

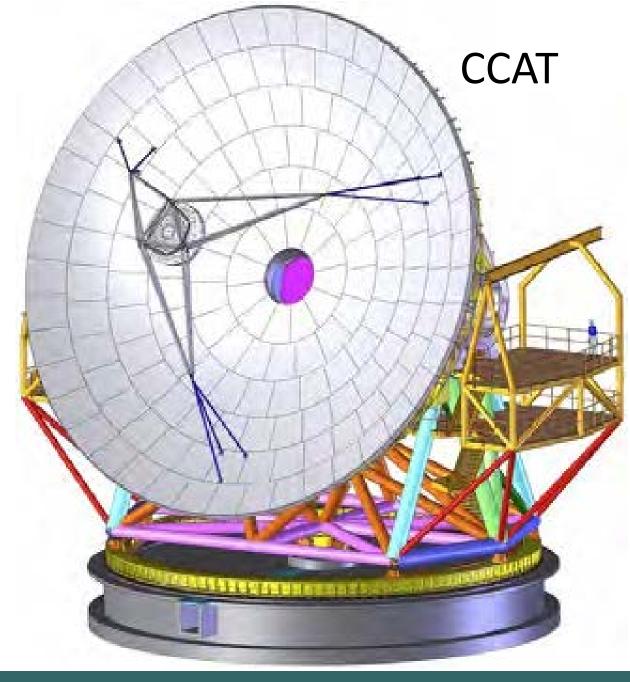
Measurement and Control of Large Telescope Surfaces CCAT example

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> D. Woody Caltech

http://www.ccatobservatory.org







25 m dia telescope on Cerro Chajnantor, 5612 m

Observe up to 350 μ m => 12 μ m RMS

CFRP primary support structure

162 compound segments, CFRP subframe with 16 machined Aluminum tiles

Active surface

In an enclosure, never sees the sun or high winds

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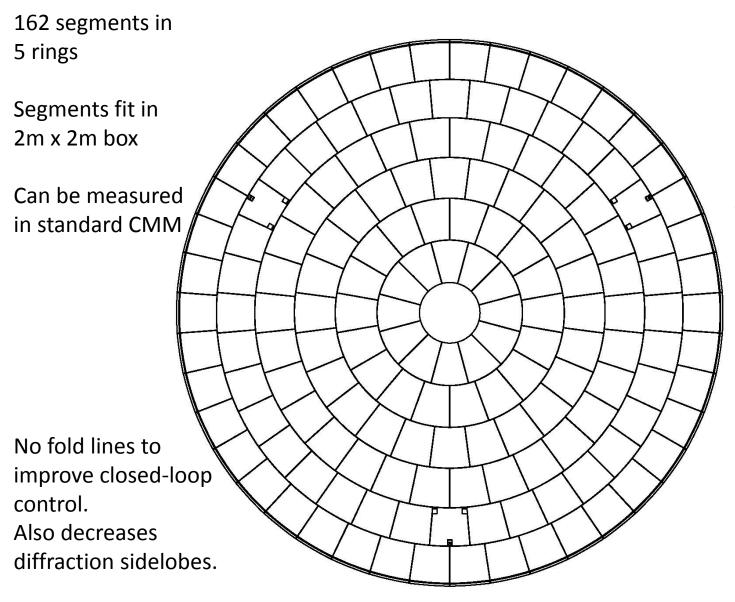
Measurement and Control of Large Telescopes

Segments are critical



- Only 9 μm RMS allocated to primary
- Segments < 4 μ m
- Want largest possible segment to minimize parts count
- Segments also serve as the reference for any closed loop control using edge sensors (like ULE segments on large optical telescopes)
- => 2 m x 2 m compound segments, CFRP subframe and machined Aluminum tiles

Segment configuration

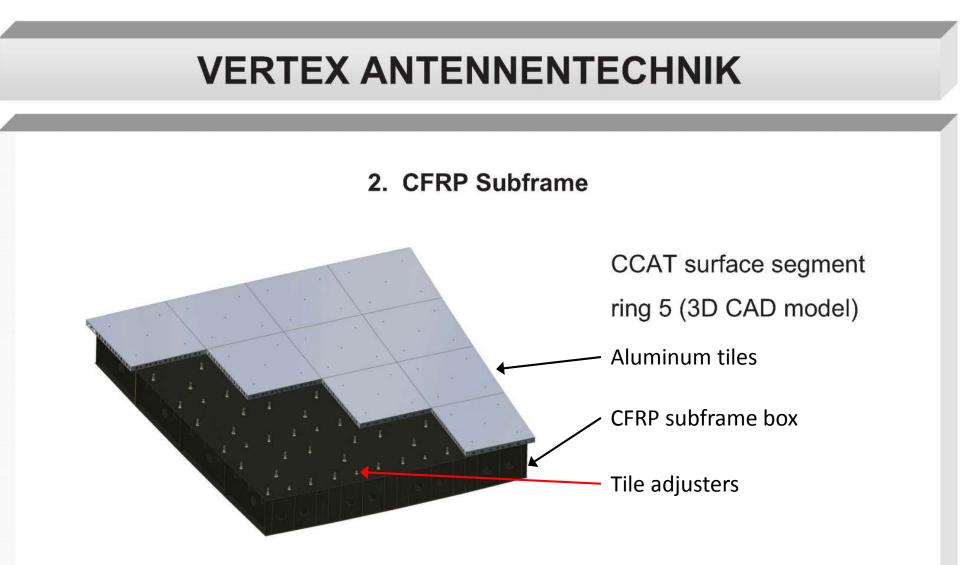




Each segment supported by three actuators

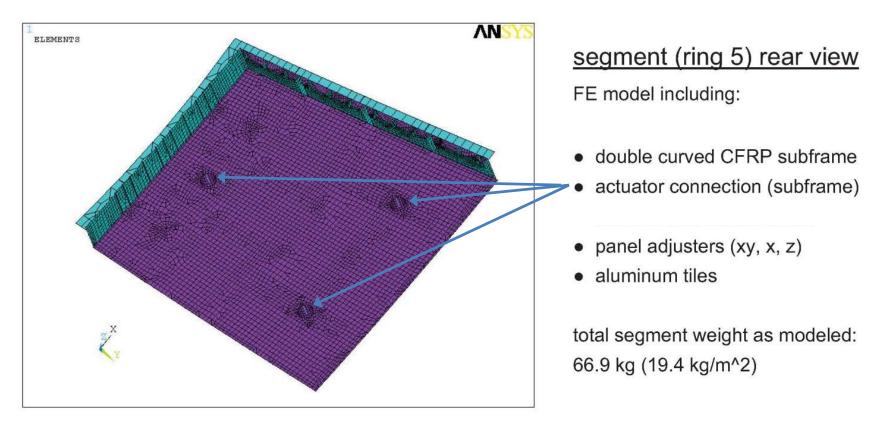
Goal of open loop control for gravity distortions

Plan for edge sensors and closed control if necessary



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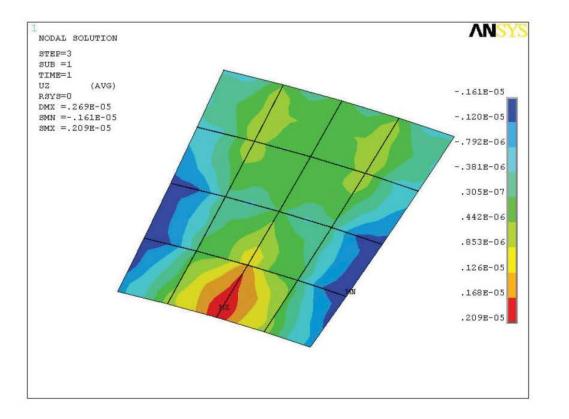
2. CFRP Subframe - FE Model



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2. CFRP Subframe - Exemplary Displacement Plots (K63712 prepregs)



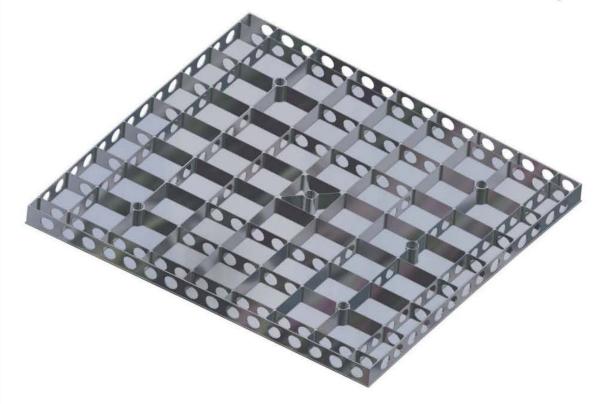
LC3: gravity along zq (actuator axis) normal-to-surface best-fit dev. [m]

rms surface error = $0.7 \mu m (001110)$

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4. Aluminum Panels (tiles)



3D CAD model of largest panel of segment ring 5

Panel areal weight ≈ 6.7 kg/m²

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5. Surface Error Budget (K63712 fiber prepregs)

Error Source	Surface Error	Source
gravity, 20° EL to 90° EL	0.8 µm	FEA, active
ambient temp. change 20 K	0.7 µm	FEA, passive
temp gradient dDT/dz = 1 K	0.3 µm	FEA, passive
wind 1 m/s	< 0.1 µm	FEA, passive
flexure heating dT= 1 K	< 0.1 µm	FEA, passive
1 mm actuator stroke	0.8 µm	FEA, active
panels	3.0 μm	best manufactured to date
adjustment	2.0 µm	estimate
others	1.0 µm	estimate
Total (RSS)	4.0 µm	Spec.(CCAT) : rms!≤ 3.8 µm

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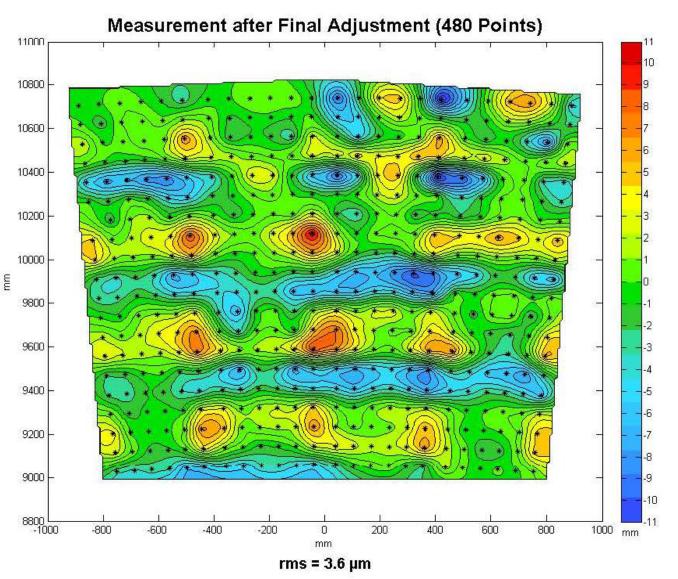


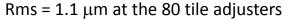
Telescopes



Weight 60 kg

Ring 5 segment measurement



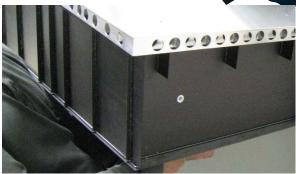


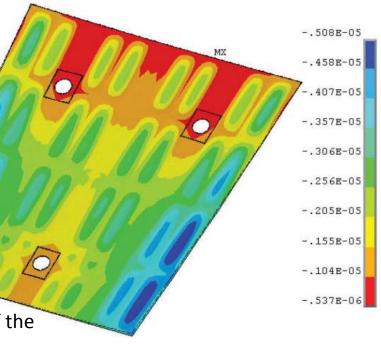




Closed-loop control of segment actuators

- Difficulties
 - Full wavefront measurements will be infrequent
 - CFRP sub-frames are not as stiff nor as stable as ULE glass
 - Segment edges are not representative of the whole segment
- Edge sensors need good long term stability
- Measuring more than just dihedral angle and piston displacements improves the surface control

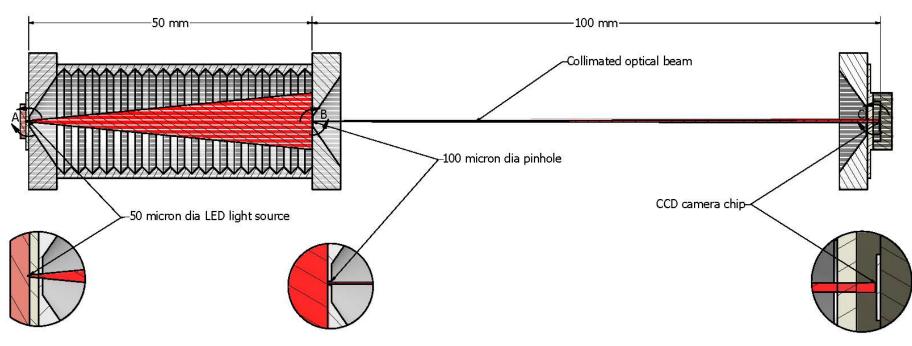




Gravity deformation of the subframe backs side

Imaging Displacement Sensor (IDS)

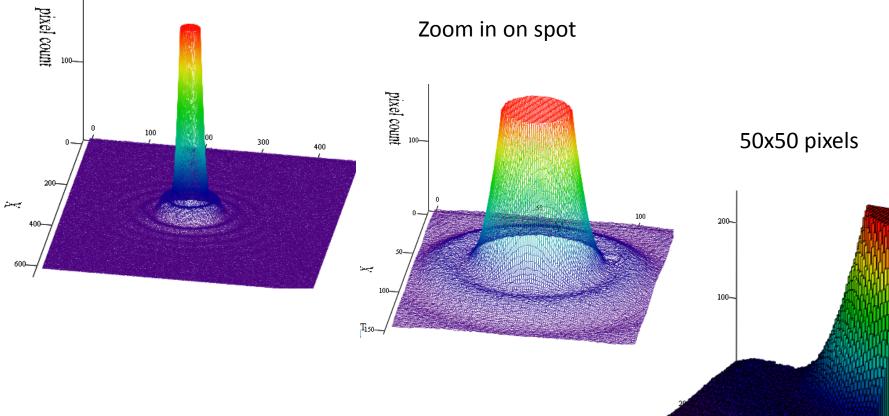




- Measures lateral displacements at a distance
- Exploit the precision lithography used for CCD or CMOS camera chips
- One device measures displacements in X and Y
- Inexpensive components, LED and standard monochrome VGA CCD
- Very simple and stable
- Single frame image centroiding error <1/100th of a pixel
 - Averaging multiple frames < 1/1000th of a pixel







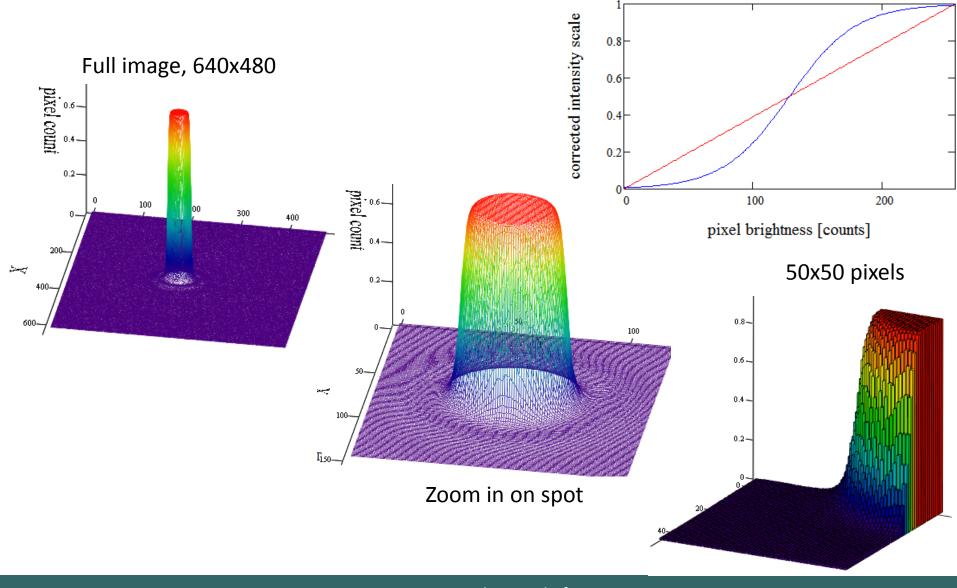
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Full image, 640x480

Apply gain correction to emphasize middle range



Gain correction to emphasize mid-range



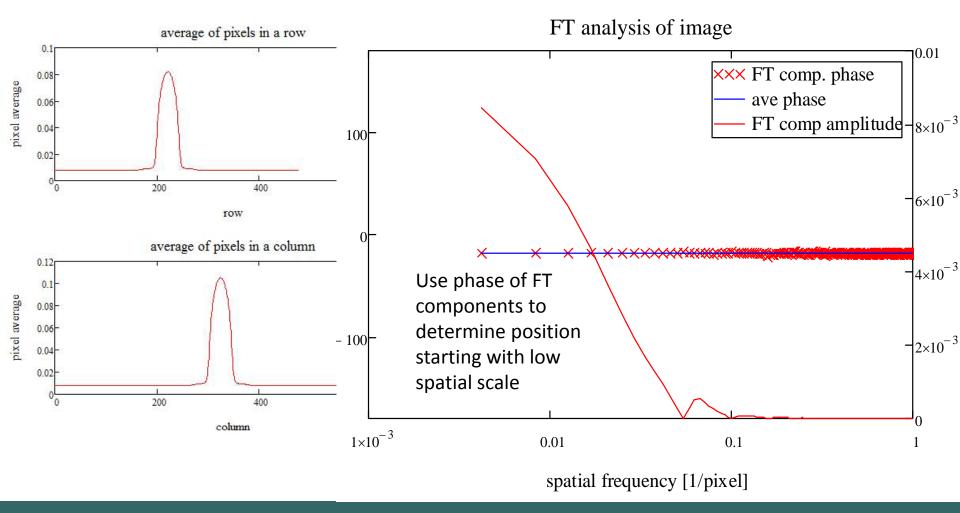
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Measurement and Control of Large Telescopes

Centroid algorithm

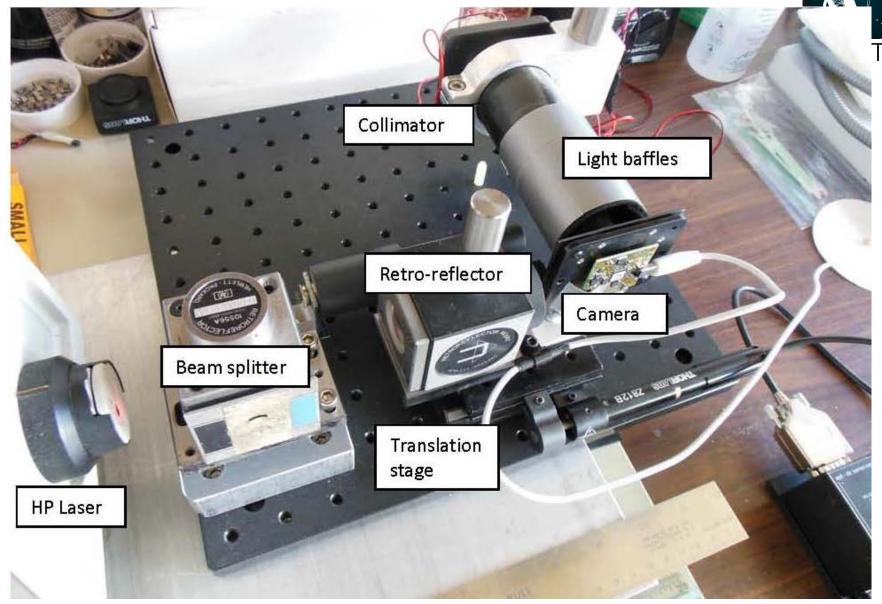
Calculate row and column averages and Fourier Transform to determine the centroid.

Use the phase of the lowest FT components to resolve ambiguity of the phase for the higher components. Find position from amplitude weighted averages of the phases.

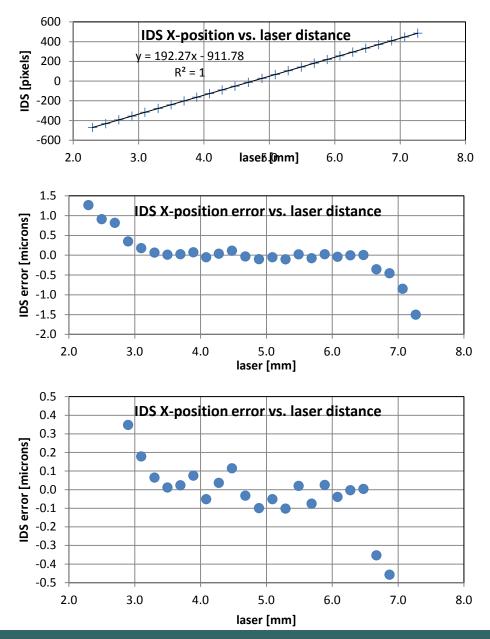




Linearity measurement setup



X-position linearity





Nominal 5.2 mm pixel gives 192.30 pixels/mm

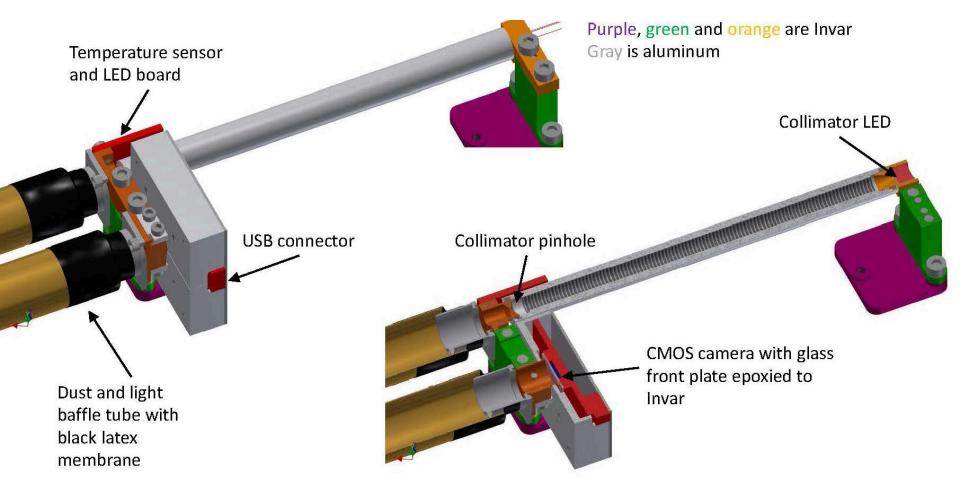
No calibration required

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IDS design model

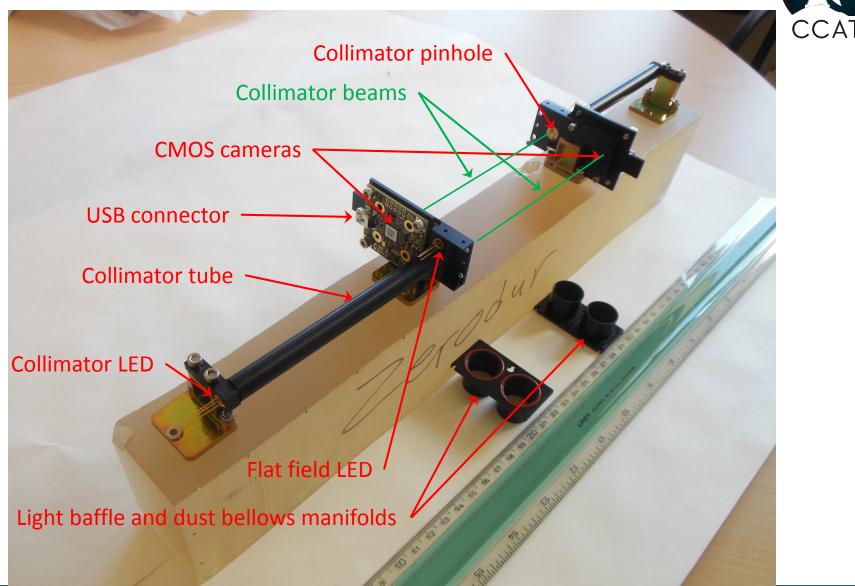




Telescopes

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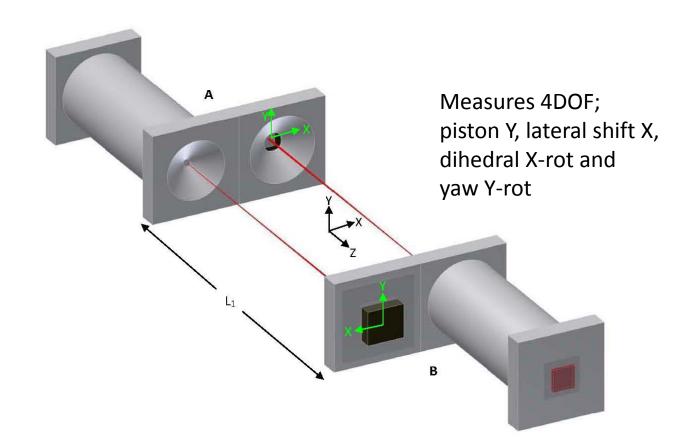
Pair of IDS modules



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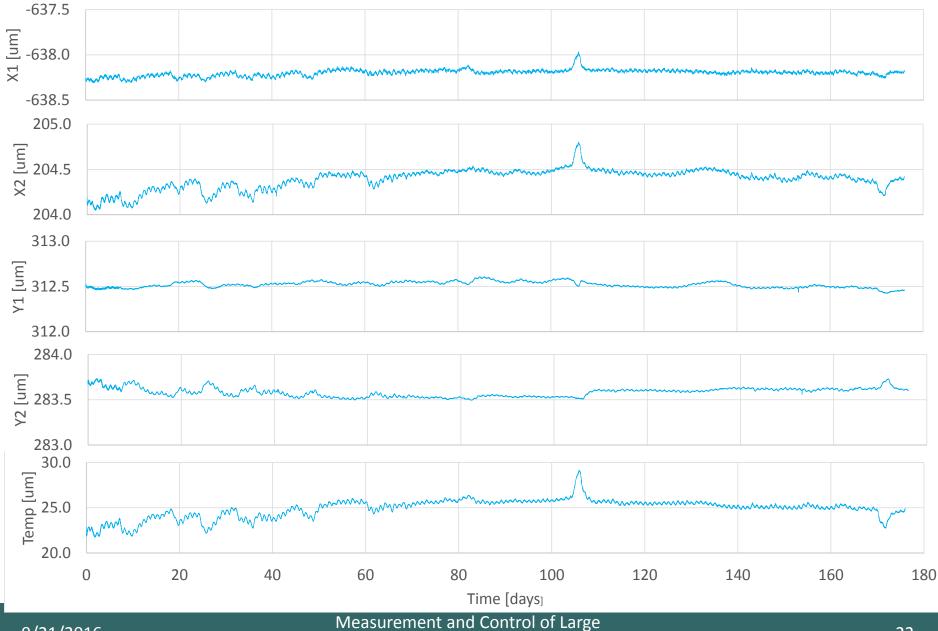
Pair of IDS modules





Long term stability over 6 months



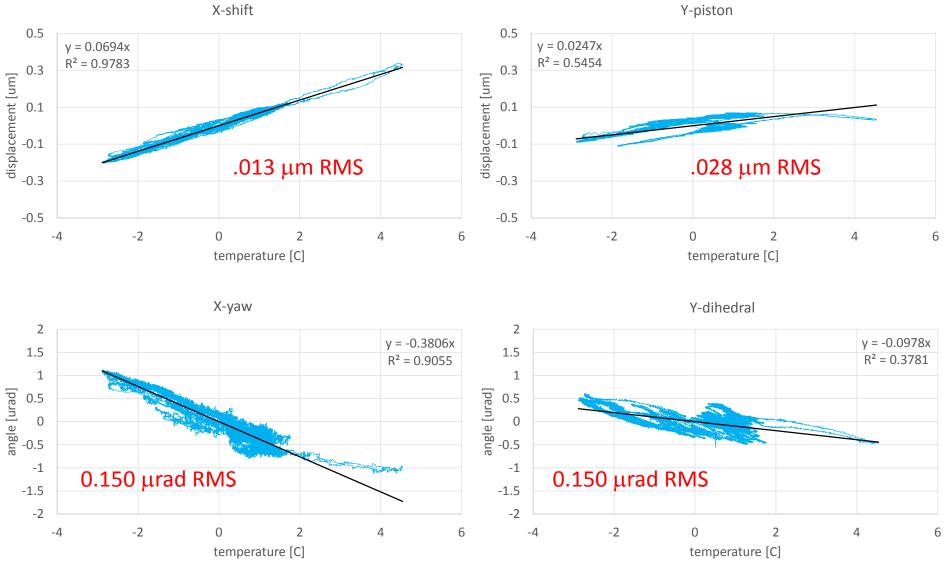


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Telescopes

Temperature dependence



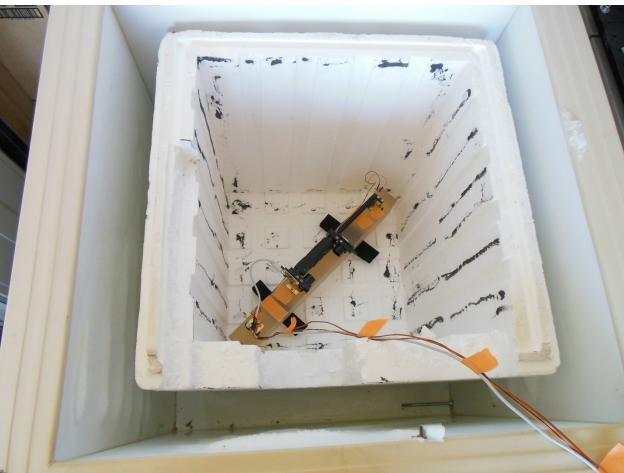


Measurement and Control of Large Telescopes

IDS temperature test

Put IDS modules mounted on Zerodur block inside Styrofoam box inside a small freezer and cycle between +20 C and -20 C.

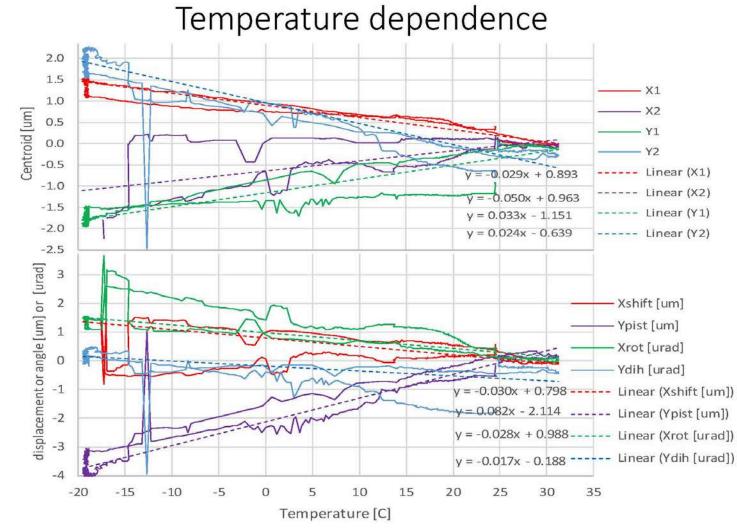




Measured <0.08 μm/C

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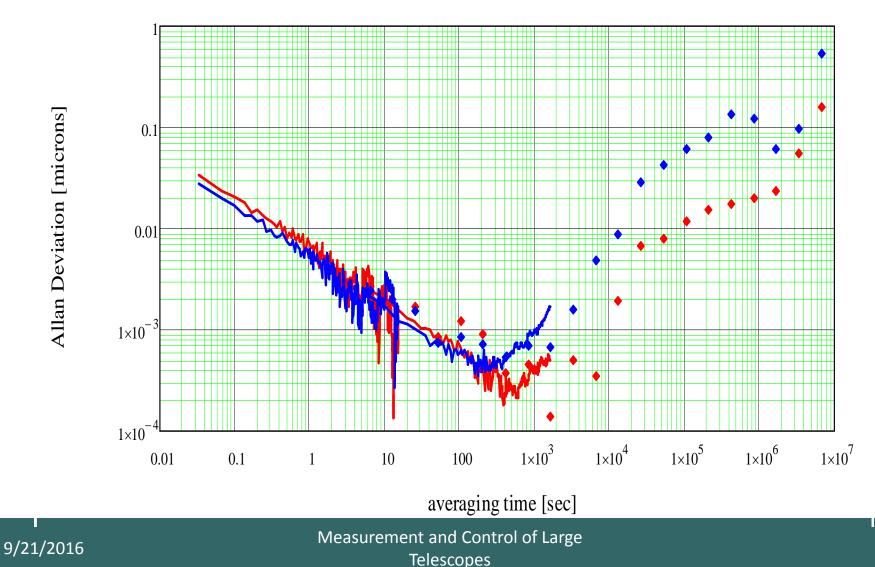
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Allan Deviation of test bed IDS



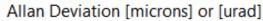


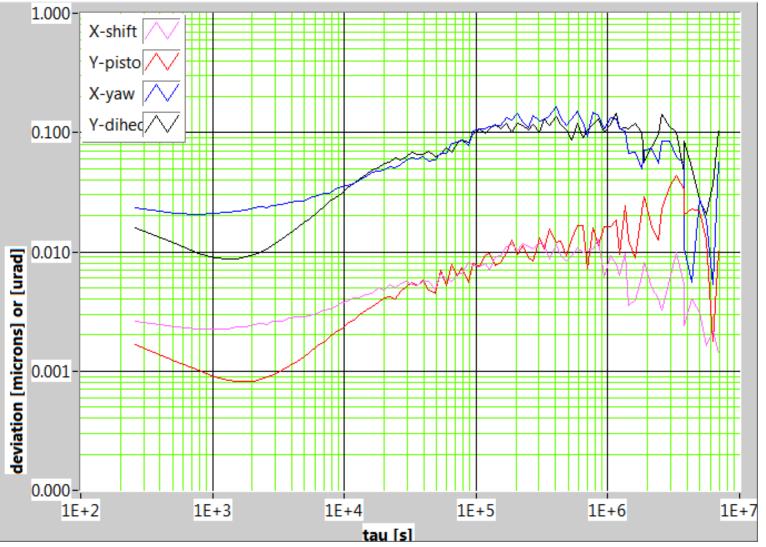


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Long term stability of prototype IDS pair

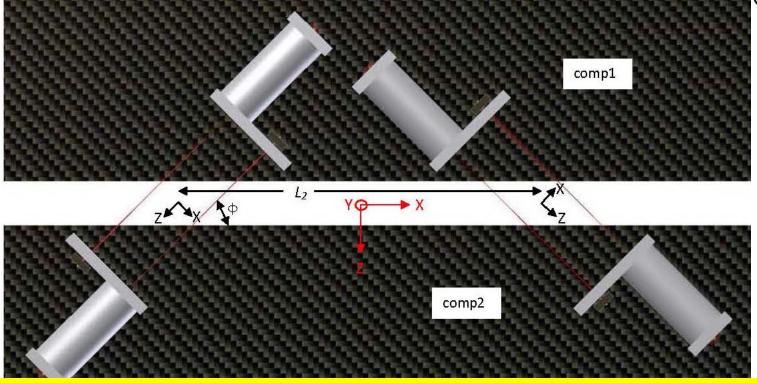






Measure 6 Degrees of Freedom

- Measuring 6DoF greatly improves the surface control performance
- Two pairs of IDS modules (4 collimators and CCDs) does this very efficiently



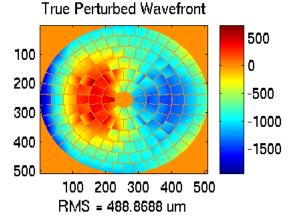
 IDSs easily meet CCAT linearity, noise and stability requirements Deviations from linear <0.1 μm over ±1 mm: requirement < 1 μm Noise <0.05 μm in one 0.1 sec frame, <0.1 in 1 sec: requirement < 0.1 μm in 1 sec Temperature dependence <0.01 μm/C with linear temperature correction: requirement < 0.01 μm/C Measures 6 DOF

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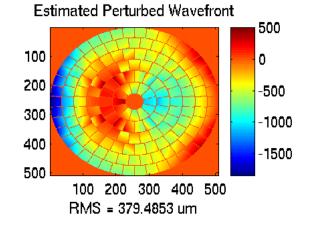
Surface control

CCAT

FEM gravity deformation at horizon

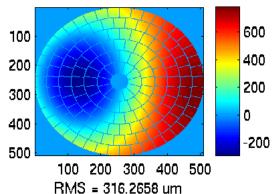


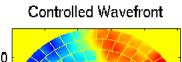
Estimated WFE state 0.5 μm RMS sensor noise

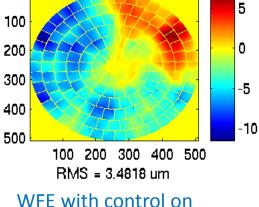


Effect of shifted secondary position

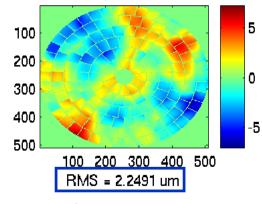
Diff WF: Estimated-True





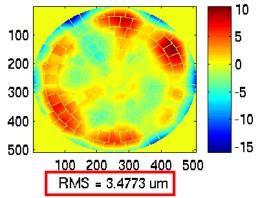


Ctrl-ed WF w/ Global Tip/Tilt/Focus/Pist Removed



Refocus and repoint

Add 1 μm uniform cupping





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Measurement and Control of Large

Telescopes