



Metrology and Control of Large Telescopes

Green Bank, WV | 19 - 24 September 2016

M1/M2 Ray Tracer

for High-Speed Mirror Metrology
in the E-ELT



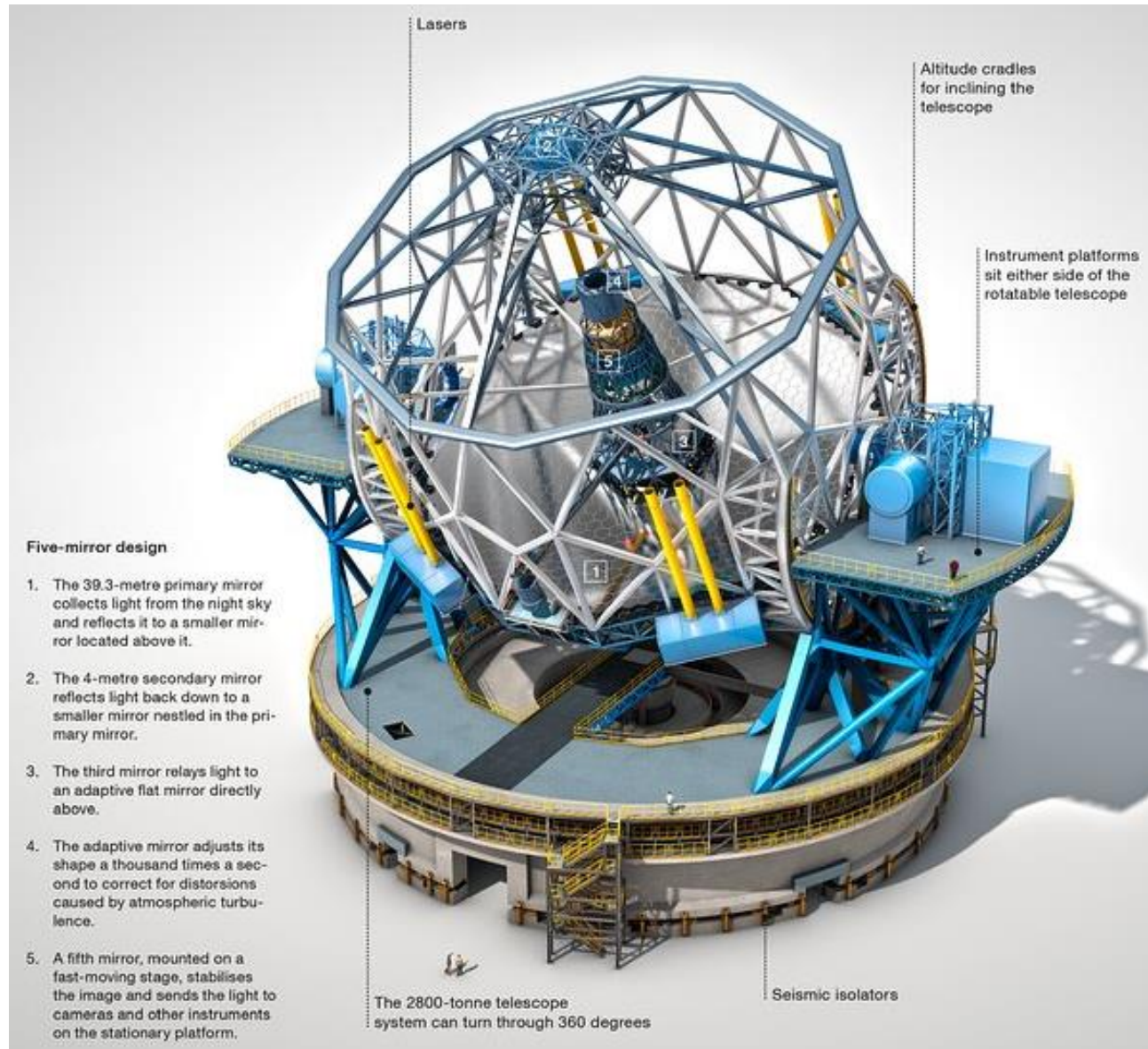
Ron Holzlöhner, 21 Sep 2016

European Southern Observatory (ESO)



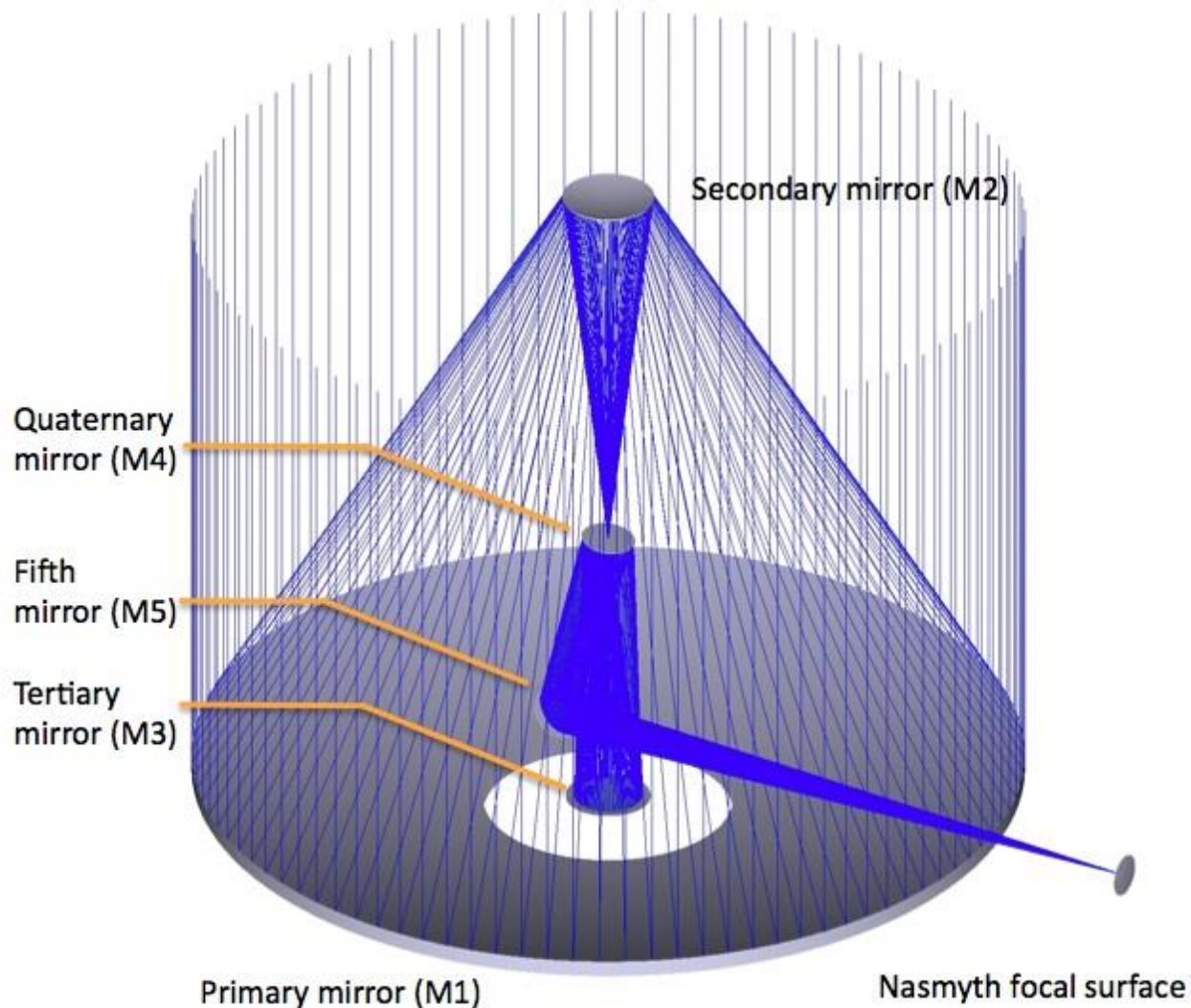
The E-ELT: 39m visible+IR Telescope

- ESO: Intergovernmental Organization, 15 member states
- EELT: 39 m visible+IR project
- Construction phase, first light 2024
- Segmented primary (~800 hexagons)
- 10 arcmin FoV
- Diffraction-limited through adaptive optics + LGS



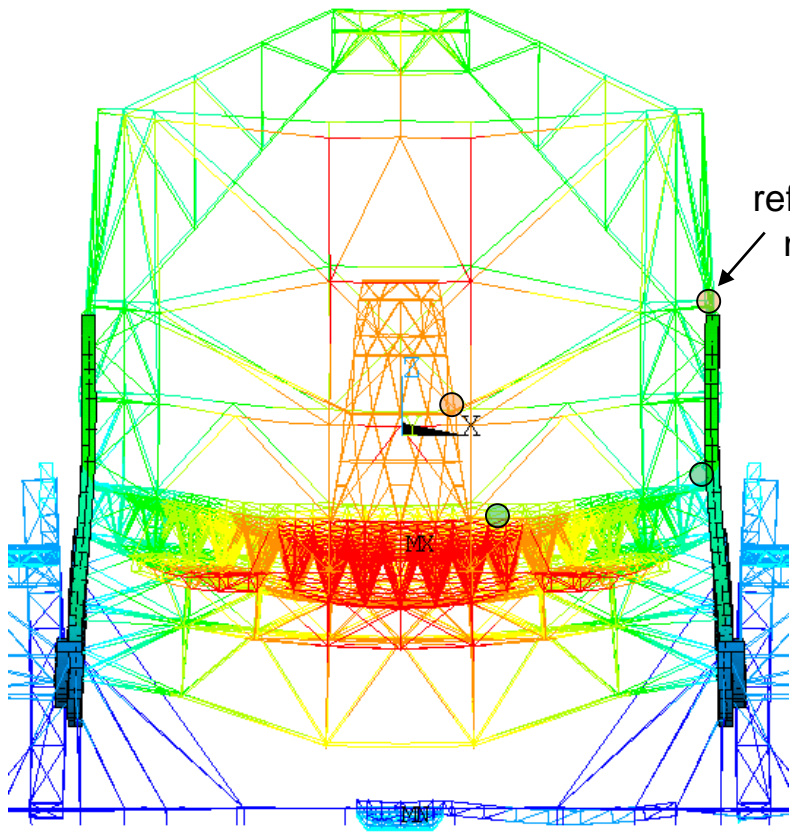
E-ELT Optical Design

- 3-mirror anastigmat
- M1–M3 powered, M4: adaptive optics, M5: fast tip/tilt
- Intermediate focus (hole in M4)
- Active optics baseline: 3 WFS around focal plane
- No online mirror position metrology in baseline

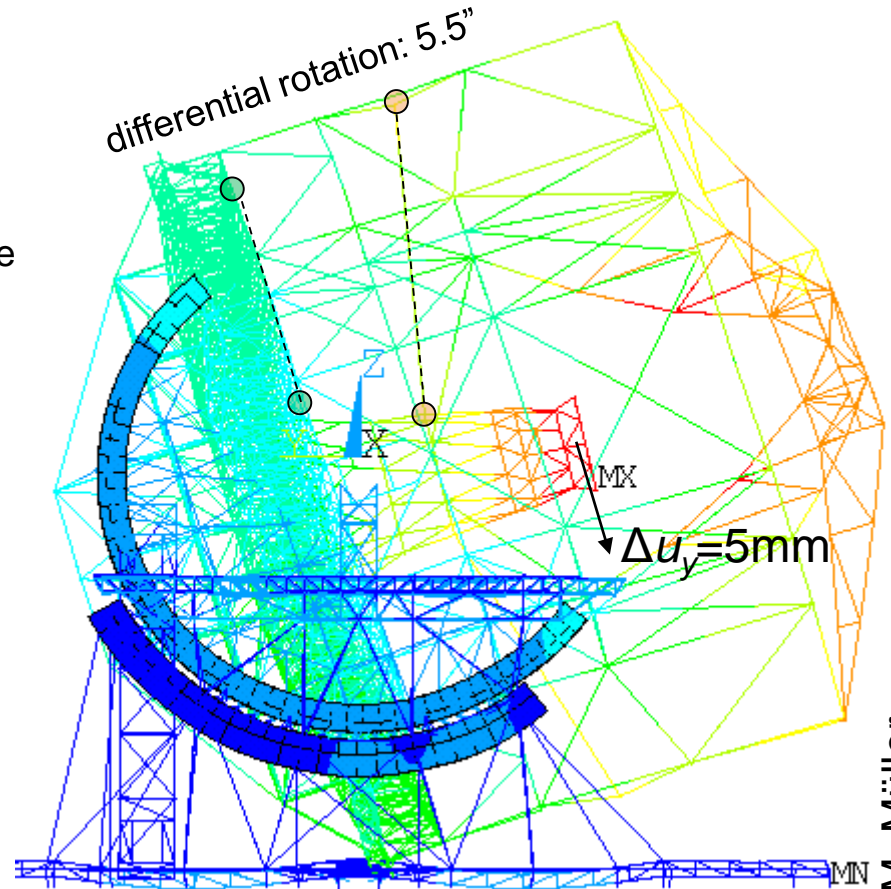


Structure Gravity Flexure

Absolute gravity flexures



zenith, red: 11 mm total xyz



$Rot_x 70^\circ$, red: 15 mm

How to do precision mirror metrology with such flexures?

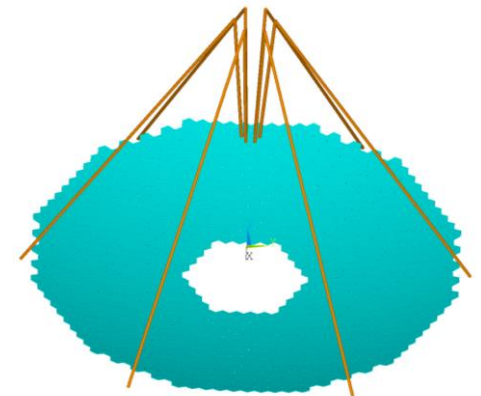
Global Mirror Metrology

- What is global mirror metrology useful for?
 - Global mirror position + figure: AIT, Commissioning
 - Applications: Blind pointing
 - Scalloping, plate scale, wind shake metrology
- Sub-micron/arcsecond accuracy over tens of meters requires lasers
- Candidate: *Etalon Multiline* (24-channel laser interferometer) extended to 30+ meters. Challenges:
 - Measurements compromised by main structure flexure
 - Local turbulence → opt. path length jitter, beam wander
 - Light pollution of science and/or WFS

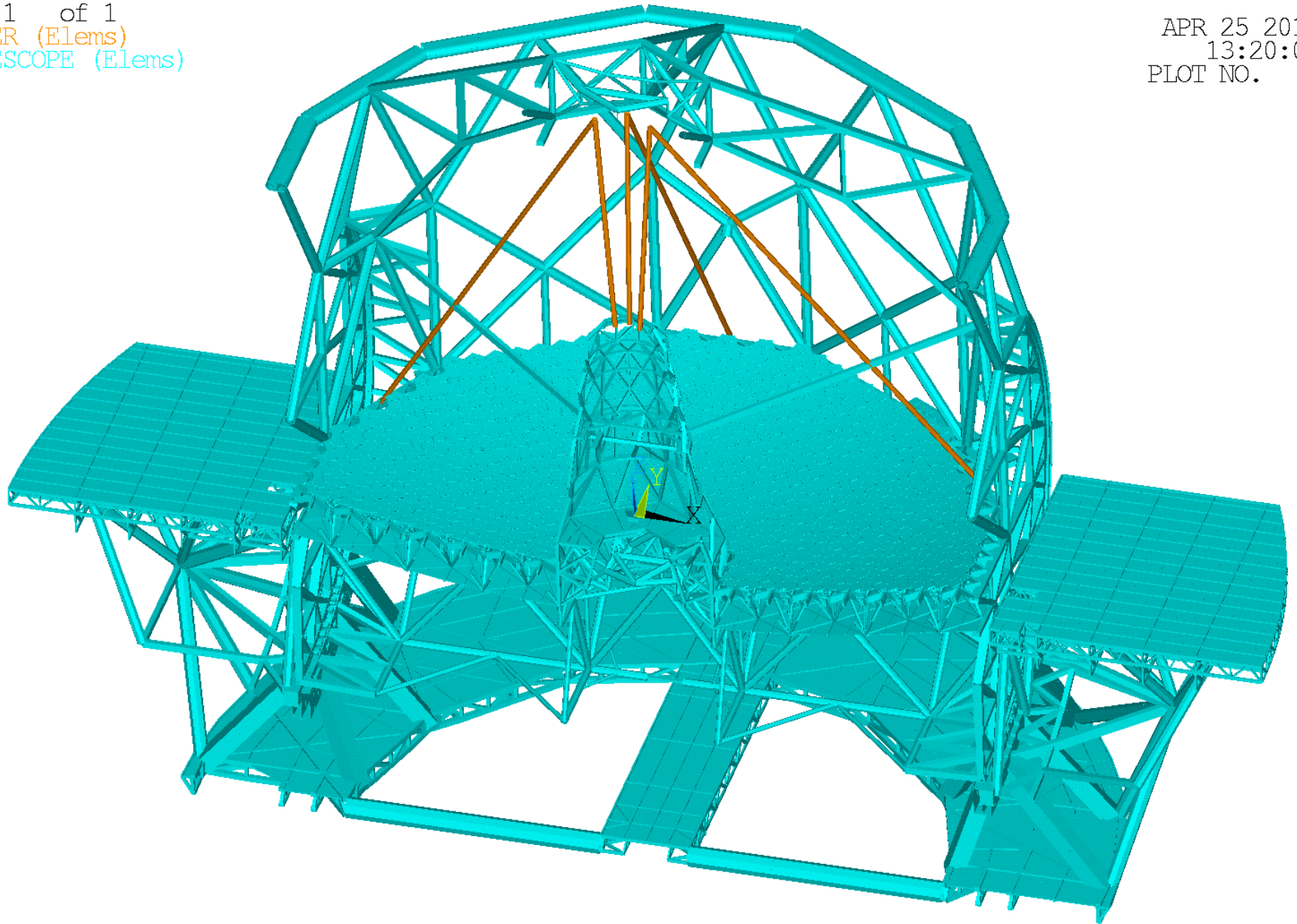
Goals & Concept

- Would be nice to...
 - Measure global M1/M2 position+figure *all the time*
 - Launch/receive beams in low-flex positions
 - Be transparent to science: Operate at 589nm or 1450nm?
 - Propagate (near-)parallel to science light: beat turbulence

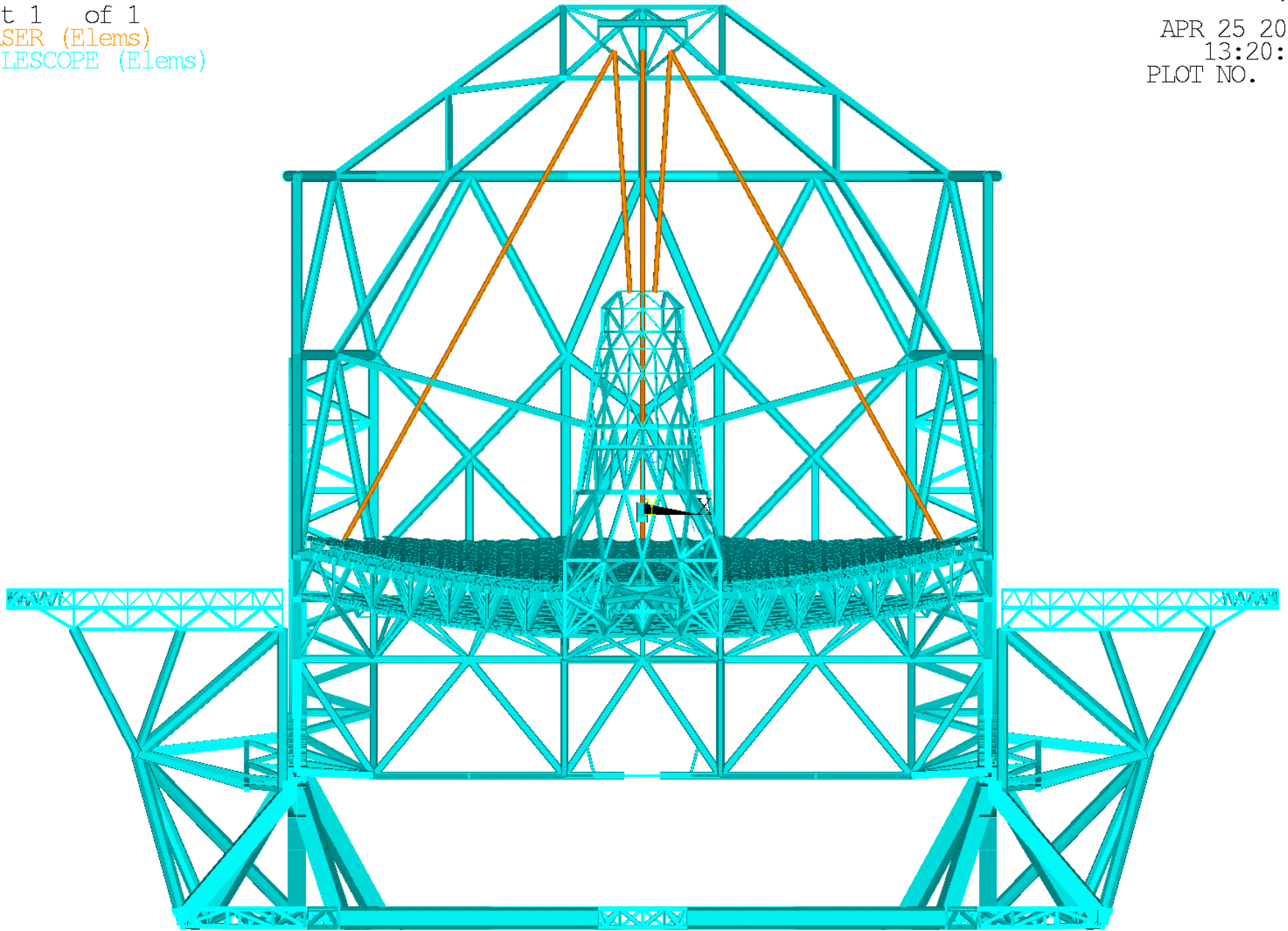
- Concept: Laser “Ray Tracer”
 - Launched from outer M1 segment edges...
 - ...reflected on M2...
 - ...and received near intermediate focus
 - Launched 13.5' off-axis (outward)
 - 70 mm beam, 100 μ W (eye safe)



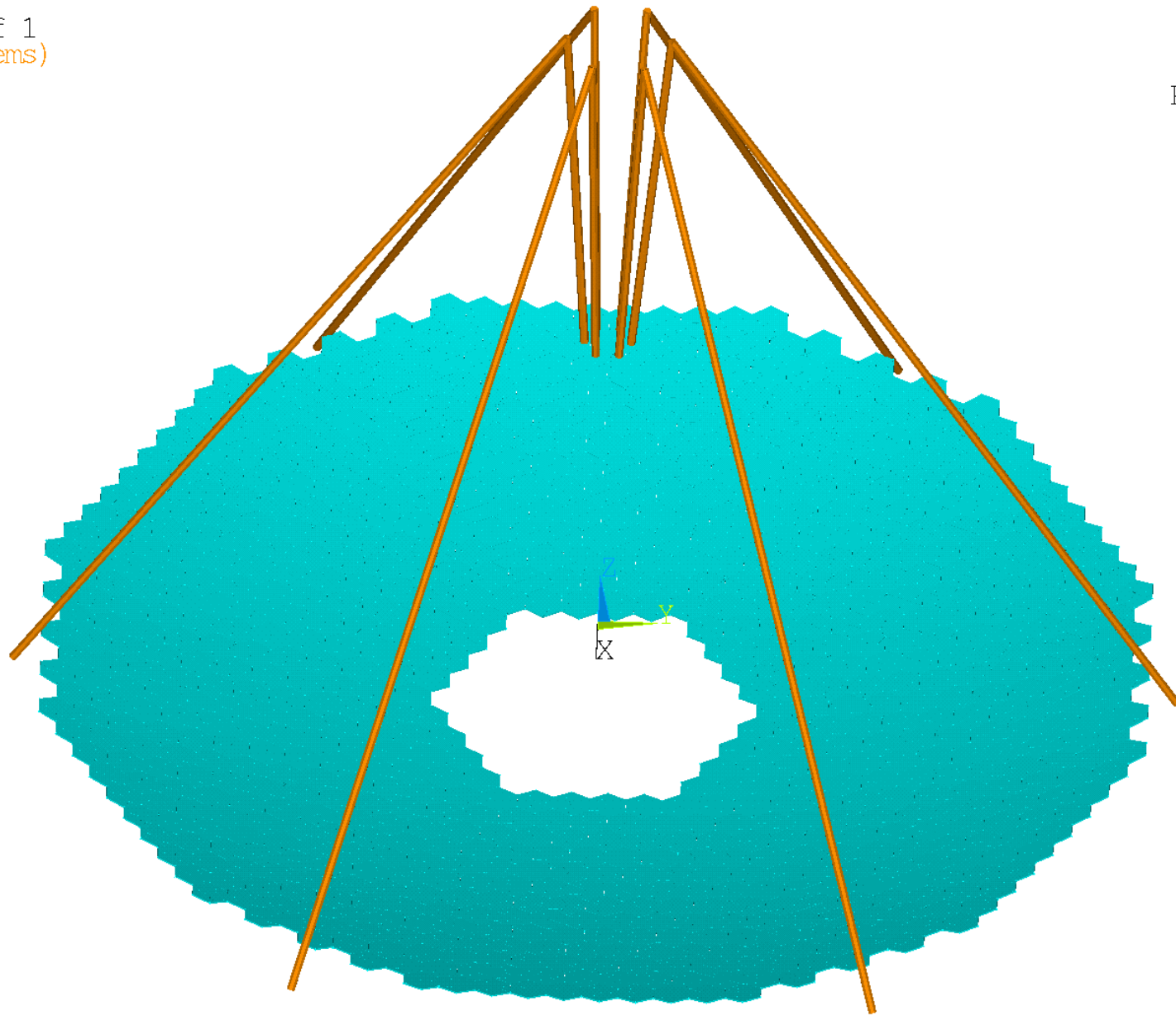
1 COMPONENTS
Set 1 of 1
LASER (Elems)
TELESCOPE (Elems)



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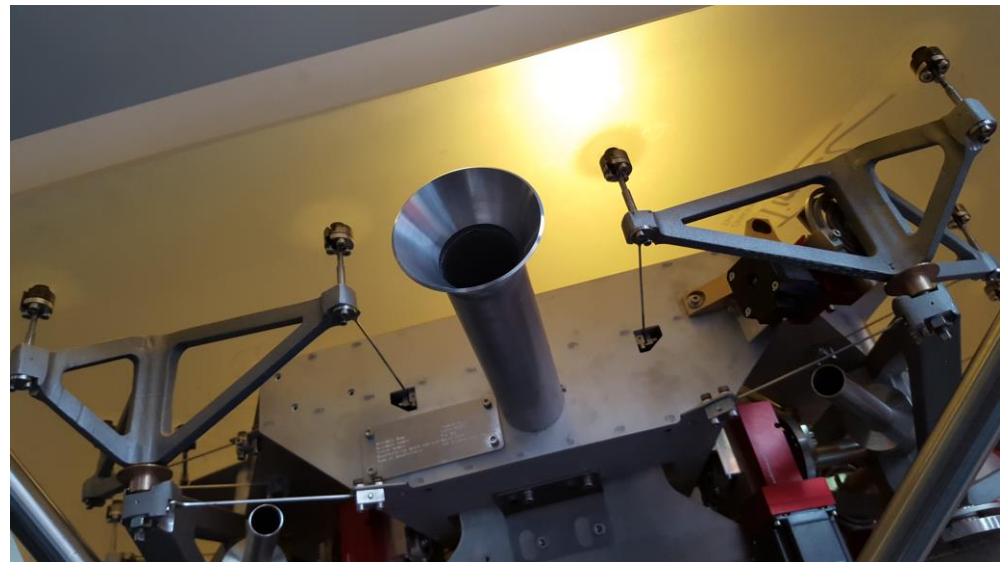
1 COMPONENTS
Set 1 of 1
LASER (Elems)
M1 (Elems)





Launch from Segment Support

- Projectors mounted on M1 segment support
 - Mount onto Moving Frame (holds mirror)
 - Requires very stiff arm (< 1" flexure)
 - Removable for segment exchange
 - Launch beams just outside of pupil and field
 - ...but close to both to emulate science light



Hardware

■ Lasers:

- Commercial narrow-band diode lasers
- Single-mode fiber coupled, not mounted on arm



■ Projectors:

- Diffraction-limited refractive 70mm beam expanders
- Monochromatic + on-axis (COTS item)
- One-time manual alignment



■ Position Sensitive Detector (PSD):

- COTS item, precision 10^{-5}
- Four-cathode photodiode, analog
- Frame rate 40 Hz and more
- Automatic 2-mirror beam acquisition assembly



Sense & Sensitivities

■ Gaussian beams

- Beam path lengths: $M1 \rightarrow 32.7 \text{ m} \rightarrow M2 \rightarrow 13 \text{ m} \rightarrow \text{det.}$
- Example: Weakly convergent launch, $2w = 39 \text{ mm}$
 - ➔ Footprints: 4.5 mm (M2), 3.0 mm (detector)
- $\text{RoC}_{\text{Launch}} = 40\text{m}$ (\sim focus on M2)
- Diameter far below r_0 , thus low beam distortion

■ Sense position (+angle) with Position Sensitive Det.

■ Raw mechanical sensitivities:

- 1" tip/tilt on M1: -0.6 mm , $-7''$ (M2 amplified)
- 1" tip/tilt on M2: $0.11 \times -0.12\text{mm}$, $1.8'' \times -1.9''$
- M1 RoC + 250 μm : -0.12 mm , $-1.4''$
- Piston M2 by 100 μm : -0.2 mm , $-2.1''$

Pros & Cons

■ Pros:

- Independent of turbulence outside M1–M2–detector
- Laser power can be tuned for sensor/frame rate
- Does not require accurate M1 phasing, independent of AO
- Can be used prior to M3–M6 commissioning
- Can be used anytime (daytime, during science)
- Fast enough to sense windshake
- Eye safe, non-disruptive
- Affordable (order 20k€ per laser/receiver)

■ Challenges / ToDo:

- Needs stiff launcher arm mounted to segment support
- Light pollution to science instruments to be studied



What Can Be Sensed?

- Assume 6 lasers equally spaced around M1
 - Yields 6 *xy*-positions + 6 *xy*-angles: 24 scalars total
 - M1 defocus, global tip/tilt and astig, M2: 5 rigid-body DoF
 - Can sense 11 scalar aberration modes in total
 - Use Generalized Least Squares method to assess errors:

Aberration	σ	Related EELT Allocation (preliminary)
M1 defocus (Z4)	1.2 nm RMS	5 μm RMS defocus (blind acq.); 3.5 μm RMS (plate scale)
M1 astigmatism (Z5,Z6)	1.7 nm RMS	
M1 trefoil (Z9,Z10)	1.0 nm RMS	
M2 tip/tilt vs. M1	0.031 μrad	25 μrad (blind acquisition)
M2 lateral displacement vs. M1	310 nm	100 μm (blind acquisition)
ARU Tower tip/tilt at top	0.15 μrad	

- The above values do not include projector jitter, turbulence, ARU Tower vibrations etc.

Conclusions

- Strength of the Ray Tracer lies in fast sensing of *relative* M1/M2 misalignment/misfigure/windshake
- Much smaller delay than laser interferometers like the *Etalon Multiline* (currently around 1s/target)
- Internal turbulence is sensed in a way relevant to the science PSF (real beam raytracing)
- Complements laser trackers which yield *absolute distances*, indispensable for AIV and coarse alignment after presets
- Cost estimate for six lasers/detectors: 110 k€

Seem like a good cost / benefit concept?