
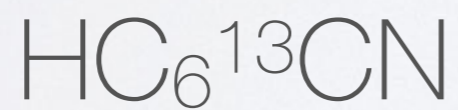
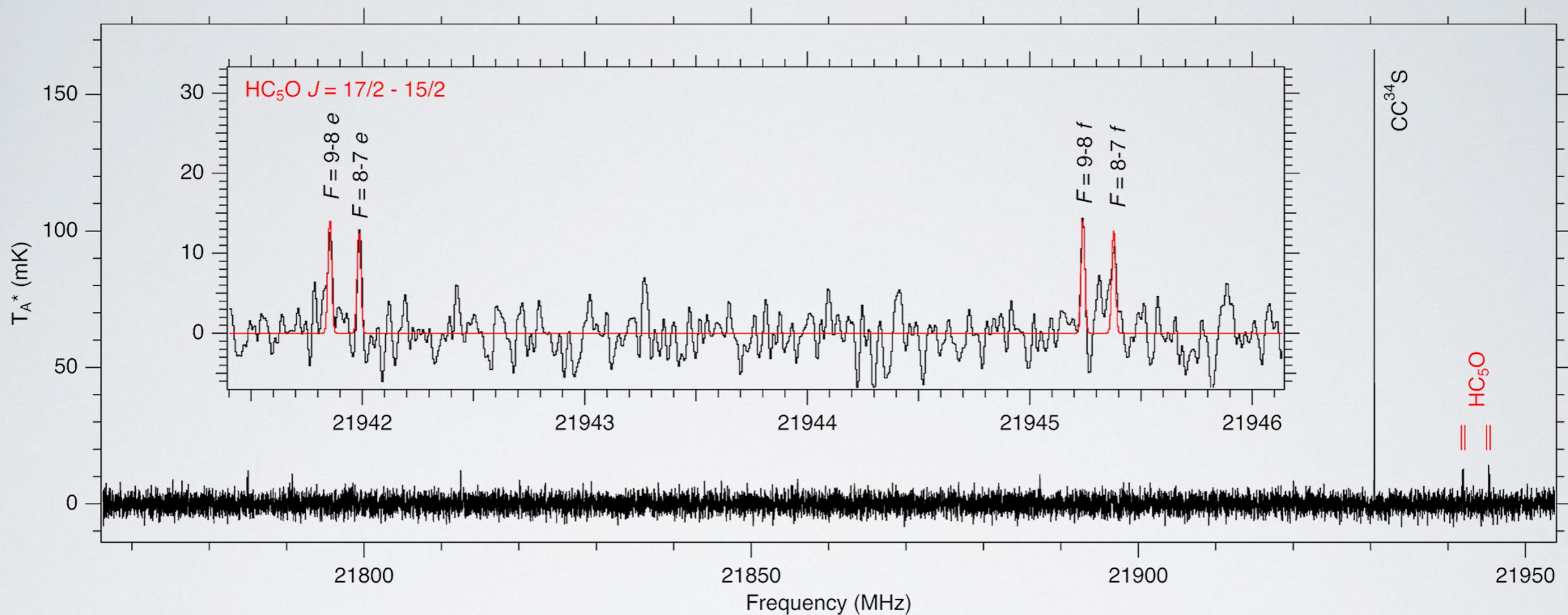
~1000 Hours w/
GBT Spectrometer

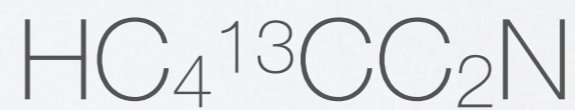
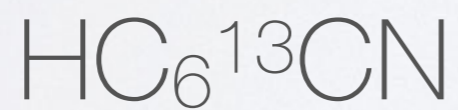
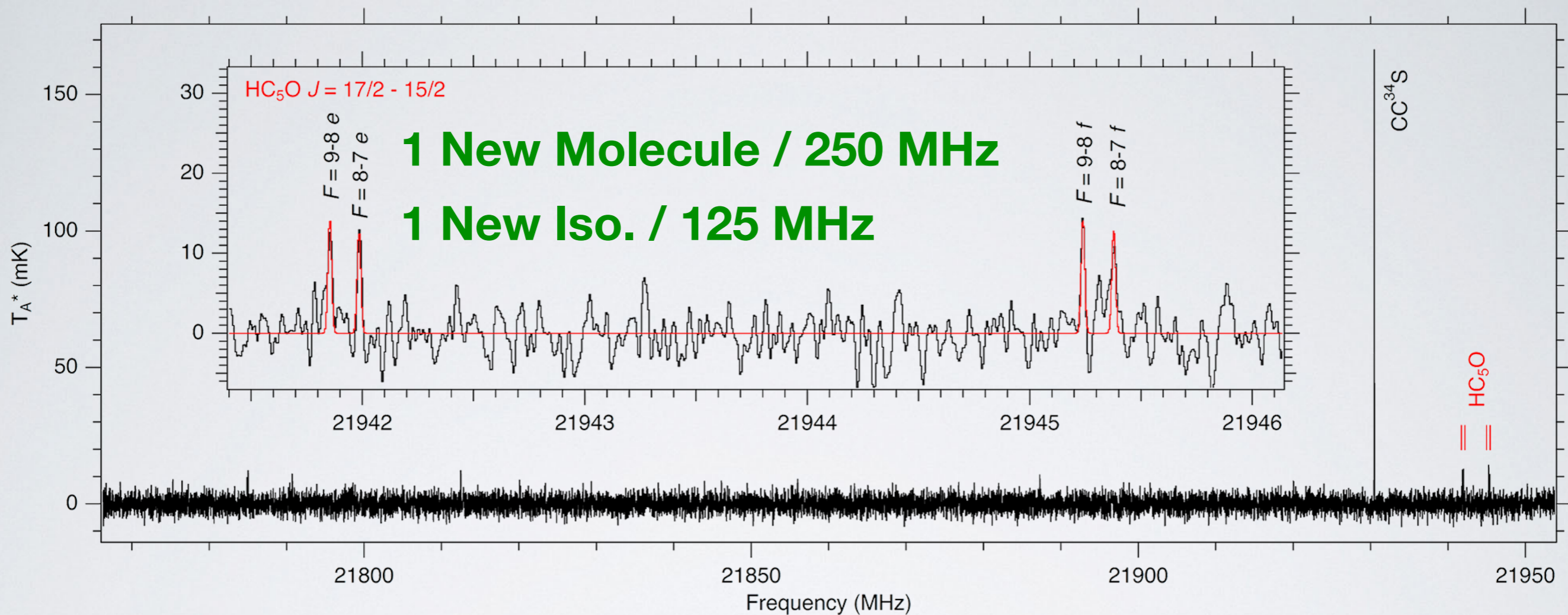
~300 Hours w/
VEGAS

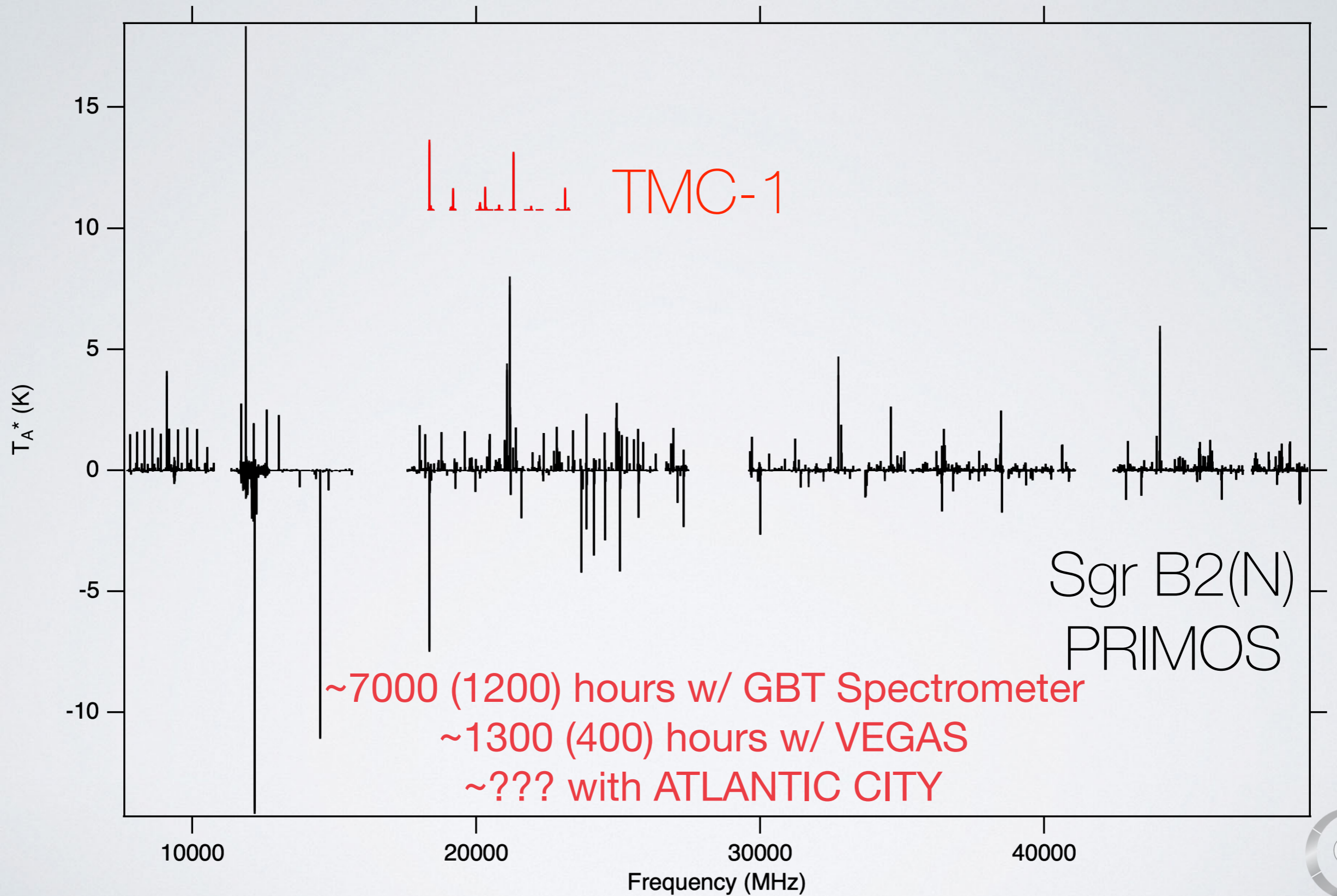
~??? Hours w/
ATLANTIC CITY











What we really need:

A shift in the way we as a community view
astrochemical research

What we really need:

A shift in the way we as a community view
astrochemical research



The primary purpose of astrochemistry is to study
chemistry

Actual '*weakness*' received on an
astrochemistry proposal:

The proposal would benefit from a clearer
description of the relation between the objectives
[...] and studies of **star formation** on Galactic
scales [...].

Could easily be recast for a star formation proposal:

The proposal would benefit from a clearer description of the relation between the objectives [...] and studies of [chemical evolution](#) on Galactic scales [...].

If both of these are not fair review criteria,
neither are

WE ARE ALL MADE OF STARS

WITHOUT STARS, WE WOULDN'T BE HERE. STARS
PRODUCE BASIC ELEMENTS, LIKE CARBON, THE
FOUNDATION OF THE COMPLEX MOLECULES WHICH
MAKE UP LIFE.

Photo/Video | Photo Album

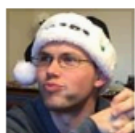


What's on your mind?

Friends

Post

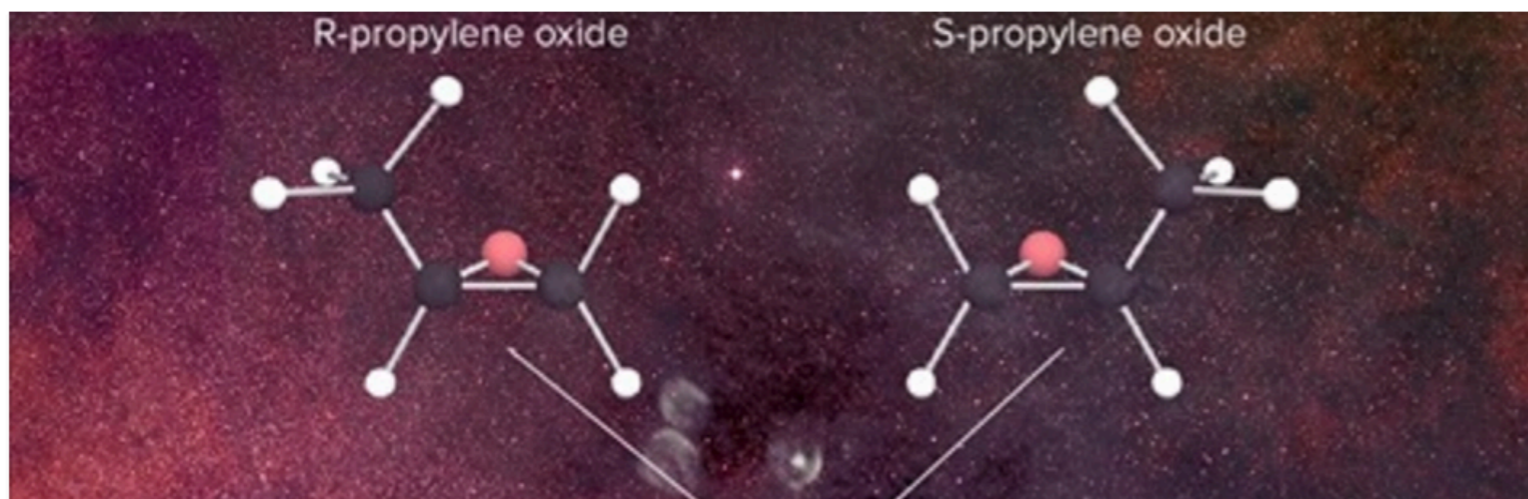
Geoffrey A. Blake commented on your post.



Brett McGuire with Ryan Loomis and 2 others.

11 hrs · Friends

Brandon and I did a thing with Geoff, Ryan, Ian, Phil, and Tony!!!



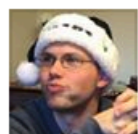
18 Annual Midsummer's P... on June 18

TODAY'S GAMES

TRENDING

- Alexander Skarsgård:** Actor Says 'True Blood' Co-Star Ryan Kwanten Is a 'Good Kisser'
- Donald Glover:** Actor in Talks to Join Cast of Upcoming 'Spider-Man: Homecoming' Film, Report Says
- Mara Wilson:** Actress Says She Has 'Embraced the Bi/Queer Label Lately'
- Yasmine Bleeth:** Former 'Baywatch' Actress Celebrates 48th Birthday on June 14
- Jessie Graff:** Stuntwoman Competes in 'American Ninja Warrior' While Dressed as Wonder Woman
- Game of Thrones:** HBO Releases Titles, No Descriptions for Final 2 Episodes of Season 6
- Sniper Elite 4:** Tactical Shooter's Release Delayed Until February 2017
- Chiral Molecules:** Scientists Discover Compounds Resembling Human Hands Outside Solar System

Photo/Video | Photo Album

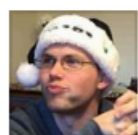


What's on your mind?

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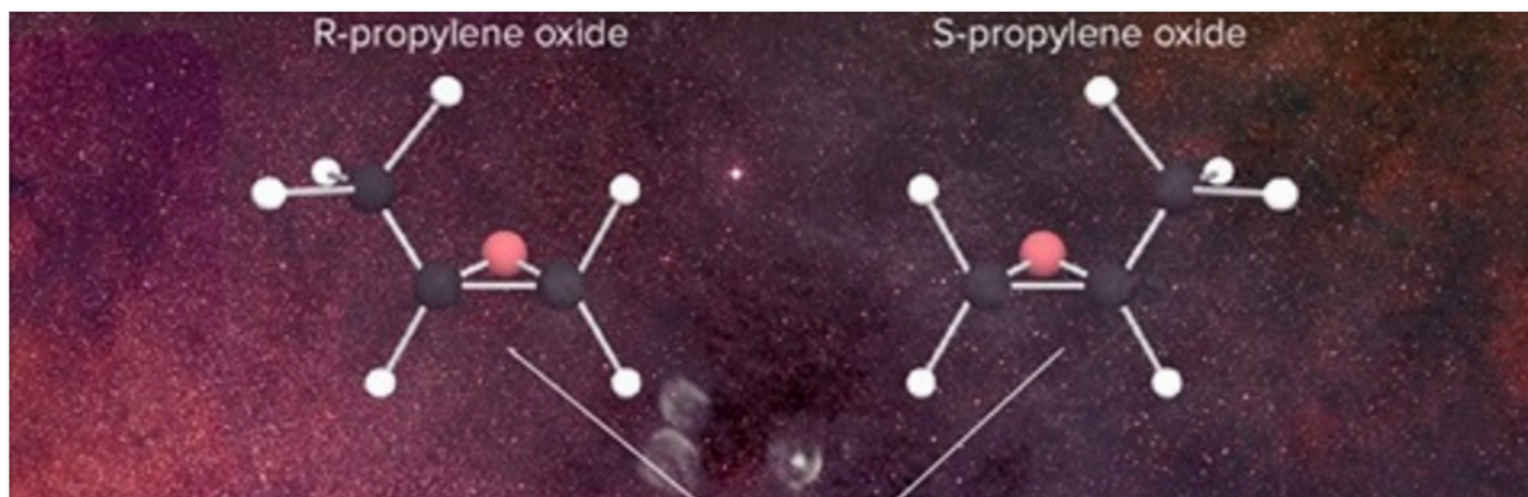
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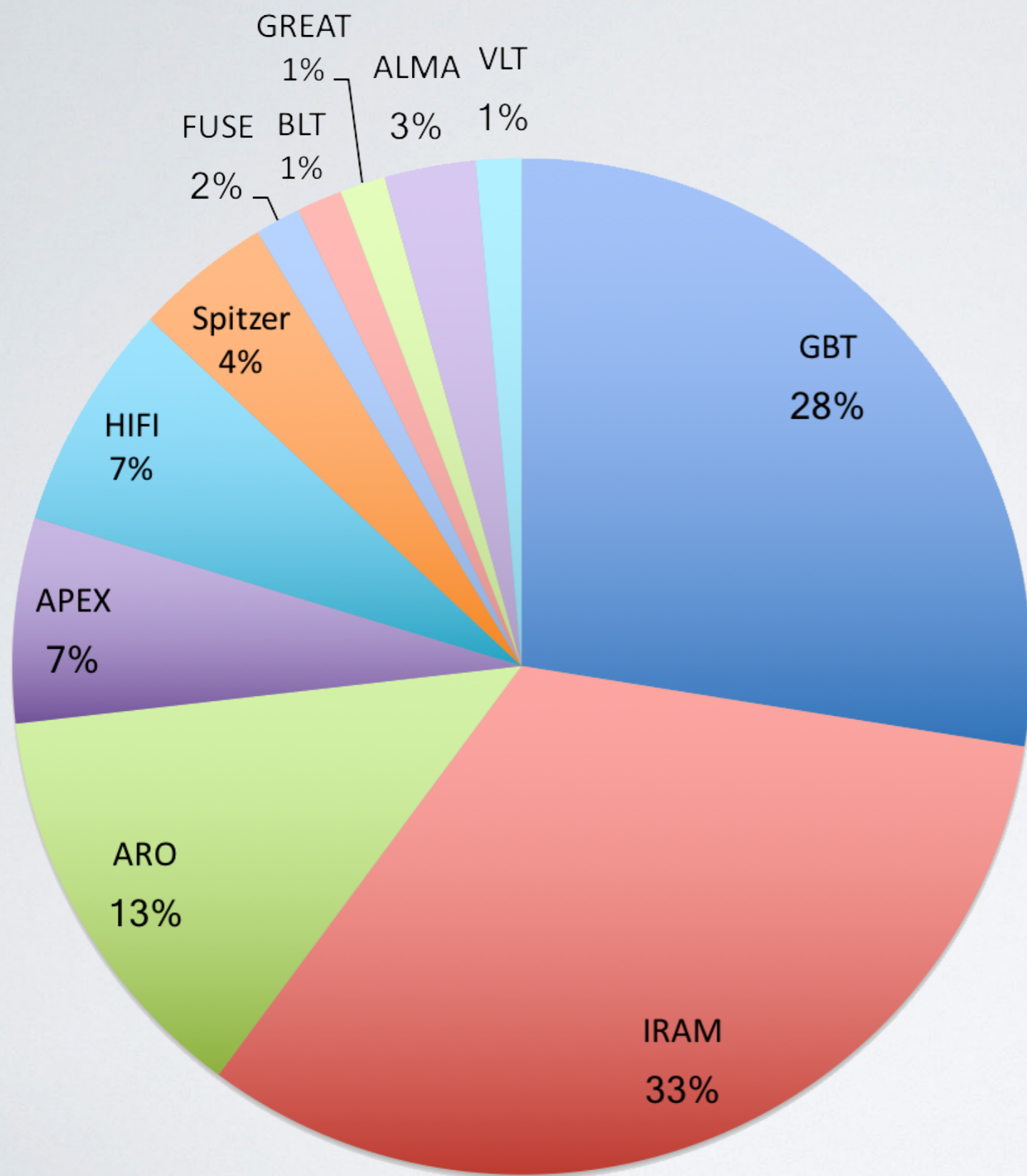


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| CH ₂ CNH | HC ₅ O |
| CH ₃ C ₅ N | HC ₇ O |
| C ₆ H ⁻ | |

GBT facilities are, and always have been, at the forefront
of astrochemical discovery

‘Minor’ technical upgrades will enable significant new
science

The upgrade we really need is political, not technical